

Notes for lectures by K. Kajantie, Jyväskylä 18-22.8.08

QCD & N=4 SYM

Known to everybody:

broken by $m_f = g_1^3$

flavor $SU(N_f)$

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu}^a F^{\mu\nu a} + \sum_f \bar{\Psi}_f i \gamma^\mu (\partial_\mu + i g A_\mu) \Psi_f \quad \uparrow \Psi_f \text{ is}$$

$$F_{\mu\nu}^a = \partial_\mu A_\nu^a - \partial_\nu A_\mu^a + g f_{abc} A_\mu^b A_\nu^c$$

$$[T_a, T_b] = i f_{abc} T_c \quad \left\{ \begin{array}{l} \text{exact local} \\ \text{color } SU(N) \end{array} \right.$$

- fundamental rep $\Psi^i \quad i = 1, \dots, N_c$ spinor
- adjoint rep $a = 1, \dots, N_c^2 - 1$ vector

$$S = \int d^d x \mathcal{L}$$

flat space, $dx^2 = \eta_{\mu\nu} dx^\mu dx^\nu \quad \eta_{\mu\nu} = (-+++)$

Classically: $S = \int d^d x (\partial A + g A^3)^2 \quad \alpha = \frac{e^2}{4\pi\epsilon_0 \hbar c} \Rightarrow \frac{e^2}{\hbar}$ is dim

dim: $\hbar = x^d (\frac{1}{x^2} A^2 + g^2 A^4 + \dots)$

$$\hbar = x^{d-2} A^2 \quad \hbar = g^2 x^d \frac{\hbar^2}{x^{2d-4}}$$

$$g^2 \hbar = x^{d-4} = \left(\frac{\hbar}{p}\right)^{d-4} = 1 \text{ for } d=4$$

Quantize theory, physics is in expectation values

$$\langle \phi(x_1) \dots \phi(x_n) \rangle$$

$$Z[J] = e^{iW[J]} = \int \mathcal{D}\phi(x) e^{iS[\phi(x)] + i \int d^d x J(x) \phi(x)}$$

↑ set of fields $x = (t, \vec{x})$

Thermal:

$$Z = e^{-\frac{F(T)}{T}} = e^{P \frac{V}{T}} = Z[J=0, \text{periodic}] = \int \mathcal{D}\phi(\tau, \vec{x}) e^{-\int_{\tau=0}^{\tau=1} d\tau \int d^d x \mathcal{L}[\phi(\tau, \vec{x})]}$$

$\phi(0, \vec{x}) = \phi(1, \vec{x})$

Some conventions:

Minkowski \leftrightarrow Euclidean:

$$\begin{aligned} \Gamma &= i^{\frac{1}{2}} & x_{\mu} x^{\mu} &= x_0^2 - \vec{x}^2 \rightarrow -x_0^2 + \vec{x}^2 \\ x_4 &= i x^0 & x^0 &= -i x_4 \\ \partial_0 &\rightarrow i \partial_4 & \partial_3 &\rightarrow i \partial_0 & A_0 &= i A_4 \end{aligned}$$

$$\begin{aligned} \int d^4x \delta(x^0 - t) &= \int d^3x \delta(x^0 - t) = \int d^3x \delta(x^0 - t) \frac{1}{i} \frac{\partial}{\partial x^0} (-i x^0) \\ &= \int d^3x \delta(x^0 - t) = \int d^3x \delta(x^0 - t) = \int d^3x \delta(x^0 - t) = \int d^3x \delta(x^0 - t) \end{aligned}$$

Mostly - \leftrightarrow Mostly +

$$+ \dots \leftrightarrow - \dots \quad \eta_{\mu\nu} \leftrightarrow -\eta_{\mu\nu} \quad g_{\mu\nu} \leftrightarrow -g_{\mu\nu}$$

$$\mathcal{L}[\varphi] = \frac{1}{2} \eta^{\mu\nu} \partial_{\mu} \varphi \partial_{\nu} \varphi - V(\varphi) \rightarrow -\frac{1}{2} \partial_{\mu} \varphi \partial^{\mu} \varphi - V(\varphi)$$

$$\mathcal{L}[A^{\mu}] = -\frac{1}{4} F_{\mu\nu}^2 = -\frac{1}{4} g^{\mu\nu} \partial_{\mu} A^{\rho} \partial_{\nu} A_{\rho} \rightarrow -\frac{1}{4} F_{\mu\nu}^2$$

$$x^{\mu} x_{\mu} = x_0^2 - \vec{x}^2 \rightarrow -x_0^2 + \vec{x}^2$$

$$\Gamma = g^{\mu\nu} \partial_{\mu} \partial_{\nu} \rightarrow \Gamma$$

$$R^{\mu\nu\rho\sigma} = 2\Gamma^{\rho} - \Gamma^{\sigma} \rightarrow R^{\mu\nu\rho\sigma}$$

$$R_{\mu\nu\rho\sigma} = \rightarrow -R_{\mu\nu\rho\sigma}$$

$$R_{\mu\nu} = g^{\rho\sigma} R_{\mu\rho\nu\sigma} \rightarrow R_{\mu\nu}$$

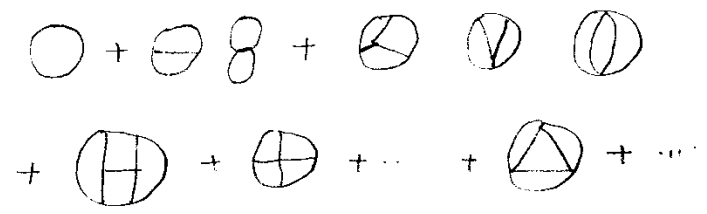
$$R = g^{\mu\nu} R_{\mu\nu} \rightarrow -R$$

$$R_{\mu\nu} = \frac{1}{2} R g_{\mu\nu} + \Lambda g_{\mu\nu} \rightarrow R_{\mu\nu} = \frac{1}{2} R g_{\mu\nu} - \Lambda g_{\mu\nu}$$

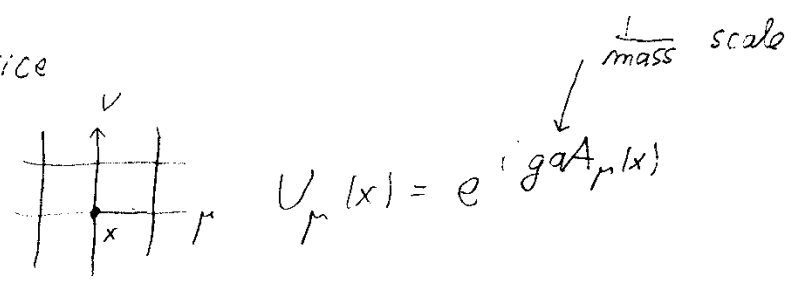
Regularize maintaining local GI:

- \overline{MS} $d = 4 - 2\epsilon$ expand in g

mass scale $\rightarrow \mu^{2\epsilon} \int d^{4-2\epsilon} x$



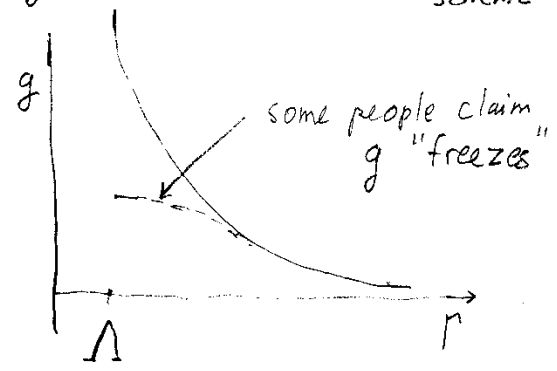
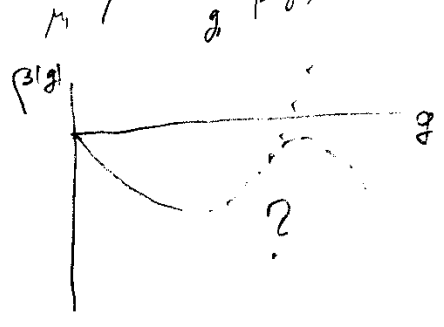
- Lattice



$\Rightarrow \mu \frac{\partial g}{\partial \mu} \equiv \beta(g) = -\beta_0 g^3 - \beta_1 g^5 - \beta_2 g^7 - \beta_3 g^9 - \dots$

Scheme independent, same for \overline{MS} , \overline{MS} , latt,

$\int_{\mu}^{\mu_2} \frac{d\mu}{\mu} = \int_{g}^{g_2} \frac{dg}{\beta(g)} \Rightarrow \text{integration constant } \Lambda_{\text{scheme}}$



running of g breaks conformal invariance

Chiral symmetry is trivially there in \mathcal{L}_{QCD}

$\Psi_L \rightarrow U_L \Psi_L \quad \Psi_R \rightarrow U_R \Psi_R$

spontaneously broken to $SU_V(4) \Rightarrow$ essential

$\mathcal{L}_{N=4 \text{ SYM}} = ?$

A canonical way of deriving this (Diverchia, -th/9803026)

$S = \int d^{10}x \left\{ -\frac{1}{4} F_{AB}^a F^{aAB} - \frac{1}{2} \bar{\lambda}^a i \Gamma_A (D^A \lambda)^a \right\}$
 (Annotations: $1, \dots, N_c^2-1$ for F_{AB}^a ; $0, 1, \dots, 9$ for A ; $10d$ Dirac for Γ_A ; $\gamma^{AB} (\frac{1}{2} \lambda^a - g f_{abc} A_A^{b,c})$ for the second term)

Susy generators $Q_a^i, \bar{Q}_a^i \quad i=1, \dots, N$

$\downarrow \begin{pmatrix} 1 \\ \frac{1}{2} \end{pmatrix}$ $N=1$ supersymm. theory in 10d

$8 = d - 2$ bosonic dots = $\frac{1}{4} 2^{\frac{d}{2}} = 8$ fermionic dots $|\neq d_A$

Invar. under $\begin{cases} \delta A_A^a = \frac{i}{2} (\bar{\lambda}^a \Gamma_A \xi - \bar{\xi} \Gamma_A \lambda^a) \\ \delta \lambda_a = \sigma_{AB} F_a^{AB} \xi \quad \delta \bar{\lambda}_a = \dots \end{cases}$

spinor reps. of $O(d)$:
 $(x_1^2 + \dots + x_d^2) = (\gamma_1 x_1 + \dots + \gamma_d x_d)$
 $\{\gamma_i, \gamma_j\} = 2\delta_{ij}$
 Find a $2^{\frac{d}{2}}$ -dim rep for this = spinor

Do dimensional reduction:

$X^A = (x^0, \bar{x}, x^4, \dots, x^9)$
 x^r fields do not depend on these x^i

$\Rightarrow F_{ij}^a = g f_{abc} A_i^b A_j^c \rightarrow$ get 6 adjoint scalars ϕ_i

$F_{\mu i} = \partial_\mu A_i - \cancel{\partial_i A_\mu} - g A_\mu A_i \sim D_\mu A_i \uparrow$

$\Rightarrow S[A_\mu^a, \phi_i^a, \psi^a, \bar{\psi}^a] = \int d^4x \left[-\frac{1}{4} F_{\mu\nu}^2 + \frac{1}{2} (D_\mu \phi_i)^2 - \frac{1}{2} \bar{\psi}^a i \not{D} \psi^a - \bar{\psi} \phi \psi - \phi^4 \right]$

$\begin{matrix} \downarrow 1 \\ \downarrow \frac{1}{2} \\ \downarrow 0 \\ \downarrow -\frac{1}{2} \\ \downarrow -1 \end{matrix}$
 $\begin{matrix} 1 \\ 4 \\ 6 \\ 4 \\ 1 \end{matrix}$
 $N=4$ susy
 This theory is said to have identically $\beta(g) = 0 \Rightarrow \frac{\partial g}{\partial t} = 0 \Rightarrow g = \text{number}$
 $\frac{8b+8f \text{ dots}}{g \ll 1}$ weakly coupled, pert theory
 $\left. \begin{matrix} g \gg 1 \\ (g^2 N_c \gg 1) \end{matrix} \right\}$ strongly " , AdS/CFT

No "hadrons"! No energy gap!

Conformal invariance

Invariance group of Maxwell

- Lorentz 1892: Lorentz $O(1,3)$ + Translations

$x^\mu = \Lambda^\mu_\nu x^\nu$ preserves $-t^2 + \vec{x}^2$

$$e^{\frac{i}{2} \omega_{\mu\nu} M^{\mu\nu}} \quad 6 \quad + \quad 4 \quad = 10 \text{ param.}$$

(Poincaré group)

- Cunningham-Bateman 1902: also

- Dilatations $D: x^\mu \rightarrow \lambda x^\mu$ (1)

- Special conf. transformations K_μ : "translation of inversion, $\frac{1}{x^\mu} \sim \frac{1}{x} + b$ "

(4)
$$x^\mu \rightarrow \frac{x^\mu + b^\mu x^2}{1 + 2x \cdot b + b^2 x^2} = x'^\mu$$

with indices
 \Rightarrow take a square
 \downarrow

$$\frac{x'^\mu}{x'^2} = \frac{x^\mu}{x^2} + b^\mu \quad \frac{1}{x'^2} = \frac{1}{x^2} + 2 \frac{b \cdot x}{x^2} + b^2$$

Total $10 + 5 = 15 = \frac{d(d-1)}{2}$ for $d=6 \Rightarrow$ the group is $O(2,4)$

Conformal algebra: (cf. ang. mom $[L_i, L_j] = i \epsilon_{ijk} L_k$)

mass dimension		P_α	$M_{\alpha\beta}$	D	K_α
1	P_μ	0	$i(\eta_{\alpha\mu} P_\beta - \eta_{\beta\mu} P_\alpha)$	$-i P_\mu$	$2i(\eta_{\mu\alpha} D + M_{\mu\alpha})$
0	$M_{\mu\nu}$		$i(\eta_{\mu\beta} K_\nu - \eta_{\nu\beta} K_\mu)$	0	$i(\eta_{\mu\alpha} K_\nu - \eta_{\nu\alpha} K_\mu)$
0	D		antisymm	0	$-i K_\alpha$
-1	K_μ				0

Generalisations:

Poincare \rightarrow SUSY algebra

Have to know reps of $N(4) \sim SL(2, C)$
 $M_{\psi_a} \quad M^* \bar{\psi}_a \quad M, M^*$ not equivalent
 Brann et al 0806.2531

Mass dim:			P_α	$M_{\alpha\beta}$	Q_b	\bar{Q}_b
1	4	P_μ	0	$i\gamma P$	0	0
0		$M_{\mu\nu}$	$-i\gamma P$	$i\gamma M$	$\sigma_{\mu\nu} Q$	$\bar{Q} \bar{\sigma}_{\mu\nu}$
$\frac{1}{2}$	2	Q_a	antis		0	$2\sigma_{ab}^M P_\mu$
$\frac{1}{2}$	2	\bar{Q}_a			symm	

key relation!

can have diff # of bos & ferm generators: have to count states

Conformal \rightarrow Superc conformal algebra

superanalogue for $D, K_\mu = ?$

mass dim

Q_a^i	\bar{Q}_a^i	\bar{S}_a^i	S_a^i	$i = 1, \dots, 4 = N$
8	8	8	8	
	$+\frac{1}{2}$		$-\frac{1}{2}$	

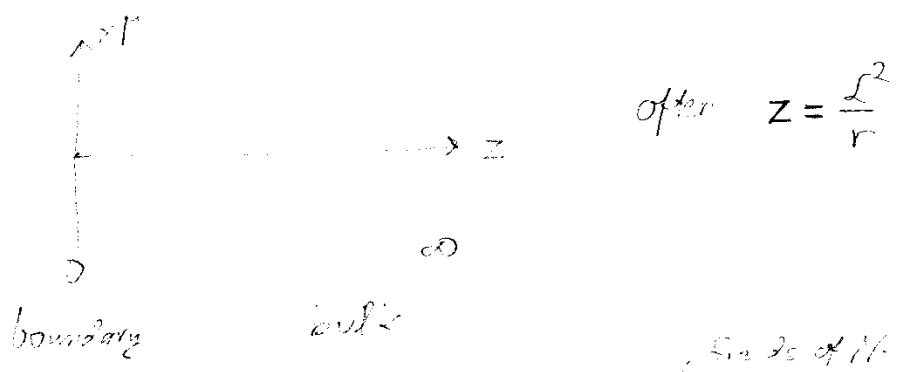
$[Q, K] \sim \sigma S$

see d'Hoker, sect. 3.3
 Lyng Petersen, " 3.6

$\{Q_a^i, \bar{Q}_a^j\} = \delta_{ij} \sigma_{aa}^\mu P_\mu \Rightarrow U(N)$ invariance
 $SU(4) \sim SO(6)$
 $4^2 - 1 = 15 \quad \frac{6 \cdot 5}{2} = 15$

2. Master Formula

4d QFT \Leftrightarrow 5d classical gravity in AdS
 x x, z radius L



$$Z[\mathcal{O}] = e^{iW[\mathcal{O}]} = \int \mathcal{D}\psi e^{iS[\psi(x)] + i \int d^4x J(x) \psi(x)}$$

ends of AdS spacetime or related

"single trace op" in SYM₄

For any operator $\mathcal{O}(x)$ (combination of ψ 's, like $F_{\mu\nu}^2(x)$)

a generating functional of \mathcal{O} 's: $\langle \mathcal{O}(x_n) \rangle$ is

$$\frac{1}{Z[0]} \int \mathcal{D}\psi e^{iS[\psi(x)] + i \int d^4x \varphi_0(x) \mathcal{O}(x)} \quad (\Rightarrow \varphi_0: \varphi, \mathcal{O})$$

have to fix dims by this!

$$\langle e^{i \int d^4x \varphi_0 \mathcal{O}} \rangle = e^{i S_{\text{grav}}[\varphi(x, z), \varphi(x, z=0) = z^{4-\Delta} \varphi_0(x)]}$$

mostly written in Euclidean form!

$0 = \Delta - 4 + 4 - \Delta$

in the $N_c \gg 1$ $g^2 N_c \gg 1$ limit

Simplest for the thermal partition fn $\langle \mathcal{O} \rangle = \frac{\int \mathcal{D}\psi e^{iS[\psi]}}{\int \mathcal{D}\psi e^{iS[\psi]}}$

$$Z(\tau) = \langle 1 \rangle = e^{-S_{\text{grav}}^E}$$

before normalisation

Need class gravity

3. Classical gravity (or weak field quantum)

$m=0 \quad s=1 \quad \overset{\text{SU(2) matrix}}{\downarrow} \quad A_\mu(x) \quad \int d^4x \left(\frac{1}{4} F_{\mu\nu}^2 \right) \quad \text{gauge inv.}$
 $=0 \quad s=2 \quad g_{\mu\nu}(x) \quad \int d^4x \sqrt{-g} R \quad \text{inv under } x^\mu \rightarrow x'^\mu(x)$

Given:

coordinates x : $(t \ x \ y \ z) \quad (t \ r \ \theta \ \varphi)$

$(t \ x \ y \ r \ z) \quad (r \ \eta \ x^2 \ x^3 \ z)$

metric

$$g_{\mu\nu} = \begin{pmatrix} & t & r & \theta & \varphi \\ & -1 & & & \\ & \frac{a(t)}{1-kr^2} & & & \\ & 0 & a(t)r^2 & & \\ & & & a(t)r^2 \sin^2\theta & \end{pmatrix} \quad \text{FRW}$$

often diagonal

$ds^2 = g_{\mu\nu}(x) dx^\mu dx^\nu \quad g^{\mu\alpha} g_{\alpha\nu} = \delta^\mu_\nu \quad g = \det g_{\mu\nu}$

$V^\mu = \frac{\partial x'^\mu}{\partial x^\alpha} V^\alpha \quad V^\mu_\nu = \frac{\partial x^\alpha}{\partial x'^\nu} V^\alpha \quad g^{\mu\nu} = \dots \quad g_{\alpha\beta}$

Covariant derivative $D_\mu = \partial_\mu + ig A_\mu$

$\nabla_\nu V^\mu = \partial_\nu V^\mu + \underbrace{\Gamma^\mu_{\nu\alpha}}_{\text{Christoffel}} V^\alpha \rightarrow \frac{\partial x'^\mu}{\partial x^\alpha} \frac{\partial x^\beta}{\partial x'^\nu} \nabla_\beta V^\alpha$
 $\Gamma^\mu_{\nu\lambda} = \frac{1}{2} g^{\mu\alpha} [\partial_\nu g_{\lambda\alpha} + \partial_\lambda g_{\nu\alpha} - \partial_\alpha g_{\nu\lambda}]$

$\nabla_\nu V_\mu = \partial_\nu V_\mu - \Gamma^\alpha_{\nu\mu} V_\alpha$

To find Γ I would find the eq. for a geodesic line by extremising $L(x^\alpha, \dot{x}^\alpha) = g_{\mu\nu}(x^\alpha) \dot{x}^\mu \dot{x}^\nu \quad \dot{x}^\mu \equiv \frac{dx^\mu}{ds}$

$\Rightarrow \frac{\partial L}{\partial x^\alpha} - \frac{d}{ds} \frac{\partial L}{\partial \dot{x}^\alpha} = 0 \quad \Rightarrow \ddot{x}^\sigma + \underbrace{\Gamma^\sigma_{\mu\nu}}_{\text{read } \Gamma \text{ from the result}} \dot{x}^\mu \dot{x}^\nu = 0$

$\Gamma \sim g (\partial g + \partial g - \partial g)$

Riemann $F_{\mu\nu} = \frac{i}{g} [D_\mu, D_\nu]$

$$[\nabla_\mu, \nabla_\nu] V^\alpha = R^\alpha_{\beta\mu\nu} V^\beta$$

$$R^\alpha_{\beta\mu\nu} = \underbrace{\partial_\mu \Gamma^\alpha_{\nu\beta} - \partial_\nu \Gamma^\alpha_{\mu\beta}}_{\text{contains } \partial^2 g} + \underbrace{\Gamma^\alpha_{\mu\gamma} \Gamma^\gamma_{\nu\beta} - \Gamma^\alpha_{\nu\gamma} \Gamma^\gamma_{\mu\beta}}_{\text{here } (\partial g)^2}$$

Ricci $R_{\mu\nu} = R^\alpha_{\mu\alpha\nu} \equiv g^{\alpha\beta} R_{\alpha\mu\beta\nu}$ $R = g^{\mu\nu} R_{\mu\nu}$
 conventions! $R_{\mu\nu\alpha\beta} = 0$ $R_{\mu\nu\alpha\beta} = 0$ curvature scalar

$R_{\mu\nu\alpha\beta}$ has d $\frac{1}{12} d^2 (d^2 - 1)$ indep. comps

2	1
3	6
4	20
5	50

one can construct d $\frac{1}{12} d(d-1)(d-2)(d+3)$ scalars from $R_{\mu\nu\alpha\beta}$ $g_{\mu\nu}$

2	0 → 1
3	3
4	14
5	40 — Kretschmann scalar

$R^{\mu\nu\alpha\beta} R_{\mu\nu\alpha\beta}$

$$\nabla_\mu V^\nu = \partial_\mu V^\nu + \Gamma^\nu_{\lambda\mu} V^\lambda \quad V^\lambda \neq \partial^\lambda \phi$$

$$\nabla_\mu \Gamma^\nu_{\lambda\sigma} = \partial_\mu \Gamma^\nu_{\lambda\sigma} - \Gamma^\nu_{\rho\mu} \Gamma^\rho_{\lambda\sigma} + \Gamma^\rho_{\lambda\mu} \Gamma^\nu_{\rho\sigma} - \Gamma^\rho_{\lambda\sigma} \Gamma^\nu_{\rho\mu}$$

$$\nabla_\mu \Gamma^\nu_{\lambda\sigma} = \frac{1}{2} g^{\nu\rho} (\partial_\mu g_{\rho\lambda} + \partial_\mu g_{\rho\sigma} - \partial_\rho g_{\lambda\sigma}) + \Gamma^\nu_{\rho\mu} \Gamma^\rho_{\lambda\sigma} - \Gamma^\rho_{\lambda\mu} \Gamma^\nu_{\rho\sigma} - \Gamma^\rho_{\lambda\sigma} \Gamma^\nu_{\rho\mu}$$

$$= 0 \text{ if } \Gamma^\nu_{\lambda\sigma} = -\Gamma^{\nu\lambda}_{\sigma}$$

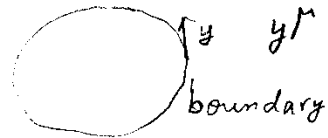
Einstein-Hilbert with Gibbons-Hawking boundary term + matter
lots of them in AdS!

$$S = \frac{1}{16\pi G_{d+1}} \left\{ \int_M d^{d+1}x \sqrt{-g} (R + 2\Lambda) - \int_{\partial M} d^d y \sqrt{-\gamma} 2K \right\} + S_m$$

$$l = \frac{1}{G_{d+1}} x^{d+1} \cdot \frac{1}{x^2}$$

$$\Rightarrow \frac{l^{d-1}}{G_{d+1}} = \text{dimless}$$

$$\text{will meet } \frac{l^3}{4\pi G_5} = \frac{N_c^2}{2\pi^2}$$



induced metric

$$\gamma_{\mu\nu} = g_{MN} \frac{\partial x^M}{\partial y^\mu} \frac{\partial x^N}{\partial y^\nu}$$

extrinsic curvature

$$K = g^{MN} K_{MN}$$

$$\text{EOM from } \frac{\delta S}{\delta g^{\mu\nu}} = 0: \quad g^{\mu\nu} \delta g_{\mu\nu} = -g_{\mu\nu} \delta g^{\mu\nu}$$

$$\left[R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} - \Lambda g_{\mu\nu} = 8\pi G T_{\mu\nu} \quad T_{\mu\nu} = \frac{-g}{\sqrt{-g}} \frac{\delta S_m}{\delta g^{\mu\nu}} \right]$$

$$\nabla^\mu T_{\mu\nu} = 0 \quad \nabla^\mu g_{\mu\nu} = 0 \quad \text{since } \nabla^\mu g_{\mu\nu} = \nabla^\mu \delta_{\mu\nu} = 0$$

and this is a tensor

$$\text{AdS}_{d+1}: \quad \Lambda = \frac{d(d-1)}{2L^2} = \frac{6}{L^2} \quad \text{for AdS}_5, \quad d=4$$

$$R = -\frac{d(d+1)}{L^2} \quad R_{\mu\nu} = \frac{-d}{L^2} g_{\mu\nu} \quad R + 2\Lambda = \frac{-2d}{L^2}$$

Sign of Λ is a convention: here $\Lambda_{\text{AdS}} > 0$

$$\text{Sign of } T_{00} \text{ matters, } S_m = 0 \Rightarrow 8\pi G T_{00}^{\text{AdS}} = \Lambda g_{00} = -\Lambda < 0$$

$$T_{00}^{\text{AdS}} > 0$$

(-++++)

$$\int d^5x \sqrt{-g} e^{-2\Phi} [R + 4g^{MN} \partial_M \phi \partial_N \phi + \dots] = \int d^5x \sqrt{-g} [R - \frac{1}{2} g^{MN} \partial_M \phi \partial_N \phi + \dots] \quad -10-$$

"String frame" $g_{MN} \rightarrow e^{2\phi/3} g_{MN}$ "Einstein frame"

More fields: + for $\eta = (+ - - - -)$

$$S_{EH} + \int d^d x \sqrt{-g} \left\{ -\frac{1}{2} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi - V(\phi) - \frac{1}{4} g^{\mu\alpha} g^{\nu\beta} F_{\mu\nu} F_{\alpha\beta} + \dots \right\}$$

$$\Rightarrow R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} = 8\pi G \left[\partial_\mu \phi \partial_\nu \phi + g_{\mu\nu} L[\phi] + F_{\mu\alpha} F_{\nu}^{\alpha} - \frac{1}{2} g_{\mu\nu} F^2 \right]$$

$$\frac{\partial L}{\partial \phi} - \partial_r \frac{\partial L}{\partial \dot{\phi}} = 0 \Rightarrow \square \phi - V'(\phi) = 0$$

sign $(-\partial_0^2 + \vec{\partial}^2) \phi - m^2 \phi$
 $V = \frac{1}{2} m^2 \phi^2 \quad V' = m^2 \phi = (\omega^2 - k^2 - m^2) \phi$

$$\frac{\partial L}{\partial A^\alpha} - \partial_r \frac{\partial L}{\partial \dot{A}^\alpha} = 0 \Rightarrow g^{\mu\alpha} \nabla_\mu F_{\nu\alpha} = J_\nu = 0 \quad \text{Maxwell}$$

if ϕ charged, e.g.
 $\nabla_\mu F_{\nu\alpha} = \partial_\mu A_\nu - \partial_\nu A_\mu$
 String & 3d gravity \rightarrow 4d BH

solves $R_{\mu\nu} = 0$ for a spherically symm ansatz static! Birkhoff

$$ds^2 = -f(r) dt^2 + \frac{dr^2}{g(r)} + r^2 (d\theta^2 + \sin^2\theta d\phi^2)$$

$$f(r) = g(r) = 1 - \frac{r_s}{r}$$

$$r_s = 2G_4 M \quad \text{NOT in EDM}$$

$$R_{\mu\nu} = 0$$

$$T_H = \frac{\sqrt{f'(r_s) g'(r_s)}}{4\pi} = \frac{1}{4\pi r_s}$$

charged BH: Reissner-Nordström

$$f(r) = 1 - \frac{2GM}{r} + \frac{G(q^2 + p^2)}{r^2}$$

$$S = \frac{A}{4G_4} = \frac{r_s^2 \Omega_2}{4G_4} = 4\pi G_4 M^2$$

$$R_{\mu\nu\alpha\beta} R^{\mu\nu\alpha\beta} = \frac{19 r_s^2}{r^6}$$

true sing at $r=0$

Strings & 2d gravity

10'
(to p. 10!)

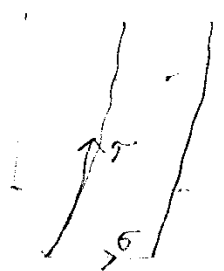
Relativistic ple:

$$S = -mc^2 \int_{\tau_1}^{\tau_2} d\tau = -mc \int_{t_1}^{t_2} dt \sqrt{-\eta_{\mu\nu} \dot{x}^\mu \dot{x}^\nu} = -mc \int_{t_1}^{t_2} dt \sqrt{1 - \frac{1}{c^2} \left(\frac{dx}{dt}\right)^2}$$

"reparam. invariant"

NR
= const
+ $\frac{1}{2} m \bar{v}^2$
sign OK!

String:



$$x^\mu = X^\mu(\tau, \sigma)$$

$$ds^2 = G_{\mu\nu} dx^\mu dx^\nu = G_{\mu\nu} \frac{\partial x^\mu}{\partial \sigma^a} \frac{\partial x^\nu}{\partial \sigma^b} d\sigma^a d\sigma^b \equiv h_{ab}$$

$$S_{NG} = -T \int d^2\sigma \sqrt{-\det h_{ab}} = S[X^\mu, X^\nu]$$

$$= -T \int d^2\sigma \sqrt{-h} \frac{1}{2} h^{ab} h_{ab}$$

$$= -\frac{T}{2} \int d^2\sigma \sqrt{-h} h^{ab} \left[\eta_{\mu\nu} \partial_a X^\mu \partial_b X^\nu \right]$$

$$\Rightarrow \text{EDM} \quad T_{ab} = T \left[\partial_a X \cdot \partial_b X - \frac{1}{2} h_{ab} h^{cd} \partial_c X \partial_d X \right] - \bar{\Psi}^\mu i g^{ab} \partial_a \Psi_\nu$$

classical $T^a_a = 0$ superstring

Much of string theory is what happens

- in quantum theory
- when $\eta_{\mu\nu} \rightarrow G_{\mu\nu}(X)$; background indep

Sean Carroll

Other coords:

$$-f(r) dt^2 + \frac{dr^2}{f(r)} + r^2 d\Omega^2$$

factor out f by $\frac{dr}{f(r)} = dr^*$

$$\frac{dr^2}{f} = f \left(\frac{dr}{f} \right)^2$$

$$ds^2 = +f(r) [-dt^2 + 2dr_*^2] + r^2(r_*) d\Omega^2$$

transformation in one step

Further: $t_+ = t + r^*$ $dt = dt_+ - dr_* = dt_+ - \frac{dr}{f(r)}$
 $= dt_+ + \frac{dr}{f(r)}$

$$ds^2 = f [-dt_+^2 + 2 dt_+ dr_* - dr_*^2 + dr_*^2] + r^2 d\Omega^2$$

$$ds^2 = -\left(1 - \frac{r_s}{r}\right) dt_+^2 + 2 dt_+ dr + r^2 d\Omega^2$$

Eddington-Finkelstein

$ds = 2 dr$ const
 $\begin{cases} dt_+ = 0 \\ \frac{dr}{dt_+} = 1 - \frac{r_s}{r} \end{cases} \begin{cases} < 0 \text{ } r < r_s \\ > 0 \text{ } r > r_s \end{cases}$

original r !

Pure AdS₅ $\begin{cases} x^\mu \rightarrow \lambda x^\mu \\ z \rightarrow \lambda z \end{cases}$

$$ds^2 = \frac{r^2}{z^2} (-dt^2 + d\vec{x}^2 + dz^2)$$

pull out one coord⁻²!
 need a scale L (length⁻¹)

$$r = \frac{r^2}{z} \quad \begin{cases} x^\mu \rightarrow \lambda x^\mu \\ r \rightarrow \frac{1}{\lambda} r \end{cases}$$

"Pure" means maximal symmetry; AdS_d:

$$R_{\mu\nu\alpha\beta} = \# (g_{\mu\alpha} g_{\nu\beta} - g_{\mu\beta} g_{\nu\alpha}) = -\frac{1}{L^2} (g_{\mu\alpha} g_{\nu\beta} - g_{\mu\beta} g_{\nu\alpha})$$

$$R_{\nu\beta} = g^{\mu\alpha} R_{\mu\nu\alpha\beta} = \# (d+1) g_{\nu\beta} - g_{\nu\beta} = \# \cdot d g_{\nu\beta} = \frac{-d}{L^2} g_{\nu\beta}$$

AdS_5 is the surface

$$-t_1^2 - t_2^2 + x_1^2 + x_2^2 + x_3^2 + x_4^2 = -L^2$$

can be parametrised by $t_i = \frac{L^2}{z} \left(\frac{t}{L} + \frac{y^2 + z^2 - t^2}{2z} \right)$

with

$$-\infty < t < \infty$$

$$0 < z < \infty$$

$$-\infty < y_i < \infty$$

$$\begin{cases} \frac{t}{L} + x_4 = -\frac{r^2}{z} \\ \frac{t}{L} - x_4 = -\frac{y^2 + z^2 - t^2}{z} \end{cases}$$

$$\begin{aligned} & \left[\frac{L^2}{z^2} (-t^2 + y^2) - \frac{1}{4z^2} (L^2 + y^2 + z^2 - t^2)^2 - (-L^2 + y^2 + z^2 - t^2)^2 \right] \\ & \quad - 4L^2 \left(\frac{y^2 + z^2 - t^2}{z} \right)^2 \\ & = \frac{L^2}{z^2} [-t^2 + y^2 - (y^2 + z^2 - t^2)] = -L^2 \quad !! \end{aligned}$$

in the space with flat metric

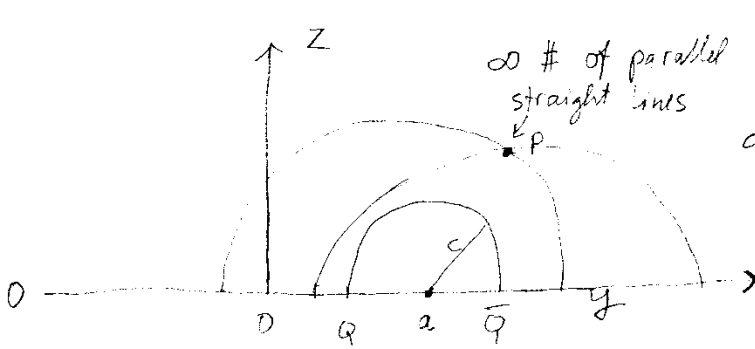
$$\begin{aligned} ds^2 &= -dt_1^2 - dt_2^2 + dx_1^2 + \dots + dx_4^2 \\ &= \frac{r^2}{z^2} (-dt^2 + dy^2 + dz^2) \quad \leftarrow \text{after some work} \end{aligned}$$

The surface and thus AdS_5 has the symmetry $O(2, 4)!$
 = conformal group in 4d!

BC really is $AdS_5 \otimes S^1$

The $N=4$ superconformal rotational symmetry is mapped to rotating S^5 in 6d space!

Poincare plane; non-Eucl. geometry:



Why this?

$$ds^2 = \frac{L^2}{z^2} (dz^2 + dy^2)$$

Geodesics:
$$S = \int ds = L \int \sqrt{\frac{dz^2 + dy^2}{z^2}} = L \int dz \frac{1}{z} \sqrt{1 + y'(z)^2}$$

$$= L \int \frac{dy}{z} \sqrt{1 + z'(y)^2} \quad \int dz L(y'(z))$$

$$= \int dy L(z(y), z'(y)) \quad \text{Simpler!}$$

$$\frac{\partial L}{\partial y} - \frac{d}{dy} \frac{\partial L}{\partial z'} = 0$$

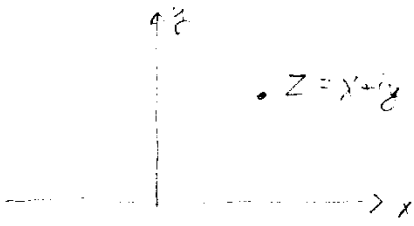
$$\frac{\partial L}{\partial z'} = \frac{y'(z)}{z \sqrt{1 + y'^2}} = \text{const} = \frac{1}{c}$$

$$z^2 \left[1 + \left(\frac{dz}{dy} \right)^2 \right] = c^2 \left(\frac{dz}{dy} \right)^2 = \frac{c^2}{z^2} - 1$$

$$\frac{dz}{dy} = \frac{1}{z} \sqrt{c^2 - z^2}$$

$$\frac{d \frac{1}{2} z^2}{\sqrt{c^2 - z^2}} = dy \Rightarrow \boxed{(y-a)^2 + z^2 = c^2}$$

• In complex coordinates:



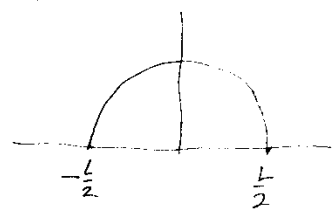
$$d_{\text{Poincaré}}(z_1, z_2) = \left| \frac{z_1 - z_2}{z_1 - \bar{z}_2} \right|$$

$$g(z_1, z_2) = \log \frac{|z_1 - \bar{z}_2| + |z_1 - z_2|}{|z_1 - \bar{z}_2| - |z_1 - z_2|}$$

distance between z_1, z_2 in metric $\frac{1}{y} z(dx^2 + dy^2)$

• Energy in z plane (Exercise):

$$S = \frac{T}{2\pi\alpha'} (\Delta \text{time}) \cdot \int_{-L/2}^{L/2} dx \frac{1}{z^2} \sqrt{1 - \frac{z^4}{z_0^4} + z'(x)^2}$$



4. Master formula, correlators

$$\frac{1}{Z} \int \mathcal{D}\bar{\phi} \phi e^{-\int d^4x \left[\frac{1}{2} g^{\mu\nu} \partial_\mu \phi \partial_\nu \bar{\phi} + \varphi_0(x) \bar{\phi}(x) \right] + \varphi_0(x) \bar{\phi}^\Delta + \dots}$$

mass dims ←

↑ quantum field = $\mathcal{O}(x)$

$$= \frac{1}{Z} \int d^5x \sqrt{-g} \left[\frac{1}{2} g^{\mu\nu} \partial_\mu \varphi \partial_\nu \varphi - \frac{1}{2} m^2 \varphi^2 \right]$$

pure AdS₅

does NOT fix renormalization

↑ needed to fix dimensions! find $m^2 < 0$!

$m^2 = \Delta(\Delta - 4)$!

For p-form bulk fields $(\Delta - p)(\Delta + p - 4) = m^2$

(2d) mom. space,

$$\varphi(x, z) = \int \frac{d^4k}{(2\pi)^4} e^{ikx} \varphi(k, z) \varphi_0(k) \rightarrow z^{4-\Delta} \varphi_0(x)$$

$$\varphi(k, z \rightarrow 0) = z^{4-\Delta} (1 + \mathcal{O}(z^2))$$

$$\int d^4x \varphi_0(x) \bar{\phi}(x) = \int \frac{d^4k}{(2\pi)^4} \varphi_0(k) \bar{\phi}(-k) = \int \frac{d^4k_1 d^4k_2}{(2\pi)^8} \delta^4(k_1 + k_2) \varphi_0(k_1) \bar{\phi}(k_2)$$

Fact 1:

$$\int d^5x \sqrt{-g} \left[\frac{1}{2} g^{\mu\nu} \partial_\mu \varphi \partial_\nu \varphi + \frac{1}{2} m^2 \varphi^2 \right]$$

$$= \int d^5x \frac{1}{2} \sqrt{-g} \left[(-\square + m^2) \varphi \right] \varphi + \frac{1}{2} \int d^5x \partial_N \sqrt{-g} \partial^N \varphi \cdot \varphi$$

= 0 for soln's of EOM / "on shell"

only this remains disappears, boundary contributions; holography

Fact 2:

$$\varphi(k, z) \equiv z^2 U(z)$$

$$\Rightarrow (\square - m^2) \varphi = \frac{z^2}{L^2} \left\{ z^2 U'' + z U' - \left[k^2 z^2 + 4 + (mL)^2 \right] U \right\}$$

↑ $-\frac{1}{2} + \frac{1}{2} = 0$ Standard Bessel eq.

$$\Rightarrow \varphi(k, z) \sim z^2 I_\nu(kz) \quad \text{or} \quad z^2 K_\nu(kz) \quad \int K_\nu(z) = K_{\nu-1/2}$$

$k \equiv \sqrt{k^2}$ $\nu = \sqrt{4 + (mL)^2}$ ($I_m = I_{-m}$)

Euclidian or spacelike

on shell action

$\frac{L^5}{z^5} g^{zz} = \frac{z^2}{L^2}$ normalised soln of $(\square - m^2)\varphi = 0$

Witten, hep-th/9802150
3700 citations

Basically Gubser-Klebanov-Polyakov, hep-th/9802109, 3200 citations

$$\frac{1}{16\pi G_5} \int d^5x \partial_N [\sqrt{-g} g^{NM} \partial_M \varphi \cdot \varphi]$$

$$= \frac{1}{16\pi G_5} \int \frac{d^4k}{(2\pi)^4} \left\{ \int_{\epsilon}^{\infty \rightarrow z_0} \frac{L^3}{z^3} \varphi_z^*(k, z) \varphi(-k, z) + \int_{\epsilon}^{\infty} dz \frac{L^3}{z^3} (-k^2) \varphi(k, z) \varphi(-k, z) \right\}$$

- need convergence for $z \rightarrow \infty$!

"compactify μ -directions so that there is no boundary there!"

- for finite T there will be upper limit $z_0 \sim \frac{1}{T}$

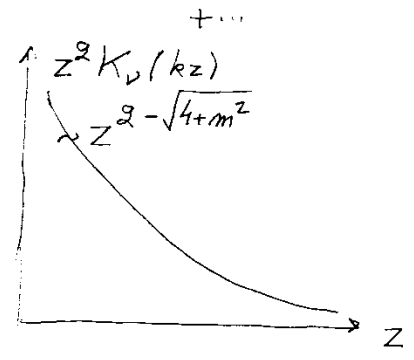
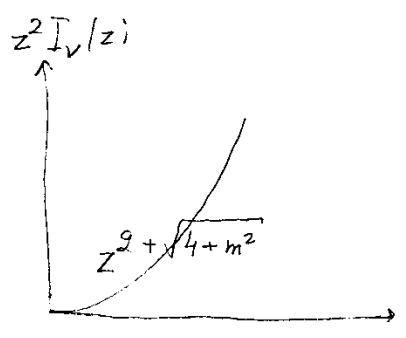
Two solutions:

Compare J's Noether current in scalar ED:

$$I_\nu(z) = \left(\frac{z}{2}\right)^\nu \frac{1}{\Gamma(1+\nu)} \left[1 + \frac{1}{1+\nu} \left(\frac{z}{2}\right)^2 + \dots \right]$$

$$J_\mu = \frac{g}{2i} (\varphi^* \partial_\mu \varphi - \partial_\mu \varphi^* \varphi)$$

$$K_\nu(z) = \frac{1}{2} \left(\frac{z}{2}\right)^\nu \left[\Gamma(\nu) - \Gamma(\nu-1) \frac{z^2}{4} + \dots \right] + (-1)^{1+\nu} \log \frac{z}{2} \cdot I_\nu(z)$$



Have to choose $z^2 K_\nu(kz)$ ($\rightarrow 0$ at $z \rightarrow \infty$)

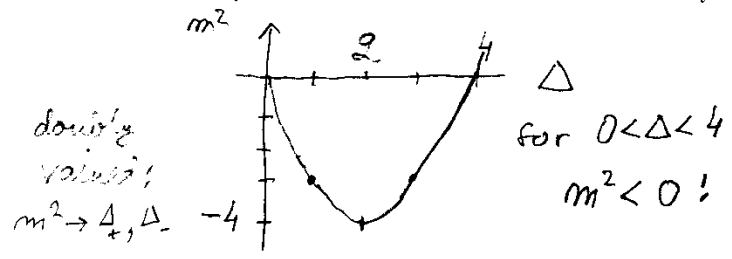
$$\nu = |\Delta - 2|$$

with normalisation

$$C = \frac{g}{\Gamma(\nu)} \left(\frac{k}{2}\right)^\nu$$

$$\varphi(k, z \rightarrow 0) = 1 \cdot z^{4-\Delta} = C \cdot \left(\frac{\Gamma(\nu)}{2} \left(\frac{kz}{2}\right)^{-\nu}\right) \cdot z^2 = C z^2 K_\nu(kz)$$

$$\Rightarrow 4 - \Delta = 2 - \sqrt{4+m^2} \Rightarrow \sqrt{4+m^2} = \Delta - 2 \Rightarrow m^2 = \Delta(\Delta - 4)$$



$$S_{\text{grav}} = \frac{L^3}{16\pi G_5} \int \frac{d^4 k_1}{(2\pi)^4} \frac{d^4 k_2}{(2\pi)^4} (2\pi)^4 \delta(k_1+k_2) \varphi_0(k_1) \varphi_0(k_2) \cdot \frac{1}{g} \frac{\varphi(k_1, z) \varphi_2'(k_2, z)}{z^3} \Big|_{z=\epsilon}$$

$$\Rightarrow \langle \mathcal{O}(k_1) \mathcal{O}(k_2) \rangle = \frac{L^3}{16\pi G_5} \frac{\varphi(k_1, z) \varphi_2'(k_2, z)}{z^3} \Big|_{z=\epsilon} \cdot (2\pi)^4 \delta^4(k_1+k_2)$$

$$\varphi(k, z) = \frac{g}{\Gamma(\nu)} \left(\frac{k}{2}\right)^\nu z^2 K_\nu(kz)$$

Ex $\Delta=3 \quad \nu=1 \quad \varphi(k, z) = k z^2 K_1(kz)$

$$\begin{cases} \varphi(k, z) = z + \frac{k^2}{4} (2 \log \frac{kz}{2} + 2\gamma - 1) z^3 + \mathcal{O}(z^5) \\ \frac{d}{dz} \varphi(k, z) = 1 + \frac{k^2}{4} (6 \log \frac{kz}{2} + 6\gamma - 1) z^2 + \mathcal{O}(z^4) \end{cases}$$

$$\varphi_k \varphi'_k = z + z^3 \cdot 2k^2 \left(\log \frac{kz}{2} + \gamma - \frac{1}{4} \right) + \dots + i \times (\dots)$$

$$\langle \mathcal{O}(k_1) \mathcal{O}(k_2) \rangle = \left[\frac{1}{\epsilon^2} + \frac{N_c^2}{8\pi^2} 2k^2 \left(\log \frac{k\epsilon}{2} + \gamma - \frac{1}{4} \right) \right] \cdot (2\pi)^4 \delta^4(k_1+k_2)$$

non-staircase
th/020505: $\frac{1}{\epsilon^2} \cdot k^2 \log k + \dots$

⇒ p. 16' for finite T & Minkowski Lorentz invariance

$$\langle \mathcal{O}(x_1) \mathcal{O}(x_2) \rangle \sim \int d^4 k_1 d^4 k_2 \delta^4(k_1+k_2) k^2 e^{ik_1 x_1 - ik_2 x_2} \sim \int d^4 k k^2 e^{ik(x_1-x_2)} \sim \frac{1}{|x_1-x_2|^6}$$

dimensional

A particularly interesting $\Delta = 3$ operator is \bar{g}_2 :

$$\int \mathcal{D}x \, m_2 \bar{g}_2(x)$$

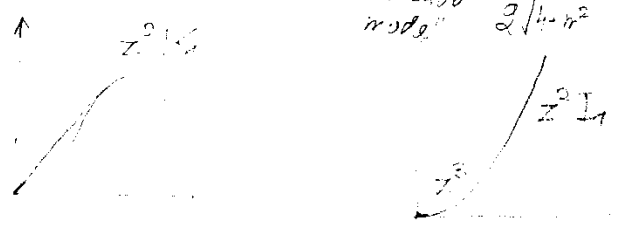
$$\phi_0(x) = m_2 \quad \phi_2(x) = m_2 z^{-2} e^{2\sigma(x)}$$

$$\varphi(z, z) = \sum m_2 + \dots \bar{g}_2 \cdot z^3 + \dots$$

"non-normalisable mode" has less powers of z when $z \rightarrow 0$ $\Delta > 2$ the full gravity dual is determined by $\langle \bar{g}_2 \rangle$

$$= z^{2-\Delta} \phi_0(x) + z^{-1} \frac{\langle \mathcal{O} \rangle}{2\Delta-4}$$

Klebanov
- Witten



\Rightarrow Holographic model for neutrinos: EKSL p1/0501128

$$\int \mathcal{D}x \, \bar{\psi} \gamma^\mu \partial_\mu \psi - A \Gamma^2 \psi$$

conserved
current
(cons. of total momentum)

$$A_\mu^\alpha(x, z) \quad \int \mathcal{D}x \, \bar{\psi} \gamma^\mu \psi$$

gauge inv. to
bulk energy

$$\int \mathcal{D}x \, T_{\mu\nu} = \delta g_{\mu\nu}$$

$$\bar{T}_{\mu\nu} = 0$$

5.57 flat black hole (in many forms)

$$ds^2 = \frac{L^2}{z^2} \left(1 - \frac{z^4}{z_0^4} \right) dt^2 + d\vec{x}^2 + \frac{dz^2}{1 - \frac{z^4}{z_0^4}}$$

inside BH

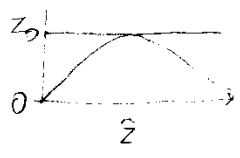
$$R_{\text{Hof}}^2 = \frac{1}{L^4} (40 + 72 \frac{z_0^8}{z_0^8}) \quad T_H = \frac{1}{\pi z_0} \quad S = \frac{A}{4G_5} = V_3 \frac{1}{4G_5 z_0^3}$$

Extrinsic curvature of a surface $z=z_0$

$$= \frac{L^2}{z^2} \left(-\frac{1 - \frac{z^4}{z_0^4}}{1 + \frac{z^4}{4z_0^4}} dt^2 + \left(1 + \frac{z^4}{4z_0^4} \right) d\vec{x}^2 + dz^2 \right)$$

$$K = -\frac{2}{L} \left[\frac{1}{\sqrt{1 - \frac{z^4}{z_0^4}}} + \sqrt{1 - \frac{z^4}{z_0^4}} \right]$$

$z = \frac{\tilde{z}^2}{1 + \frac{\tilde{z}^4}{4z_0^4}}$ FG form $g_{\mu\nu} dx^\mu dx^\nu + dz^2$ only outside of BH!

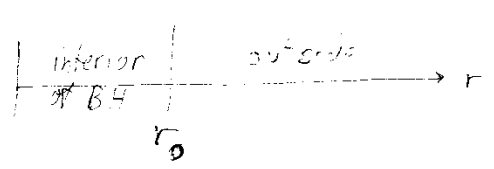


$g_{\mu\nu}(x, z=0) = g_{\mu\nu}(x) = \text{boundary metric} = (-1, +++)$

$$= \frac{L^2}{z^2} \left[-\left(1 - \frac{z^4}{z_0^4} \right) dt^2 + 2 dz dz + d\vec{x}^2 \right]$$

Eddington-Finkelstein-type

$$z = \frac{L^2}{r} \quad z_0 = \frac{L^2}{r_0} \quad dt = dt_+ + \frac{dz}{1 - \frac{z^4}{z_0^4}}$$



$$= -\frac{r^2}{L^2} \left(1 - \frac{r_0^4}{r^4} \right) dt_+^2 + 2 dt_+ dr + \frac{r^2}{L^2} d\vec{x}^2$$

$$= -\left[\frac{r^2}{L^2} \left(1 - \frac{r_0^4}{r^4} \right) \nu_\mu \nu_\nu dx^\mu dx^\nu - 2 \nu_\mu dx^\mu + \frac{r^2}{L^2} \nu_\mu \nu_\nu dx^\mu dx^\nu \right]$$

$X^\mu = (t_+, x^i) \quad u^\mu = (1, \vec{0})$ $r_0 \rightarrow r_0(x)$ $u^\mu \rightarrow u^\mu(x)$

interesting "Boosted black brane" recent work! $\text{Miyoshi et al } 0712.245$ + many related papers

$$= -r^2 a(t_+, r) dt_+^2 + 2 dt_+ dr + r^2 c(t_+, r) d\eta^2 + r^2 c(t_+, r) d\vec{x}_T^2$$

t-dep. soln with in- and outside of horizon

Nakamura et al D807.3797

Solving $R_{MN} = -\frac{4}{L^2} g_{MN}$

Use Mathematica or Maple!
 see Problems

$\mathbb{R}^4 \rightarrow \mathbb{R}^4$ coord: r, η, x^2, x^3, z

$$ds^2 = \left[\frac{1}{r^2} - \frac{\left(1 - \frac{r_0 z^4}{3 r^{4/3}}\right)^2}{1 + \frac{r_0 z^4}{3 r^{4/3}}} \right] dr^2 + \left(1 + \frac{r_0 z^4}{3 r^{4/3}}\right) (r^2 d\eta^2 + dx^2) + dz^2$$

solves $R_{MN} + \frac{4}{L^2} g_{MN} = 0$ at large r

"time dependent" $z_0 = \frac{1}{\sqrt{11}} = \frac{2^{-1/2}}{e_0} r^{1/3}$ "

corrections $\sim \frac{1}{r^{1/3}} \sim \frac{1}{r^{2/3}} \Rightarrow$ need anisotropy η
 & other 2nd order hydro coefficients

Same in EF type coordinates (Hokanson...)

$$r_0 = \frac{r^2}{r_0} = \frac{r_0}{20} r^{1/3}$$

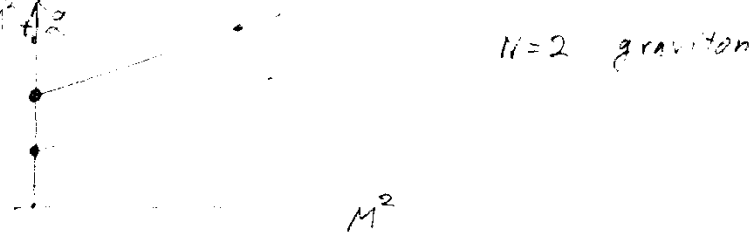
also needed!

$$ds^2 = -\frac{r^2}{L^2} \left[1 - \left(\frac{r_0}{r r_0^{1/3}}\right)^4 \right] dr^2 + 2 dr_0 dr + r^2 \left(1 + \frac{1}{r r_0}\right)^2 r^2 d\eta^2 + r^2 dx^2$$

Quadratic
Some steps on string $\hbar \rightarrow$ class gr

ST: $S = -\frac{T}{2} \int d\sigma \sqrt{-h} \left(\dot{x}^\mu \dot{x}^\nu \eta_{\mu\nu} + \dots + \text{fermionic!} \right)$

- Has excitations $\alpha_{-1}, \alpha_0, \alpha_1 \dots \propto \frac{1}{\alpha'} (1-g) N = 0, 1, \dots$
 $11 = d + M^2 \alpha'^2$



- For $d \rightarrow 0$ $M^2 \rightarrow \infty$ unless $M=0$

\Rightarrow strings \rightarrow point \Rightarrow field theory,
 T-dual Supergravity

$S = \frac{1}{16\pi G_{10}} \int d^5x \sqrt{-g} e^{-2\phi} \left[R + 4(\partial\phi)^2 - \frac{1}{4} F^2 + \dots \right]$



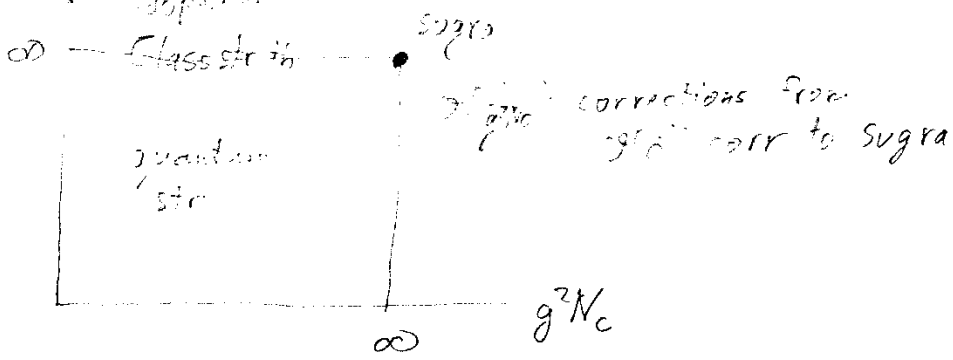
$16\pi G_{10} = (2\pi)^7 \alpha'^4 g_s^2$
 $g_s = \frac{g^2}{4\pi}$

$L^2 = \alpha' \sqrt{g^2 N_c}$

$\Rightarrow \frac{1}{4\pi G_5} = \frac{N_c^2}{g^2 \pi^2}$

$\frac{1}{16\pi G_{10}} \int d^5x \sqrt{-g} = \frac{1}{16\pi G_{10}} \int d^4y \sqrt{-g} \int d^1x \sqrt{-g_5} = \frac{1}{16\pi G_5}$

\uparrow no string loops, corrections



6. Finite T matter on the boundary

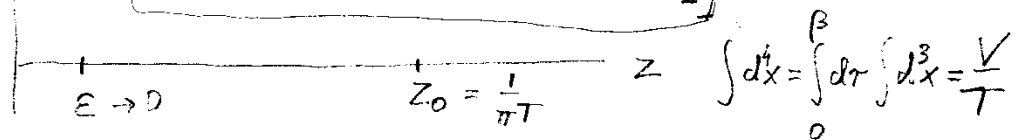
(1) $Z_{\text{CFT}} = e^{-\frac{E}{T}} = e^{\frac{P(T)V}{T}} = \int \mathcal{D}\varphi e^{-S_{\text{EL}}[\varphi; g_{\mu\nu}^{(0)}(x)]}$ no ext current!
 Master formula: $\varphi(\beta, \vec{x}) = \varphi(0, \vec{x})$

$= e^{-S_{\text{Grav}}[G_{MN}(x, z; z_0)]}$ the G_{MN} giving the $\uparrow T$
 min to ∞ (regularised!!) smallest S_{Grav} !

(2) Determine full $T_{\mu\nu}(x)$

(1) For flat 5d BH: $a = 1 - \frac{z^4}{z_0^4}$
 $ds^2 = \frac{L^2}{z^2} \left\{ -a(z) dt^2 + dx^2 + \frac{dz^2}{a(z)} \right\}$ $\sqrt{-g} = \frac{L^5}{z^5}$ $\sqrt{-g} = \frac{L^4}{z^4} \sqrt{a}$

$S_{\text{Grav}} = \frac{1}{16\pi G_5} \left[\int d^5x \frac{L^5}{z^5} \frac{-8}{L^2} - \int d^4x \frac{L^4}{z^4} \sqrt{a} \left[-\frac{4}{L} \left(\frac{1}{\sqrt{a}} + \sqrt{a} \right) + \frac{6}{L} + \frac{L}{2} R(x) \right] \right]$
 $\left[\int d^5x \frac{L^5}{z^5} \frac{-8}{L^2} \right] - \left[\int d^4x \frac{L^4}{z^4} \sqrt{a} \left[-\frac{4}{L} \left(\frac{1}{\sqrt{a}} + \sqrt{a} \right) + \frac{6}{L} + \frac{L}{2} R(x) \right] \right] = 2K$



$= \frac{L^3}{16\pi G_5} \frac{V}{T} \left\{ \int_{\epsilon}^{z_0} dz \frac{-8}{z^5} + \frac{4}{\epsilon^4} \left(1 + 1 - \frac{\epsilon^4}{z_0^4} \right) - \frac{6}{\epsilon^4} \sqrt{1 - \frac{\epsilon^4}{z_0^4}} + \frac{1}{2} L^2 R(x) \right\}$
 intrinsic curv. of $z = \epsilon$

Note: also bulk action contributes, EFT does not

$\left\{ \frac{2}{z_0^4} - \frac{2}{\epsilon^4} + \frac{8}{\epsilon^4} - \frac{4}{z_0^4} - \frac{6}{\epsilon^4} + \frac{3}{z_0^4} \right\}$

signs to be adjusted!

$\Rightarrow P(T) = \frac{L^3}{16\pi G_5} \cdot \frac{3-4+3}{z_0^4} = \frac{\pi^2 N_c^2}{8} T^{-4}$ \uparrow CT tuned to cancel $\frac{1}{\epsilon^4}$

$\frac{L^3}{4\pi G_5} = \frac{N_c^2}{2\pi^2}$

very subtle !!
 best to do with diff geo. m

Do a coord. transf. in bulk

→ may change $g_{\mu\nu}^{(0)}(r)$ by a conf. transf

Does not matter if bound. theory is conf. invariant!

Ex: Work out $p(r)$ from

$$g_{MN} = \frac{L^2}{z^2} \begin{pmatrix} -\frac{a^2}{b} & & & \\ & b & & \\ & & b & \\ & & & b \end{pmatrix} \quad a = 1 - \frac{z^4}{4z_0^4} \quad z_0 = \frac{1}{\pi T}$$

$$b = 1 + \frac{z^4}{4z_0^4}$$

$$\sqrt{-G} = \frac{L^5}{z^5} ab \quad \left\{ dx^4 \hat{=} \frac{V}{T} \quad \sqrt{-g} = \frac{L^4}{z^4} ab \quad K = \frac{1}{L} \left(-4 + \frac{za'}{a} + \frac{zb'}{b} \right) \right.$$

$$S_{gr} = \frac{L^3}{16\pi G_5} \cdot \frac{V}{T} \left\{ \int_{\epsilon}^{z_0\sqrt{2}} \frac{dz \sqrt{-g}}{z^5} \left(1 - \frac{z^8}{16z_0^8} \right) - \frac{1}{L^3} \sqrt{-g} \left[2K + \frac{6}{L} + \frac{1}{2} R(g) \right] \right\}$$

$$\int_{\epsilon}^{z_0\sqrt{2}} \left(2z^{-4} + \frac{2z^4}{16z_0^8} \right) - \frac{1}{\epsilon^4} [-8 + 6] + O(\epsilon^4)$$

$$\frac{2}{4z_0^4} - \frac{2}{\epsilon^4} + \frac{4}{8z_0^4}$$

cancel

$$= \frac{L^3}{16\pi G_5} \frac{1}{z_0^4} \cdot \frac{V}{T} \quad \text{just as for standard flat BH,}$$

but now all comes from bulk action!

$$\frac{1}{L^3} \sqrt{-g} \left(2K + \frac{6}{L} \right) = \frac{1}{z^4} [-8(a-b) + 2ab]$$

7. Comparing actions: Hawking-Page transition
 Herzog hep-th/0608151 toy model

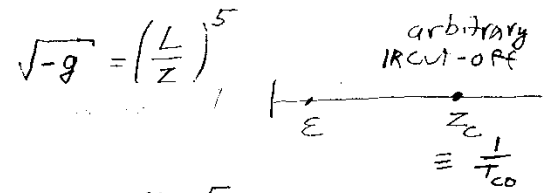
$$S = \frac{1}{16\pi G_5} \int d^5x \sqrt{-g} \left(R + \frac{19}{L^2} \right) = -\frac{1}{16\pi G_5 L^2} \int d^5x \sqrt{-g} \cdot 8$$

$$R = -\frac{20}{L^2}$$

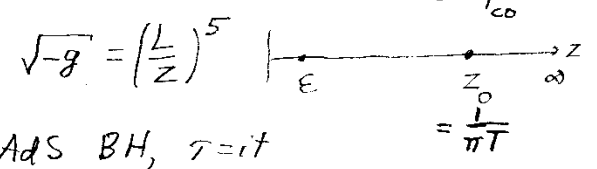
Boundary term is neglected in this toy model! "hard wall model"

Two metrics: pure AdS

(1) $ds^2 = \frac{L^2}{z^2} (d\tau^2 + d\vec{x}^2 + dz^2)$
 ↑ Euclidian time



(2) $ds^2 = \frac{L^2}{z^2} \left[a(z) d\tau^2 + d\vec{x}^2 + \frac{dz^2}{a(z)} \right]$
 $a(z) = 1 - \frac{z^4}{z_0^4}$ Flat AdS BH, $\tau = it$



$$z_{\text{CFT}} = \langle 1 \rangle = e^{-\frac{V}{T}} = e^{+P \frac{V}{T}} = e^{S_{\text{grav}}}$$

$$\int d^5x = \int d\tau \int d^3x \int dz$$

$\Rightarrow V_3$

value diverges at $z \rightarrow 0$; difference will be finite

Limits are obvious for (2) $\int dz 8z^{-5} = -2z^{-4}$

$$S_{(2)}/V_3 = -\frac{L^3}{16\pi G_5} \int_0^{\beta} d\tau \int_{\epsilon}^{\min(z_0, z_c)} \frac{8 dz}{z^5} = -\frac{L^3}{16\pi G_5 T} \left(\frac{9}{\epsilon^4} - \frac{9}{(\min(z_0, z_c))^4} \right)$$

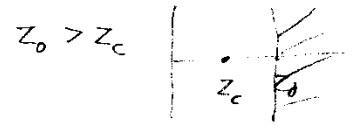
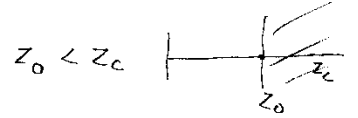
No. bdy term!?

! To compare actions at same T demand same physical period
 $\Delta\tau_{(1)} = \sqrt{a(\epsilon)} \Delta\tau_{(2)}$

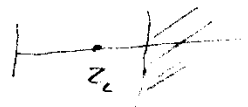
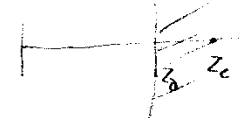
$$S_{(1)}/V_3 = -\frac{L^3}{8\pi G_5} \frac{1}{T} \sqrt{1 - \frac{\epsilon^4}{z_0^4}} \left(\frac{1}{\epsilon^4} - \frac{1}{z_c^4} \right)$$

$$\frac{1}{\epsilon^4} - \frac{1}{z_c^4} - \frac{1}{2z_0^4}$$

$$\begin{aligned} [S_{(2)} - S_{(1)}] / V_3 &= + \frac{L^3}{8\pi G_5 T} \left[\frac{1}{z_0^4} - \frac{1}{z_c^4} - \frac{1}{2z_0^4} \right] \\ &= \frac{P_{(2)} - P_{(1)}}{T} \left[\frac{1}{z_c^4} - \frac{1}{z_c^4} - \frac{1}{2z_0^4} \right] \end{aligned}$$



$$\begin{cases} \frac{L^3}{8\pi G_5 T} \left(\frac{1}{2z_0^4} - \frac{1}{z_c^4} \right) \\ \frac{L^3}{8\pi G_5 T} \left(-\frac{1}{2z_0^4} \right) \end{cases}$$

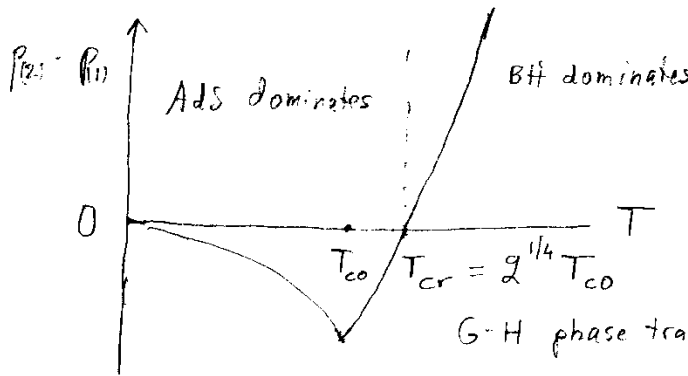


one can also soften the cut-off by $\theta(z-z_c) \Rightarrow e^{-cz^2}$ "soft model"

Put here, for concreteness $\frac{L^3}{4\pi G_5} = \frac{1}{2\pi^2} \frac{1}{z_c^2}$

$$\frac{1}{z_0} = \pi T$$

$$\Rightarrow P_{(2)} - P_{(1)} = \begin{cases} \frac{\pi^2 N_c^2}{4} \left(\frac{T^4}{9} - T_{co}^4 \right) & T > T_{co} \\ \left(-\frac{T^4}{9} \right) & T < T_{co} \end{cases}$$

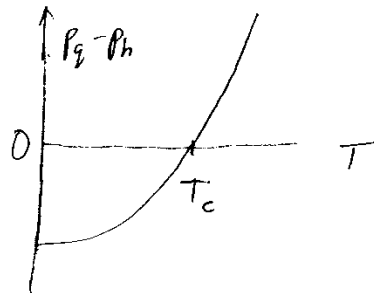


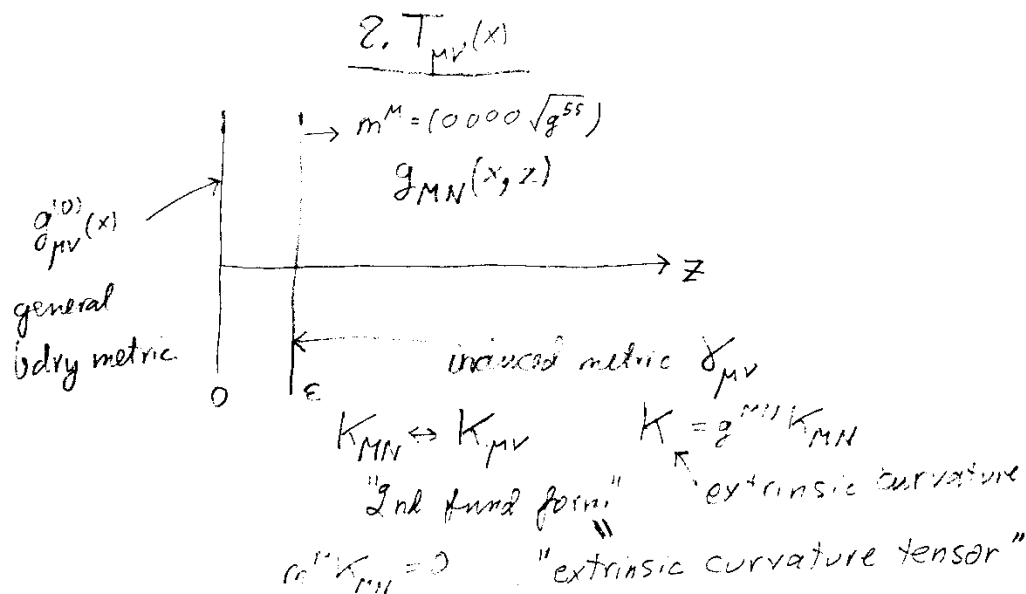
correct coefficient for $p(T) = aT^4$ for $N=4$ hot SYM!

5d bulk transition!

Compare bag model for hot QCD:

$$\begin{cases} P_g = a_g T^4 - B \\ P_h = a_h T^4 \\ a_g > a_h \end{cases}$$





$$S = \frac{1}{16\pi G} \int d^5x \sqrt{-g} \left(R + \frac{2}{L^2} \right) - \int d^4x \sqrt{-\gamma} 2K + S_{CT}$$

$$\Rightarrow \langle T_{\mu\nu} \rangle = \frac{-2\delta S}{\sqrt{-\gamma} \delta \gamma^{\mu\nu}} \stackrel{L=1}{=} \frac{1}{8\pi G_5} \left(K_{\mu\nu} - K \gamma_{\mu\nu} - 3\gamma_{\mu\nu} + \frac{d-1}{2} [R_{\mu\nu}(\gamma) - \frac{1}{2} R(\gamma) \gamma_{\mu\nu}] \right) \Big|_{z \rightarrow 0}$$

CT's

variation of $\int d^5x$

just gives EDM \Rightarrow no contrib!

"simply" has to be evaluated!

K_{MN} = extrinsic

$K = g^{MN} K_{MN} = \text{tr}[\text{Inverse}[\text{metric}] \cdot \text{extrinsic}]$

$\sqrt{-\gamma} = (\det(\text{induced}))^{1/2}$

Some generat^les:

- Write $g_{MN} = \frac{L^2}{z^2} \begin{pmatrix} g_{\mu\nu}(x,z) & 0 \\ 0 & 1 \end{pmatrix}$ "gauge choice"
FG form

- EOM $R_{MN} + \frac{4}{L^2} g_{MN}$ becomes $(L = \frac{2}{z^2})$
Skenderis's th/0009230 eq 2.5

$$\begin{cases} \mu\nu & R_{\mu\nu}(g) - 2g'_{\mu\nu} - g'^{\alpha\beta} g_{\alpha\beta} g_{\mu\nu} + z^2 [2g''_{\mu\nu} - 2g'_{\mu\alpha} g'^{\alpha\nu} + g'^{\alpha\beta} g'_{\beta\nu}] = 0 \\ \mu z & \Delta^M(g) (g'_{\mu\nu} - g'^{\alpha\beta} g_{\alpha\beta} g_{\mu\nu}) = 0 \\ z z & g^{\mu\nu} g''_{\mu\nu} - \frac{1}{2} g'^{\alpha\beta} g'_{\alpha\beta} = 0 \end{cases}$$

Try to solve: $g_{\mu\nu}(x,z) = g_{\mu\nu}^{(0)}(x) + g_{\mu\nu}^{(2)}(x)z^2 + g_{\mu\nu}^{(4)}(x)z^4 + \dots$
 $g'_{\mu\nu}(x,z) = g_{\mu\nu}^{(2)}(x) + 2g_{\mu\nu}^{(4)}(x)z^2 + \dots$
 $g''_{\mu\nu}(x,z) = 2g_{\mu\nu}^{(4)}(x) + \dots$

$z^2 = 0$: $g_{\mu\nu}^{(0)} \left[R_{\mu\nu}(g^{(0)}) - 2g_{\mu\nu}^{(2)} - g'^{\alpha\beta} g_{\alpha\beta}^{(0)} g_{\mu\nu}^{(0)} \right] = 0$

$R(g^{(0)}) - 2\text{Tr}g^{(2)} - 4\text{Tr}g^{(2)} = 0 \quad \text{Tr}g^{(2)} = \frac{1}{6}R(g^{(0)})$

($\mu\nu$)

$g_{\mu\nu}^{(2)} = \frac{1}{2} \left[R_{\mu\nu}(g^{(0)}) - \frac{1}{6}R(g^{(0)})g_{\mu\nu}^{(0)} \right]$

For flat bdry metric $g_{\mu\nu}^{(2)} = 0!$

($z z$)

$2\text{Tr}g^{(4)} - \frac{1}{2}\text{Tr}(g_{(2)})^2 \Rightarrow \text{Tr}g^{(4)} = \frac{1}{4}\text{Tr}g_{(2)}^2$

$$\begin{cases} g^{\mu\alpha} g'^{\alpha\nu} = \frac{1}{4} (R_{\mu\alpha} R^{\alpha\nu} - \frac{1}{3} R R_{\mu\nu} + \frac{1}{36} R^2 g_{\mu\nu}^{(0)}) \\ \text{Tr}g_{(2)}^2 = \frac{1}{4} R^{\alpha\beta} R_{\alpha\beta} - \frac{1}{18} R^2 \end{cases}$$

Main point is that the vector μz equation to leading order (put $z^2=0$) gives

$$\Delta^\mu(g^{(0)}) [g_{\mu\nu}^{(2)} - \text{Tr} g_2 g_{\mu\nu}^{(0)}] = \Delta^\mu(g^{(0)}) \left[\frac{1}{2} R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu}^{(0)} \right] = 0$$

↑
put in $g_{\mu\nu}^{(2)}$
from $\mu\nu$ -eq
↑
Einstein tensor
is conserved!

but to next order in z^2 one finds that it can be written in the form

$$\Delta^\mu(g^{(0)}) \left[g_{\mu\nu}^{(4)} - \frac{1}{8} \left(R_{\mu\alpha} R_{\nu}^{\alpha} - \frac{1}{2} R R_{\mu\nu} + \left(\frac{5}{36} R^2 - \frac{1}{4} R^{\alpha\beta} R_{\alpha\beta} \right) g_{\mu\nu}^{(0)} \right) \right] = 0$$

curvature tensors of boundary metric

$$\Rightarrow g_{\mu\nu}^{(4)} = \frac{1}{8} \left(\underbrace{\quad}_{\equiv A_{\mu\nu}^{(4)}} \right) + \underbrace{t_{\mu\nu}}$$

$$\text{Tr} A^{(4)} = \frac{1}{8 \cdot 18} R^2 = \frac{1}{4} (\text{Tr} g_2)^2$$

$$\langle T_{\mu\nu} \rangle = \frac{4L^3}{16\pi G_5} (g_{\mu\nu}^{(4)} - A_{\mu\nu}^{(4)})$$

some conserved tensor $\Delta^\mu(g^{(0)}) t_{\mu\nu} = 0$
 Initial condition, not specified by Einstein!
 Energy momentum tensor!

From ZZ eqz

$$t_{\mu}^{\mu} = \text{Tr} g^{(4)} - \text{Tr} A^{(4)} = \frac{1}{4} \text{Tr} g_2^2 - \frac{1}{4} (\text{Tr} g_2)^2$$

$$= \frac{1}{4} \left[\text{Tr} g_2^2 - (\text{Tr} g_2)^2 \right] = \frac{1}{4} \left[\frac{1}{4} R^{\alpha\beta} R_{\alpha\beta} - \frac{1}{18} R^2 - \frac{1}{36} R^2 \right]$$

$$= \frac{1}{4} \left(\frac{1}{4} R^{\alpha\beta} R_{\alpha\beta} - \frac{1}{18} R^2 \right)$$

$\frac{3}{3 \cdot 18} = \frac{1}{18}$

Trace anomaly $\frac{16\pi^2}{16\pi^2} \cdot 2 \cdot \left(\frac{1}{4} R_{\alpha\beta}^2 - \frac{1}{12} R^2 \right)$ from direct computation

$$\langle T^{\mu}_{\mu} \rangle = \frac{L^3}{16\pi G_5} \left(\frac{1}{4} R^{\alpha\beta} R_{\alpha\beta} - \frac{1}{12} R^2 \right) \quad T_{\mu\nu} = \begin{pmatrix} \epsilon & p & p & p \end{pmatrix} \quad T^{\mu}_{\mu} = -\epsilon + 3p = 0$$

$\epsilon = 7p - p = T^{\mu}_{\mu} = 6p - 1 \cdot p = 5p$

AdS₅/CFT₄ encodes remarkably the trace anomaly of CFT₄ in curved space (computed with the field content of N=4 SYM & $L^3/(4\pi G_5) = N_c^2/g\pi^2$)

$g_{\mu\nu}^{(0)}(x)$ $T_{\mu\nu}(x) \leftarrow g_{\mu\nu}^{(0)}(x, 0)$ are initial conditions for integrating $g_{MN}(x, z)$ from $R_{MN} + \frac{4}{L^2} g_{MN} = 0 !!$
 For second order diff eq. you need both $f(0)$ and $f'(0)$!

Ex: $\langle T^{\mu}_{\mu} \rangle$ if bdry metric is $\begin{pmatrix} -1 & & & \\ & r^2(t) & & \\ & & r^2(t) & \\ & & & r^2(t) \end{pmatrix} = \text{FRW}$

$$g_{MN} = \frac{L^2}{z^2} \begin{pmatrix} -a(t, z) & & & 0 \\ & b(t, z) & & \\ & & b(t, z) & \\ & & & b(t, z) \\ & & & & 1 \end{pmatrix} \quad \left. \begin{matrix} a(t, z) \\ b(t, z) \end{matrix} \right\} \text{ given explicitly in th/0612226}$$

$$\Rightarrow \langle T^{\mu}_{\mu} \rangle = -\epsilon + 3T^1_1 = \frac{3N_c^2}{8\pi^2} \frac{-r^2 r''}{r^3}$$

so you have gravity dual of FRW, but AdS of course does not provide $r(t)$! Need Einstein in bdry theory

Ex. Now it is trivial to redo flat BH (p.18):

Transform it 1st to FG form

$$\frac{L^2}{z^2} \left\{ - \frac{(1 - \frac{z^4}{4z_0^4})^2}{1 + \frac{z^4}{4z_0^4}} dt^2 + (1 + \frac{z^4}{4z_0^4}) d\vec{x}^2 + dz^2 \right\} \quad \pi T = \frac{1}{z_0}$$

$$(g_{\mu\nu} + g_{\mu\nu}^{(4)} z^4) dx^\mu dx^\nu$$

$$g_{\mu\nu}^{(4)} = \begin{pmatrix} 3 & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix} \frac{1}{4z_0^4} \quad T_{\mu\nu} = \frac{N_c^2}{8\pi^2} \cdot \frac{1}{4} \pi^4 T^4 \begin{pmatrix} 3 & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix}$$

$$T_{\mu\nu} = \begin{pmatrix} \epsilon & & & \\ & p & & \\ & & p & \\ & & & p \end{pmatrix} \quad p = \frac{\pi^2 N_c^2}{8} T^4$$

Ex

$$ds^2 = -F(r) dt^2 + \frac{dr^2}{F(r)} + r^2 d\Omega^2$$

$$F(r) = \hat{r}^2 + k - \frac{M}{\hat{r}^2}$$



Application 4: Expectation values of Wilson loops

$$P \exp \left[ig \int_C A^\mu dx_\mu \right] \quad C = \text{closed loop,} \quad \text{Tr is gauge invariant}$$

Expectation value of a Wilson loop ²¹ in the boundary field theory = Action of the string hanging from the loop in the 5th dimension.

Take Q at $x = -L/2$, \bar{Q} at $L/2$. How deep does the string connecting them hang in the z direction, i.e., what is $z = z(x, t) = z(x)$ for the extremal configuration (expected to be static, no t)?

Particle action = $-m \int dt$ \Rightarrow String action = $-T \int dA$. $T = \frac{1}{2\pi\alpha'}$ = Tension.

String $X^\mu(\tau, \sigma)$ moving in a space with metric $ds^2 = g_{\mu\nu} dx^\mu dx^\nu$ has the action:

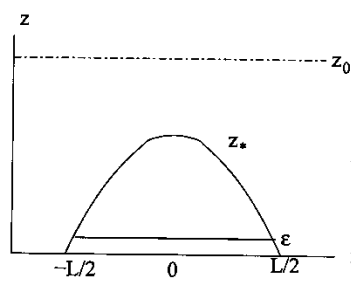
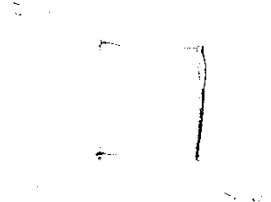
$$S = -T \int d\tau d\sigma \sqrt{-\det h_{ab}}, \quad h_{ab} = g_{\mu\nu} \frac{\partial X^\mu}{\partial \sigma^a} \frac{\partial X^\nu}{\partial \sigma^b}, \quad \sigma^a = (\tau, \sigma).$$

Nambu-Goto: $g_{\mu\nu} = \eta_{\mu\nu}$

$$h_{ab} = \begin{pmatrix} \dot{X} \cdot \dot{X} & \dot{X} \cdot X' \\ X' \cdot \dot{X} & X' \cdot X' \end{pmatrix} \quad X \cdot Y \equiv \eta_{\mu\nu} X^\mu Y^\nu \quad \cdot = \partial/\partial\tau, \quad ' = \partial/\partial\sigma.$$

²¹For a discussion of a Wilson loop and its relation to a $Q - \bar{Q}$ (free)energy at $T = 0$ and $T \neq 0$, and real $t \Leftrightarrow$ imaginary time $\tau = it$, see H.J. Rothe, Lattice gauge theories, Ch. 7 and 8 and Sect. 20.2.

String hanging ²² from a Q and a \bar{Q} at $z = 0$ in the metric $ds^2 = g_{tt}(z)dt^2 + g_{xx}(z)dx^2 + g_{zz}(z)dz^2$. 35



$$\left. \begin{aligned} g_{tt} &= -g(z) \\ g_{xx} &= 1 \\ g_{zz} &= \frac{L}{g(z)} \end{aligned} \right\} \cdot \frac{L^2}{z^2}$$

$$\sqrt{-h} = \sqrt{g(z) + z'^2 - \frac{1}{g} z'^2} \cdot \frac{L^2}{z^2}$$

$$\int dt dz \sqrt{g(z) + z'^2 - \frac{1}{g} z'^2} \cdot \frac{L^2}{z^2}$$

$$\sigma^1 = t, \quad \sigma^2 = x, \quad X^\mu = (t, x, 0, 0, z(t, x) \rightarrow z(x))$$

$$h_{ab} = \begin{pmatrix} g_{tt} + g_{zz} \dot{z}^2 & g_{zz} \dot{z} z' \\ g_{zz} \dot{z} z' & g_{xx} + g_{zz} z'^2 \end{pmatrix} \quad \sqrt{-h} = \sqrt{-g_{tt}g_{xx} - g_{tt}g_{zz}z'^2 - g_{xx}g_{zz}\dot{z}^2}$$

Extremize ($T = 1/(2\pi\alpha')$ = tension, $\dot{z} = 0$, $g(z) = 1 - z^4/z_0^4$ specialising to the 5d AdS BH)

$$S = T\Delta t \int_{-L/2}^{L/2} dx \sqrt{-h} = T\Delta t 2 \int_\epsilon^{z_*} dz x'(z) \sqrt{-h} = \Delta t \frac{L^2}{\pi\alpha'} \int_\epsilon^{z_*} \frac{dz}{z^2} \sqrt{1 + g(z)x'(z)^2}$$

Key technical point: $dx L = dx L(z(x), z'(x)) = dz x'(z) L = dz L_{\text{new}}(x'(z))$

²²Rey-Yee, hep-th/9803001; Sonnenschein, hep-th/9910089

Equation of motion is

$$\frac{\partial L}{\partial x(z)} - \frac{d}{dz} \frac{\partial L}{\partial x'(z)} = 0 \Rightarrow \frac{\partial L}{\partial x'(z)} = \frac{-g_{tt}g_{xx}x'}{\sqrt{-g_{tt}g_{zz} - g_{tt}g_{xx}x'^2}} = \text{constant.} \quad (5)$$

The constant is fixed neatly so that the maximum value of z is z_* , $z(x=0) = z_*$, $z'(z_*) = 0$. This gives $z'^2 = (dz/dx)^2$ from which by integration (Exercise)

$$L = 2 \int_{\epsilon \rightarrow 0}^{z_*} \frac{dz}{\sqrt{\frac{g_{xx}}{g_{zz}} \left(\frac{g_{tt}g_{xx}}{g_{tt}^*g_{xx}^*} - 1 \right)}} = 2 \int_0^{z_*} \frac{dz}{\sqrt{g(z) \left[\frac{g(z)}{z^4} \frac{z_*^4}{g(z_*)} - 1 \right]}} \quad g_{tt}^* \equiv g_{tt}(z_*), \text{ etc.}$$

Inserting this z'^2 to the action gives the extremal action (Exercise)

$$S = T\Delta t 2 \int_{\epsilon}^{z_*} dz \sqrt{\frac{-g_{tt}g_{zz}}{1 - \frac{g_{tt}^*g_{xx}^*}{g_{tt}g_{xx}}}} = T\Delta t 2 \int_{\epsilon}^{z_*} dz \frac{\mathcal{L}^2}{z^2 \sqrt{1 - \frac{z^4}{g(z)} \frac{g(z_*)}{z_*^4}}}$$

Separate divergence of the 5d AdS BH at $z \rightarrow 0$:

$$\int_{\epsilon}^{z_*} dz \frac{1}{z^2} f(z) = \int_{\epsilon}^{z_*} dz \frac{1}{z^2} [f(z) - 1 + 1] = \frac{1}{\epsilon} + \int_0^{z_*} dz \frac{1}{z^2} [f(z) - 1] - \frac{1}{z_*}$$

throw away $1/\epsilon$. For Euclidian Wilson loop $\langle \Delta t \times R - \text{loop} \rangle \sim \exp[-\Delta t V(R)]$ so write here $V(L, z_0) = S/\Delta t$:

37

$$V(L, z_0) = 2T\mathcal{L}^2 \left[\int_0^{z_*} \frac{dz}{z^2} \left(\frac{1}{\sqrt{1 - \frac{z^4}{g(z)} \frac{g(z_*)}{z_*^4}}} - 1 \right) - \frac{1}{z_*} \right]$$

Scaling $z = yz_0$, $z_* = y_m z_0$:²³

$$L = 2z_0 \int_0^{y_m} \frac{dy}{\sqrt{(1-y^4)[q(y_m)/q(y) - 1]}} \quad q(y) \equiv y^4/(1-y^4) \quad (6)$$

$$V(L, z_0) = \frac{\mathcal{L}^2}{\alpha' \pi z_0} \left\{ \int_0^{y_m} \frac{dy}{y^2} \left[\frac{1}{\sqrt{1 - q(y)/q(y_m)}} - 1 \right] - \frac{1}{y_m} \right\} \quad (7)$$

Units of length and energy given by

$$\pi z_0 = \frac{1}{T_H}, \quad -V_{Q\bar{Q}} \equiv \frac{\mathcal{L}^2}{\alpha' \pi z_0} = \sqrt{g^2 N_c} T_H.$$

Small L (and $y_m \rightarrow 0$, $q(y) \approx y^4$):

$$V(L) = -\frac{0.2285 \sqrt{g^2 N_c}}{L}$$

Intermediate $0 < L < L_{\max}$: can fit to

$$V(L) = -\frac{4}{3} \frac{\alpha_s}{L} + \sigma L$$

Small L again for $y_m \rightarrow 1$:

$$V(L, z_0) \rightarrow V_{Q\bar{Q}} = \frac{\mathcal{L}^2}{\pi \alpha'} \int_{\epsilon}^{z_0} \frac{dz}{z^2} \Rightarrow -\frac{\mathcal{L}^2}{\alpha' \pi z_0}$$

After some y_m , for $L >$ some value (see Fig) the dominant config is that with separate $Q\bar{Q}$. This is also a solution of (5) with $x' = 0$, $z_* = z_0$.

²³Plotted with Mathematica using ParametricPlot[{L[y_m], V[y_m]}, {y_m, 0.15, 0.994}] in Fig.1

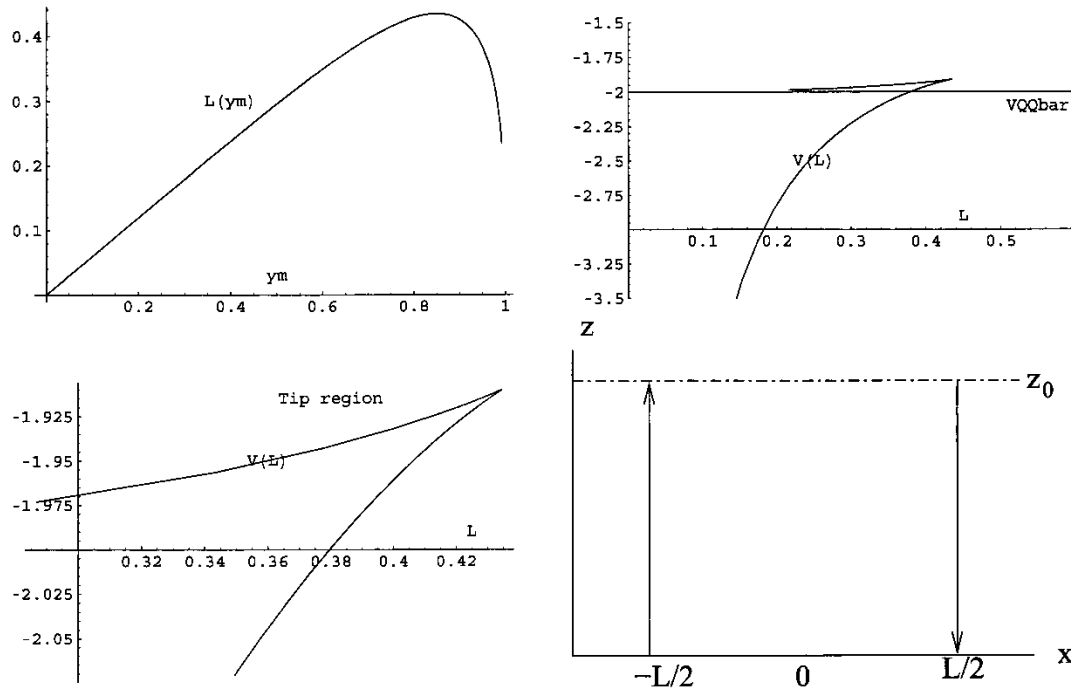


Figure 1: The distance L as function of y_m evaluated from (6), the extremal action and its tip region from (7) (all scaled by the factors outside the integral) and the $Q\bar{Q}$ configuration.

Picture from AdS/CFT:

- at small distances conformally invariant form $V \sim 1/L$, $\sqrt{\lambda}$ dependence on $g^2 N_c$ is typical of strong coupling.
- at some distance interaction is screened and $Q\bar{Q}$ separate.
- one can put in numbers ²⁴
- mathematics of the curves is pretty: the independent $Q\bar{Q}$ solution is obtained also from the string-connected solution when $z_* \rightarrow z_0$ and the two branches approach each other

This was just an example of numerous applications of AdS/CFT to Wilson loop computations. Works even for gluonic scattering amplitudes!! ²⁵

²⁴Large number of papers, mine is Kajantie-Tahkokallio-Yee, hep-ph/0609254

²⁵Alday-Maldacena, arXiv:0705.0303