



Modeling comet coma photometry in ESA Comet Interceptor

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OUTLINE

I. Why and how to study comets?

II. Light scattering behind the model

III. Modeling and analysis

IV. Model validation

V. Conclusions



What are comets?

- Small Solar System objects composed of ice and dust
 - Coma and tails caused by Sun's radiation
 - Very eccentric, oval like, orbits
 - Formed in the outer Solar System

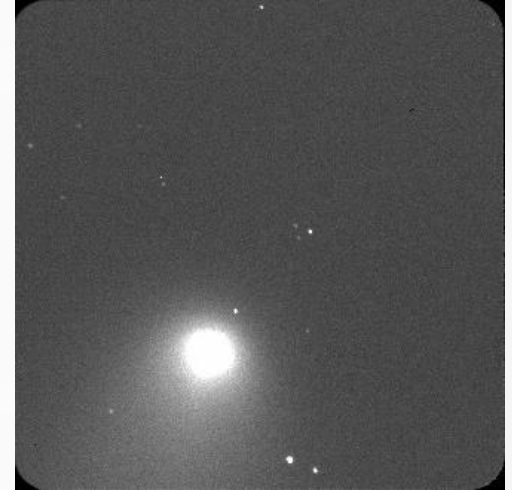


Fig 1: Image of comet C2023/A3 taken with Nordic Optical Telescope



Why study comets?

- Comets are relics from the formation of the Solar System
- Comets are less affected by space weathering than asteroids
- Challenges in studying comets
 - Most of the time very far away
 - Faint and dark objects



Fig 2: Comet C/2020 F3
(NEOWISE)



ESA Comet Interceptor



European space agency (ESA) mission aimed to explore a long-period comet entering the inner Solar System for the first time

- First ESA fast track mission
- Studies the comet's surface, shape, coma composition etc.
- Planned to launch 2029

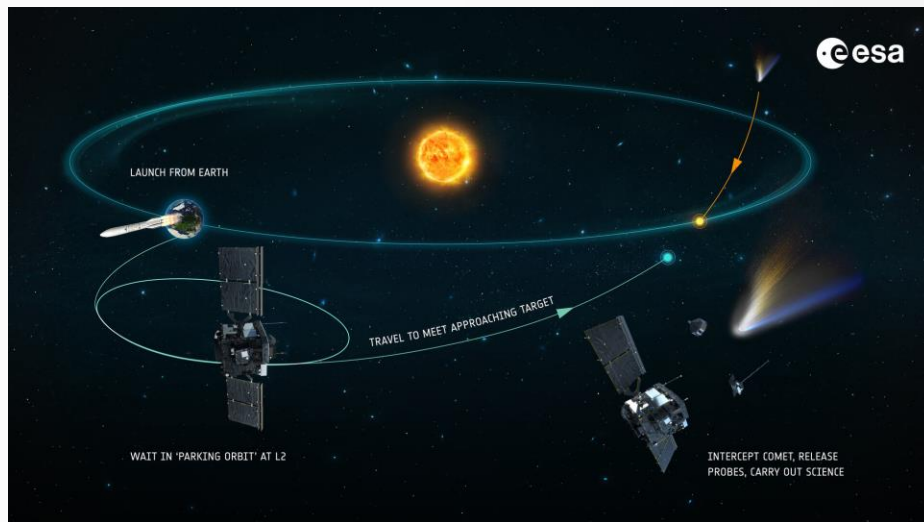


Fig 3: Comet Interceptor path to comet



Asteroid Image Simulator (AIS)

- Simulates imaging instrument operations for Solar System missions
 - Utilizes Blender combined with python-programming
 - Can simulate images from given flight paths of a spacecraft
 - Helps in missions planning

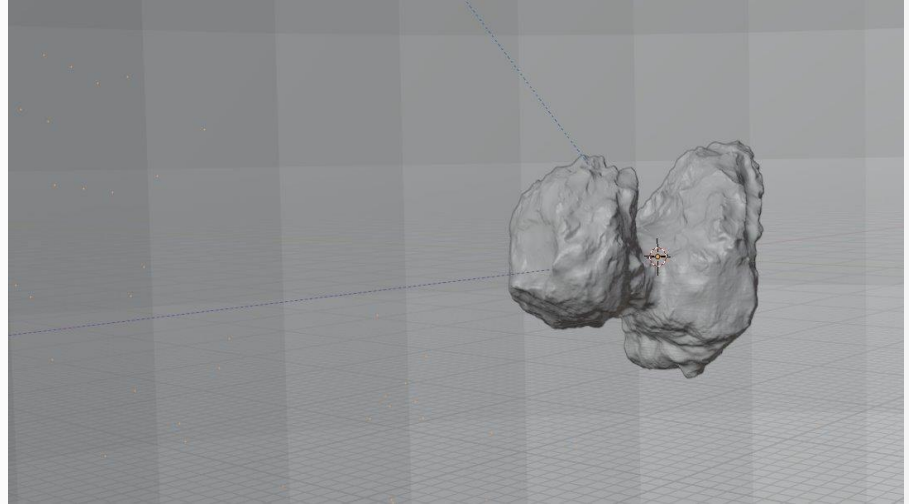


Fig 4: Comet 67/C-G in AIS



Scattering phase functions

- Phase function describes the angular distribution of scattered light
- Rayleigh scattering
 - Scattering by particles much smaller than the wavelength
 - Symmetrical phase function

$$P(\alpha) = \frac{3}{4} (1 + \cos^2 \alpha)$$

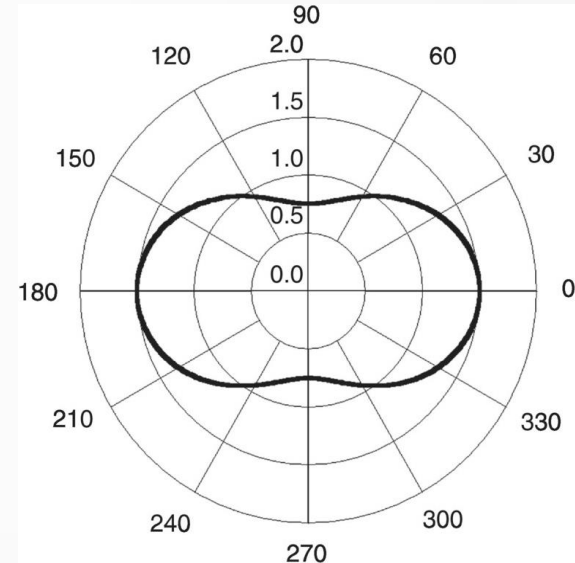


Fig 5: Rayleigh angular distribution



Scattering phase functions

- Henyey-Greenstein phase function
 - Only possibility in the model
 - Approximates light scattering from small particles, such as interstellar dust
 - Asymmetrical phase function

$$P(\theta) = \frac{1 - \xi^2}{(1 - 2\xi \cos \theta + \xi^2)^{3/2}}$$

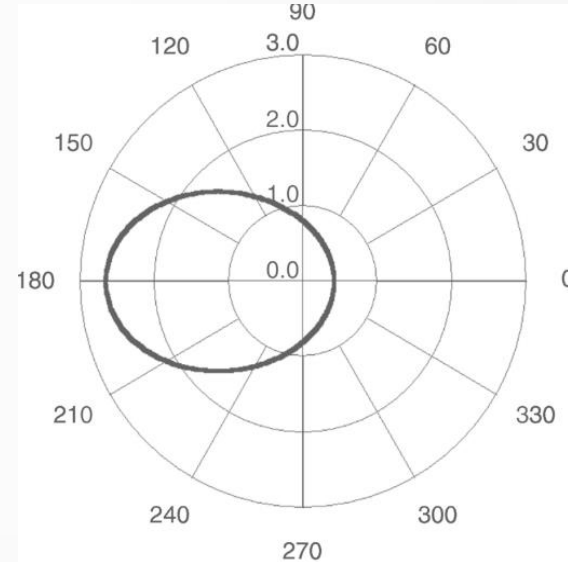


Fig 6: Henyey-Greenstein angular distribution



Comet gas/dust environment

- Crucial information for mission planning
- Engineering gas coma model (EGCM)
 - An analytical model of gas environment of comets
 - Tested if could be applied to dust -> can be
- Engineering dust coma model (EDCM)
 - A numeral spatial distribution model of a dusty-gas coma

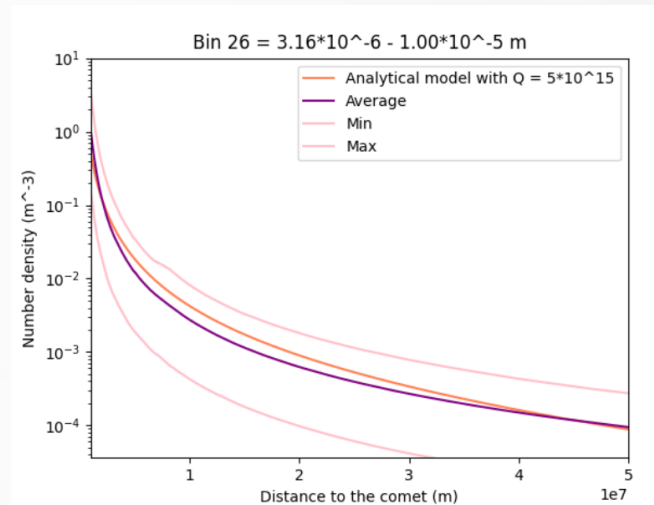


Fig 7: EDCM data compared with EGCM analytical expression



Modeling with Blender

- How is the simulation done?
 - Used shader volumetrics to simulate material inside a spherical volume
 - Blender has a 'volume scatter' node, but it can only model Henyey-Greenstein phase functions
 - > Mix 2 'volume scatter' nodes to get approximation of Rayleigh phase function

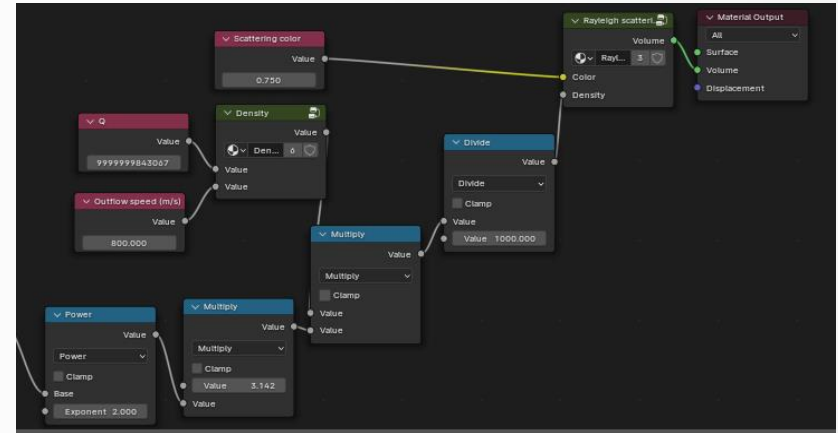


Fig 8: Example of a Blender shader node tree



Coma renders



Fig 9: With coma



Fig 10: Without coma



Photometry analysis

- Importing renders into a separate python program
 - Extract pixel color values from images
 - Calculate different metrics of brightness
 - Plot and export values into an output file



Fig 11: Animation of 67-CG simulated with AIS



Results

- Work in progress
- Coma brightness in different scenarios
- Pixel values as of now
 - Average brightness comparable to physical values

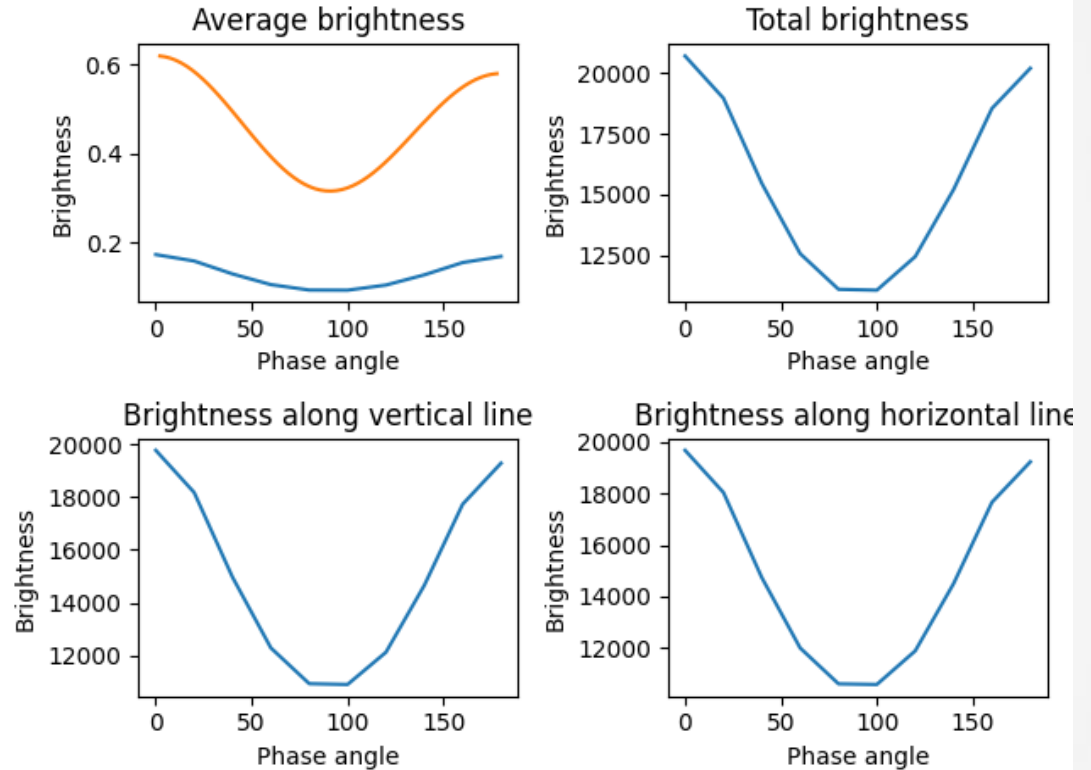


Fig 11: Preliminary results for brightness from model,
blue line = model results



The Radiative Transfer Coherent Backscattering (RT-CB)

- Electromagnetic scattering program for discrete sparse random media
- Numerically solves scattering properties of a given system as a function of angle
 - Much faster than exact computing methods

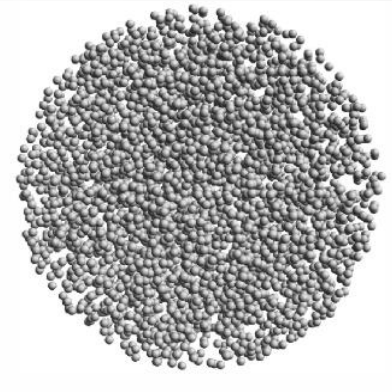


Fig 12: RT-CB random media sphere



Model validation

- RT-CB parameters:
 - Spherical volume with Rayleigh scatterers inside
 - Other parameters from coma model in Blender
- The model gives results that acceptably close to that of RT-CB
 - Difference likely caused by different phase functions

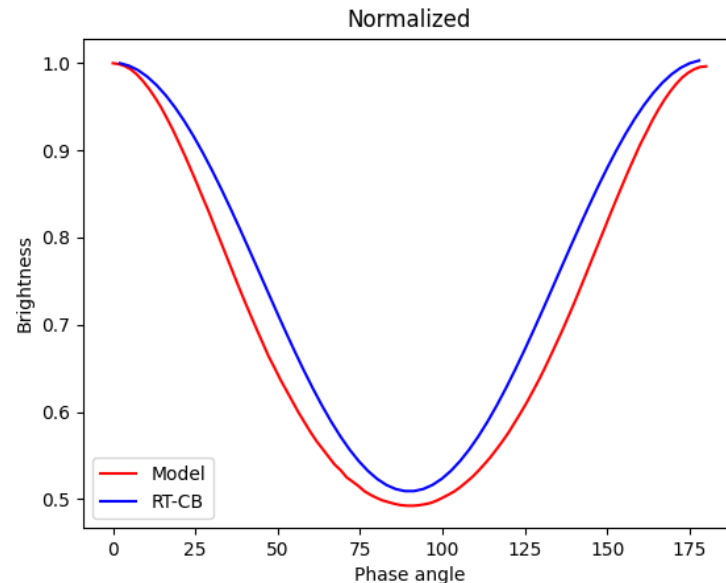


Fig 13: Normalized model photometry compared to RT-CB solution



What next?

- Without normalization a difference of π between model and RT-CB, why?
- Modeling asymmetrical phase functions
- Can the model be improved to give better results?



Conclusions

- Comets give us an insight into how the Solar System was formed
- Comet Interceptor aims to study a previously unknown comet
- AIS can be used to model what an imaging instrument onboard a spacecraft would see when flying by a comet
- The model agrees with RT-CB results with an acceptable accuracy



Questions?

Thank you for listening



References

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Image references

- Comet Interceptor logo: <https://www.cosmos.esa.int/web/comet-interceptor>
- Fig 3: https://www.esa.int/Science_Exploration/Space_Science/Comet_Interceptor
- Fig 2: <https://apod.nasa.gov/apod/ap200722.html>
- Fig 12: Vaisanen, Timo, et al. "Validation of Radiative Transfer and Coherent Backscattering for Discrete Random Media." *2016 URSI International Symposium on Electromagnetic Theory (EMTS)*, IEEE, 2016, pp. 396–99
- Fig 5&6: Shepard, Michael K. *Introduction to Planetary Photometry* / Michael K. Cambridge University Press, 2017.