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**GAUSSIAN REPRESENTATION OF ATOMIC
SCATTERING FACTORS FOR X-RAYS**

BY

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Abstract

Analytic approximations of theoretical atomic factors are discussed. For a number of atoms ($8 \leq Z \leq 37$) a Gaussian representation is given, fitted to the basic numerical values up to the last significant decimal place.

In detailed studies of electron distributions in crystals it has proved advantageous to employ analytic representations for the theoretical atomic factors, especially when electronic computers are used. Such a representation must, of course, coincide with the theoretical atomic factors accurately enough. When the representation is needed only in a limited z ($= \sin \theta/\lambda$) range, any reasonable functional form will do if it can be fitted to the tabulated values. However, when actually used in place of the theoretical values in a complete analysis, for instance with convolution methods [1], the function must also have the correct behaviour at infinity and near the origin.

In most cases the theoretical atomic factor $f(\mathbf{h})$ is spherically symmetric. Therefore only even functions $f(z)$ can be used for its description. Any odd components would make the atomic factor or some of its derivatives discontinuous at the origin causing errors at small z values. (This limitation is generally valid for the real part of the scattering amplitude of any real electron distribution in any direction.) On the other hand, we know that asymptotically $f(z) \sim e^{-az^2}$.

The Gaussian representation

$$f(z) = \sum a_i e^{-b_i z^2}, b_i > 0,$$

has all these properties. In addition, it has another remarkable advantage. The Fourier transform is easy to perform and gives directly an analytic representation for the corresponding density function.

First attempts to apply the Gaussian method evidently suffered from limited computing facilities. The fit with two or three Gaussians was not satisfactory for a detailed analysis. Therefore one turned to other approximations hoping to find simple forms with few parameters. During recent years many different propositions have been made [2–10]. Often they give good fits to the tabulated theoretical values, but do not fulfill the other conditions. Therefore they have limited applicability. One of the best fits

and most complete lists of atomic factors is given by CROMER and WABER [10], who represent their relativistic HFS scattering factors with a sum of four Gaussians and a constant. However, the constant term, or in some cases even a positive exponent spoils the behaviour at infinity and makes the corresponding density function singular. On the other hand, for the lighter elements ($Z \lesssim 36$) the older HF scattering factors with exchange are evidently more reliable [11]. In connection with our earlier works [12–16] we have been compelled to pay special attention to the possibilities of detailed analysis, and we have several times met with the need for more accurate analytic representations.

In most cases we can evidently make the fit as good as needed with a sufficient number of Gaussians. Also, in the use of present-day computers it is no appreciable disadvantage to have, say, 50 parameters instead of 5. Moreover, due to the general trend of the Gaussians, there is no danger of getting wrong values between standard points even with a great number of terms, as is the case with e.g. polynomial methods. Therefore we have found the pure Gaussian representation the most convenient. We have calculated the corresponding parameters a_i, b_i for the HF atomic factors of a number of common atoms ($8 \leq Z \leq 37$). The basic numerical values were taken from International Tables [17] except for F^- and O^{-2} for which the values of refs. [18] and [19, 20], respectively, were used. In some cases also the value at $z = 6 \text{ \AA}^{-1}$ from the tables of HANSON et al. [21] was taken into account. We needed 6 to 12 terms, depending on the case, to reach a fit up to the last significant decimal place of the theoretical values. The addition of the point $z = 6 \text{ \AA}^{-1}$ had only a slight tendency to increase this number.

There are some practical difficulties in finding what is the sufficient number of terms and in evaluating the parameters, especially because the parameters giving the desired fit need not be unique [22]. The calculations were performed on ELLIOTT 503 and 803 computers. The program was written in ALGOL. It was based on an iterative process sketched below.

As starting values one gives a set of b_i values. In the first stage the factors a_i are calculated for the given b_i values by the method of least squares, but with the limitation $a_i \geq 0$. This is done by using a special procedure written by PAATERO [23] and designed originally for the determination of lifetimes in radioactive decay. This procedure itself is iterative and its main feature is the use of gradients in a gradual evaluation of the a_i factors. Then, as a second stage, the number of a_i 's different from zero is decreased, if possible, combining adjacent b_i 's. The resulting b_i values form then the starting point of the next iteration cycle. The iteration is stopped when the sum of the squares of the residuals and the number of a_i 's do not change any more.

Another run of the calculation may give different parameters and even a different number of terms. The fit, however, is always the same. The accuracy of the resulting representations does not change regularly with Z , which was the case reported earlier for polynomial approximations [8]. The method accidentally failed to give any satisfactory representation for two ions (Cl^- , S^{-2}). This is certainly due to the fact that the Gaussians do not form a complete set of functions. Moreover, from a general point of view, we must admit that the good fits found for most atoms are exceptional.

The reliability of the representation was tested in some cases by new calculations leaving off a few standard points. Only a negligible effect on the fit was observed, unless the last points were missing, in which case the fit became less accurate at these points. If our results are compared with the values of HANSON et al. [21] for large values of κ , in the cases where these were not included in the calculation, we observe similarly that the fit is not so good at the ultimate values. In the worst cases the values at $\kappa = 6 \text{ \AA}^{-1}$ differ by a factor of 3. In general, however, it is sufficient to have the correct asymptotic behaviour, which is Gaussian. Numerical differences at large values of κ have no great significance. They could be corrected with some additional Gaussians with small a and large b , the effect of which is less important and well controlled [1].

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Table

Parameters for the Gaussian representations $f = \sum a_i \exp(-b_i \kappa^2)$ of some HF atomic factors. The number in brackets refers to the source of the basic numerical values if they are not taken from International Tables [17]. The asterisk indicates that the representation is fitted at $\kappa = 6 \text{ \AA}^{-1}$ to the values of Hanson et. al. [21].

Element	a_i	b_i	Element	a_i	b_i
O ⁻² [19]	0.20453	0.00050	Na ⁺	0.09678	0.00500
	1.55062	0.30000		1.54636	0.12860
	0.08506	2.00000		0.23579	0.75000
	3.05979	7.00000		3.07795	2.72163
	3.35961	17.00000		3.28779	5.36484
	0.97949	35.00000		1.40801	10.00000
	0.75679	50.00000		0.09262	15.00000
	0.00413	200.00000		0.25474	19.22734
O ⁻² [20]	0.20013	0.00100	Mg ⁺²	0.45642	0.00100
	0.13926	0.25000		1.42610	0.20000
	1.38808	0.30000		4.75005	2.51596
	0.12951	1.25000		3.33842	6.50000
	1.67653	5.50000		0.00643	7.00000
	1.37229	10.46827	0.02086	45.00000	
	2.07430	15.00000	Al	0.83846	0.00050
	0.27728	26.99502		0.98780	0.25000
	2.35692	34.59520		0.31253	1.00000
0.38143	105.00000	3.47904		1.92011	
F ⁻	0.76875	0.10000	3.69604	3.70307	
	0.96778	0.30000	0.28580	6.00000	
	0.09167	1.00000	0.68803	34.99805	
	2.13390	4.00000	0.16829	45.15373	
	2.27098	8.00000	1.27641	49.45225	
	1.43582	13.02203	0.91376	81.62434	
	1.12298	17.00000	0.33396	147.84744	
	1.15689	25.41793	0.01993	200.00000	
	0.05112	45.00000	Si	1.22577	0.00050
F ⁻ [18]	0.08143	0.10000		1.61095	0.75000
	1.67319	0.20718		0.93818	1.25000
	0.96358	3.18682		4.95028	2.50000
	2.48382	5.83983		0.71967	3.00000
	2.40192	11.00000		0.69155	25.00000
	1.31628	20.00000		2.58771	40.00000
	0.53276	29.26554		0.64003	50.00000
	0.54742	55.00000	0.63665	105.00000	

Table (continued)

Element	a_i	b_i	Element	a_i	b_i		
Si ⁺⁴	1.21831	0.00050	Mn	0.77362	0.00050		
	0.71084	0.50000		*	6.53870	0.30000	
	4.11611	1.50000			1.06898	0.50000	
	3.69942	3.50000			0.03077	3.00000	
	0.25449	7.00000			10.88315	5.27008	
	0.00213	40.00000			3.14224	14.19669	
K ⁺	1.37554	0.00050			0.13352	40.00000	
	6.81505	0.75000			2.09520	70.37231	
	0.65133	1.00000			0.02214	169.82470	
	2.21178	12.00000			0.31087	200.00000	
	5.81686	13.00000		Mn ⁺²	0.77363	0.00050	
	0.73405	30.00000			*	6.54018	0.30000
	0.39734	35.00000				1.07007	0.50000
Ca ⁺²	1.33169	0.00100				8.74776	5.00000
	3.05284	0.60000				2.01057	7.00000
	4.37381	0.70000			2.28022	11.00000	
	8.53855	10.52150		1.55365	19.00000		
	0.53324	27.00000		0.01763	80.00000		
	0.17106	31.00000		0.00637	300.00000		
Cr	0.72253	0.00050	Fe	0.81400	0.00050		
	* 5.08048	0.30000		*	0.42082	0.20000	
	2.69159	0.50000			7.08495	0.30000	
	2.16405	5.00000			8.91578	4.50000	
	8.45148	6.50000			2.32304	5.50000	
	2.01638	15.00000			3.15091	11.00000	
	1.47932	25.00000			0.82901	25.00000	
	1.32661	80.00000			0.63724	35.00000	
	0.06759	300.00000			1.09188	80.00000	
Cr ⁺²	0.72256	0.00050			0.71853	105.00000	
	* 5.22037	0.30000			0.01372	500.00000	
	2.50586	0.50000		Fe ⁺²	0.81423	0.00050	
	5.79508	5.50000			*	0.80580	0.25000
	4.29081	6.50000				6.69693	0.30000
	1.80522	13.00118				9.88854	4.50000
	1.06078	15.00000			0.53398	6.00000	
	0.59659	30.00000			4.12232	10.00000	
0.00199	500.00000		0.61148		19.00000		
			0.52019		25.00000		
			0.00662	500.00000			

Table (continued)

Element	a_i	b_i	Element	a_i	b_i
Co	0.85465	0.00050	Cu	0.09859	0.00050
*	4.29654	0.25000	*	0.79023	0.00100
	3.13831	0.30000		1.85939	0.10000
	0.23172	3.00000		5.64360	0.30000
	11.49973	4.23431		5.80408	3.00000
	4.17324	11.00000		8.29916	4.50000
	0.55434	30.00000		4.19229	11.00000
	0.78374	50.00000		0.94277	21.00000
	1.17554	80.00000		1.19802	56.51347
	0.29310	140.00000		0.94701	80.00000
				0.07739	222.08623
Co ⁺²	0.85439	0.00050	Cu ⁺	5.18697	0.10000
*	6.82574	0.25000		3.44572	0.50000
	0.78676	0.75000		10.93042	3.50000
	0.51046	3.00000		0.77955	5.00000
	6.40382	4.00000		3.86671	7.00000
	6.05618	5.50000		1.17649	11.00000
	3.11735	13.00000		2.34025	17.00000
	0.30823	19.00000		0.18298	45.00000
	0.13862	40.00000		0.09360	50.00000
Ni	0.88474	0.00050	Cu ⁺²	1.24509	0.00050
*	0.02954	0.15000		7.03399	0.25000
	7.33975	0.25000		5.99841	3.00000
	0.53856	2.50000		5.82549	4.00000
	11.14379	3.85286		4.41133	7.00000
	2.10810	7.00000		1.43773	11.00000
	2.99759	11.00000		1.04952	19.00000
	0.99639	26.45838	Zn	0.93770	0.00050
	1.71335	70.98541	*	0.14174	0.10000
	0.06691	105.00000		3.33422	0.15000
	0.18164	162.32611		3.90570	0.30000
Ni ⁺²	0.88488	0.00050		9.92243	3.00000
*	7.35430	0.25000		4.44998	4.50000
	0.05019	1.50000		4.42064	9.00000
	0.05583	1.75000		1.34828	30.00000
	10.20649	3.61442		0.09316	50.00000
	5.82664	7.50000		1.42132	80.00000
	1.37100	17.04543		0.02440	500.00000
	0.24744	23.00000			
	0.00281	500.00000			

Table (continued)

Element	a_i	b_i	Element	a_i	b_i	
Zn ⁺²	4.53316	0.05000	Br ⁻	0.94236	0.00050	
	4.27930	0.50000		*	6.01893	0.10000
	0.92567	2.50000			6.04228	1.25000
	12.50135	3.50000			12.92151	2.50000
	3.58286	8.00000			0.29038	5.00000
	1.44638	11.00000			0.83833	15.00000
	0.73116	21.00000			5.61768	23.00000
					2.12955	50.00000
Ge	0.49070	0.00500		1.17997	80.00000	
	7.19160	0.15000		0.01932	200.00000	
	2.13527	1.50000	Rb ⁺	2.60561	0.00100	
	8.10773	2.50000		4.81933	0.15000	
	7.79201	4.00000		10.95218	1.50000	
	1.48207	11.00000		7.24559	2.00000	
	0.76386	21.00000		7.09768	15.00000	
	1.44338	35.00000		1.89459	25.00000	
	1.88758	50.00000		1.38563	35.00000	
	0.68981	105.00000				
0.01772	140.00000					
Ge ⁺⁴	3.45118	0.00500				
	3.05246	0.25000				
	3.05962	0.75000				
	13.39202	3.00000				
	1.87172	4.00000				
	2.05435	8.00000				
	1.12008	9.00000				