Phases and phase transitions of QCD with large N_f $N_c=3>>1$ =0,3,6,9,...

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QCD thermodynamics

$$e^{p(T, \mu; m_q)} \frac{V}{T} =$$
 Maximum p dominates!

$$\int \mathcal{D}A\mathcal{D}q \ e^{-\int^{1/T} d\tau d^3x \left[\frac{1}{g^2}F^2 + \bar{q}(\partial + A)q + \frac{m_q}{\bar{q}q} + \mu q^{\dagger}q\right]}$$

Color N_c, Flavor N_f, QCD scale Λ_{QCD}

 $m_q = 0$ to have chiral symmetry $\mu = 0$ in this talk

$$QCD_{physical}$$
: m_u , m_d , m_s , m_c , m_b , m_t

Chiral symmetry:

$$\text{Action } \int d^4x \big[\frac{1}{g^2} F^2 + \bar{q} (\partial + A) q \big]$$

has $\operatorname{SU}_L(N_f) \times \operatorname{SU}_R(N_f)$ symmetry, while solutions may only have $\operatorname{SU}_V(N_f)$ symmetry $N_f^2 - 1 \approx N_f^2$ massless Goldstone bosons, "pions"

(in SM EW sector $N_f=2$ and 3 gb's are eaten by W and Z)

Order parameter: $\langle \bar{q}q \rangle$

Chiral phase transition at some $T_c \approx \Lambda_{\text{QCD}} \approx m_{\text{min}}/5$ $m_{\pi} = 0!$

No order parameter for confinement-deconfinement, unless N_f=0

Values of N_f:

$$\mu \frac{\partial g(\mu)}{\partial \mu} = -\frac{1}{16\pi^2} \left(\frac{11}{3} N_c - \frac{2}{3} N_f\right) g^3 + \dots$$

As freedom lost if $x_f \equiv \frac{N_f}{N_c} > \frac{11}{2}$

Massless dofs at T >>T
$$_{
m c}$$
 $2N_c^2+rac{7}{2}\,N_cN_f$ Equ

Equal at
$$x_f = 4$$

Conformal domain: $x_c \approx 4 < x_f < 5.5$



Universality + ε expansion suggest 1st order!

We want p(T), all T

$$\begin{split} s(T) &= p'(T) \qquad \epsilon(T) = Ts(T) - p(T) \\ & T \frac{d}{dT} \frac{p(T)}{T^4} = \frac{\epsilon - 3p}{T^4} \qquad \text{interaction measure} \end{split}$$

$$\left(T\frac{d}{dT}\right)^2 \frac{p(T)}{T^4} = \frac{(c_s^{-2} - 3)(\epsilon + p) - 4(\epsilon - 3p)}{T^4}$$

sound velocity

Massless particles,
$$\epsilon = 3p \sim T^4, \quad c_s^2 = rac{1}{3}$$

Lattice data for interaction measure:



Here computed from one theory: QCD

How do you connect if you have two different appros of QCD?

Interaction measure in cosmology:

$$R = g^{\mu\nu}R_{\mu\nu} = -8\pi G T^{\mu}_{\mu} = 8\pi G (\epsilon - 3p) = 8\pi G (\epsilon_{\rm QCD} - 3p_{\rm QCD} + \epsilon_{\rm DM} + 4\epsilon_{\Lambda})$$

At T_{QCD} the QCD sector dominates over DM the curvature of the Universe by the factor

$$\begin{split} \frac{T_{\rm QCD}}{T_0} \cdot \frac{\rho_{\rm rad,0}}{\rho_c} &\approx 10^{12} \cdot 10^{-5} \\ R &\sim \frac{1}{t^2} \frac{\epsilon_{\rm QCD} - 3p_{\rm QCD}}{T^4} \qquad {\rm t} \sim 10 {\rm km} \end{split}$$

Similarly for EW etc

Hadron gas

1.5

1.0

0.5

2

m=0:

 $x^2 K_2(\mathbf{x})$

6

8

4



over all hadrons in Particle Data Tables

or integrate over the Hagedorn spectrum

over all hadrons in Particle Data Tables
$$m=0:$$

 $p_{gb} = N_f^2 \frac{\pi^2}{90} T^4$
 $= N_c^2 T^4 \cdot x_f^2 \frac{\pi^2}{90}$
 $\rho(m, b, a, m_0) = \delta(m) + \frac{\rho_0}{m_0} \left(\frac{m}{m_0}\right)^a e^{bm} \theta(m-m_0)$
 $\int_{m_0}^{\infty} dm \exp[bm - \frac{m}{T}] m^a \dots \longrightarrow T \leq T_H \equiv \frac{1}{b}$

m_{min} from holography: $m_0 \equiv m_{\min} = 0.707\Lambda, \quad x_f = 2$ Plasma phase from holography, gauge/gravity duality " $\pi=3$ "

$$S = \frac{1}{16\pi G_5} \int d^5x \sqrt{-g} \left[R - \frac{4}{3} (\partial_z \phi)^2 + V_g(\phi) - V_f(\phi, \tau) \sqrt{1 + (\partial_z \tau)^2} \right]$$
Järvinen-Kiritsis 1112.1261

$$ds^{2} = b^{2}(z) \left[-f(z)dt^{2} + d\mathbf{x}^{2} + \frac{dz^{2}}{f(z)} \right] \quad \lambda(z) \quad \tau(z)$$

$$f(z_h) = 0, \quad -f'(z_h) = 4\pi T, \quad s = \frac{A}{4G_5} = \frac{b^{\circ}(z_h)}{4G_5}$$

p_q(T) from classical gravity black hole solutions:"tree level"
 No classical solns for p_h(T): arises from "1loop or stringy effects"!
 Mass spectrum from fluctuations around f(z)=1, T=0 solutions 1loop computation!

1st order transition $p_h(T_c)=p_q(T_c)$ is trivial: draw line at gb



2nd order transition $p_h(T_c)=p_q(T_c)$, $p_h'(T_c)=p_q'(T_c)$ is simple:



A cusp in interaction measure

Cusp: 2nd derivatives of opposite sign!

Must be equal for 3rd order!

We found that bending the 2nd derivative of HG down to negative side required including repulsive interactions between hadrons arising from their finite size!

Include in the textbook derivation of BE distribution

$$V \to V - v_0 N, \quad v_0 \equiv \frac{1}{T_0^3}$$

$$\frac{p_{\rm h}}{N_c^2 T^4} = \frac{\pi^2}{90} x_f^2 + \frac{\rho_0}{m_0} x_f^2 \int_{m_0}^{\infty} dm \, \frac{m^a}{m_0^a} \, e^{bm} \, \frac{T_0^3}{T^3} \, W\!\left(\frac{m^2 T}{2\pi^2 T_0^3} K_2\left(\frac{m}{T}\right)\right)$$



3rd order transition $p_h(T_c)=p_q(T_c)$, $p_h'(T_c)=p_q'(T_c)$, $p_h''(T_c)=p_q''(T_c)$:







$x_f = N_f / N_c$ dependence

scales with m_{min}:
$$T_c = (0.27 \pm 0.03) m_{min}$$

 $m_0 = (2.8, 0.71, 0.08)$ $x_f = (1, 2, 2.5)$



Another 3rd order PT, Tracy-Widom universality class, Gross-Witten-Wadia:



For N=infty a 3rd order PT between left and right! GWW1980 Goldstone bosons at T=0 disappear with finite m_q :



Concluding questions:

- What is the chiral transition really: 1st, 2nd, 3rd,..., continuous?

- Can lattice MC in practice give the answer: $V \to \infty$, $a \to 0$, $m_q \to 0$?

- What do you get when μ is included? Theoretical problem: what are the potentials in the action? Practical problem: how to organise the very demanding computation

Overflow



Hadron	Plasma	Energy
gas		Unit Λ



$$c_{\rm SB} + c_2 g^2 + c_3 g^3 + (c_4' \log g + c_4) g^4 + c_5 g^5 + (c_6' \log g + c_6) g^6 + c_7 g^7 + \dots$$

$$c_{2} \text{ Shuryak 78, } c_{3} \text{ Kapusta 79, } c_{4}' \text{ Toimela 83, } c_{4} \text{ Arnold-Zhai 94,} \\ c_{5} \text{ Zhai-Kastening, Braaten-Nieto 95, } c_{6}' \text{ Kajantie-Laine-Rummukainen-Schröder 03} \\ p/p_{\text{SB}} = \\ 1 - \frac{5}{2} \lambda(\bar{\mu}) + \frac{20}{\sqrt{3}} \lambda^{3/2} + \left[30 \log(\frac{2}{3}\lambda) + p_{2}b_{0} \log \frac{\bar{\mu}}{4\pi T} + 99.0784 \right] \lambda^{2} \\ + \left[\frac{3}{2} p_{3}b_{0} \log \frac{\bar{\mu}}{4\pi T} - 227.746 \right] \lambda^{5/2} + \\ \left\{ \left(-42.8187 + 60b_{0} \log \frac{\bar{\mu}}{4\pi T} \right) \log(\frac{2}{3}\lambda) - 140.915 \log \lambda + p_{2}b_{0}^{2} \log^{2} \frac{\bar{\mu}}{4\pi T} \\ + \left(p_{2}b_{1} + p_{4}b_{0} + 2b_{0} 99.0784 \right) \log \frac{\bar{\mu}}{4\pi T} + q_{0} \right\} \lambda^{3} + \mathcal{O}(\lambda^{7/2}) \\ \frac{1}{\lambda(\mu)} = b_{0} \log \frac{\mu}{\Lambda} + \frac{b_{1}}{b_{0}} \log(\log \frac{\mu}{\Lambda}) \end{aligned}$$

