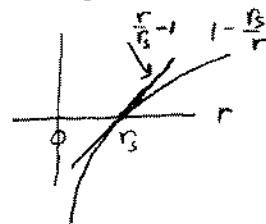


Getting  $T_H$  from the BH metric: seems like a very formal argument!

Finite  $T$  means periodicity in  $\tau = it$ :  $f(\tau + \frac{h}{T}) = f(\tau)$

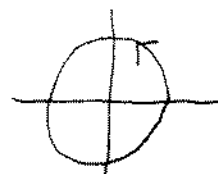
Near  $r_s$ :  $1 - \frac{2MG}{r} \approx \frac{1}{r_s}(r - r_s)$        $r_s = 2MG = 2 \frac{M}{M_{Pl}^2}$



$$ds^2 = \frac{1}{r_s}(r - r_s) d\tau^2 + \frac{r_s}{r - r_s} dr^2 + \dots$$

Try:  $g = 2\sqrt{r_s(r - r_s)}$        $dg = \sqrt{\frac{r_s}{r - r_s}} dr$        $\frac{1}{4}g^2 = r_s(r - r_s)$

$$ds^2 = dg^2 + \frac{1}{4} \frac{g^2}{r_s^2} d\tau^2 = dg^2 + g^2 d\left(\frac{\tau}{2r_s}\right)^2$$



compare with

$$\begin{aligned} x &= g \sin \varphi & dx^2 + dy^2 &= (g \cos \varphi d\varphi - \varphi g \sin \varphi dr)^2 + (g \sin \varphi d\varphi + \varphi g \cos \varphi dr)^2 \\ y &= g \cos \varphi & &= dg^2 + g^2 d\varphi^2 \end{aligned}$$

Periodicity  $\varphi + 2\pi$

$$\frac{T}{2r_s} + 2\pi \quad \text{or} \quad \tau + 4\pi r_s$$

$$\Rightarrow 4\pi r_s = \frac{1}{T_H}$$

$$T_H = \frac{1}{4\pi r_s} = \frac{1}{8\pi G M} \equiv \frac{M_{Pl}^2}{8\pi M}$$



BB radiation:

$$\frac{dM}{dt} \sim -\frac{1}{r_s^2} T_H^4 \sim -\frac{1}{r_s^2} \sim -\frac{M_{Pl}^4}{M^2}$$

$$M^2 dM = -M_{Pl}^4 dt \quad \int_{M_0}^{M/3} \frac{1}{M^3} dM = -M_{Pl}^4 t \Rightarrow \frac{1}{3} M^{3/4} = \frac{1}{3} M_0^3 - M_{Pl}^4 t$$

$$M^{3/4} = M_0^3 - 3M_{Pl}^4 t$$

$$\Rightarrow \text{lifetime of BH} \sim \frac{M_0^3}{M_{Pl}^4} \sim \left(\frac{M_0}{M_{Pl}}\right)^3 \frac{1}{M_{Pl}}$$

Example BH with lifetime  $\sim \tau_{universe}$ :

$$\frac{1}{H_0} \sim \tau_U \sim \frac{3}{h} 10^{10} \frac{1}{M_{Pl}}$$

$$M = 10^{20} M_{Pl} \Rightarrow \text{lifetime} = 10^{10} \text{ a} \sim \tau_U$$

$$\sim 10^{12} \text{ kg} \quad r_s \sim 10^{20} \frac{1}{M_{Pl}} \sim \frac{1}{0.1 \text{ GeV}} \approx 1 \text{ fm}$$

BH history:



Entropy:

$$S_{BH} = \frac{A_d}{4\hbar G_d} \Rightarrow \text{holography}$$

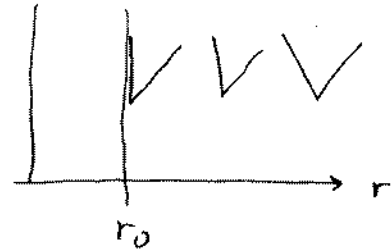
$$\dim G_d = (\text{length})^{d-2} = \text{area}$$

More generally:

$$-dt^2 = -g(r) dt^2 + \frac{1}{f(r)} dr^2 \quad \equiv \quad g_{tt} dt^2 - g_{rr} dr^2$$

$$\underset{r=r_0}{\approx} -g'(r_0)(r-r_0) dt^2 + \frac{1}{f'(r_0)(r-r_0)} dr^2$$

light:  $\frac{dt}{dr} = \frac{1}{\sqrt{g'(r_0)f'(r_0)}} \frac{1}{r-r_0}$



(1) go Euclidian,  $\tau = it$

(g) transform to

$$dg^2 + g^2 d\varphi^2 \quad \varphi \rightarrow \varphi + 2\pi - \text{periodic!}$$

$$ds^2 = \frac{1}{f'_0(r-r_0)} dr^2 + g'_0(r-r_0) dt^2$$

$$g = 2\sqrt{\frac{1}{f'_0}(r-r_0)} \quad g^2 = \frac{4}{f'_0}(r-r_0) \quad 2g dg = \frac{4}{f'_0} dr$$

$$dr = \frac{f'_0}{4} \cdot 4\sqrt{\frac{1}{f'_0}(r-r_0)} dg \quad dr^2 = f'_0(r-r_0) dg^2$$

$$\Rightarrow ds^2 = dg^2 + g_0 \cdot \frac{1}{4} g^2 f'_0 dt^2 = dg^2 + g^2 \cdot \frac{g'_0 f'_0}{4} dt^2$$

periodic under  $\tau \rightarrow \tau + \frac{4\pi}{\sqrt{g'_0 f'_0}}$

$$T_H = \frac{1}{4\pi} \sqrt{g'_0(r_0) f'_0(r_0)}$$



Usual BH:  $g'_0 = \frac{1}{r_s} = f'_0 \quad T_H = \frac{1}{4\pi r_0}$

If the period is not  $2\pi$ , there is a conical singularity; singularity is removed only for this period

Two other cases:

(1) de Sitter

in "static coordinates"  
see later!

$$ds^2 = -(1 - H^2 r^2) dt^2 + \frac{1}{1 - H^2 r^2} dr^2 - \dots$$

$$r_0 = \frac{1}{H} \quad g'(r_0) = -2H = f'(r_0)$$

$$\Rightarrow \boxed{T_H = \frac{H}{2\pi}} \quad (R/H \sim e^{Ht})$$

(2) BH in  $AdS_{m+1}$

$$ds^2 = -\left(1 - \frac{\omega_m G_{m+1} M}{r^{m-2}} + \frac{r^2}{b^2}\right) dt^2 + \frac{1}{(\text{same})} dr^2 + r^2 d\Omega_{m-1}^2$$

↑ anti-dS  $b = \frac{1}{H}$

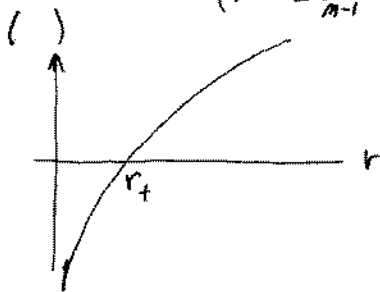
$$m \Rightarrow d-1 \quad G_d = \frac{1}{M^{d-2}}$$

$$\omega_m = \frac{16\pi}{(m-1)\Omega_{m-1}}$$

$$\omega_3 = \frac{16\pi}{2 \cdot 4\pi} = 2$$

$$\dim G_{m+1} = \frac{1}{M^{m-1}}$$

$$[G_{m+1} M] = \frac{1}{M^{m-2}} \equiv r_s^{m-2}$$



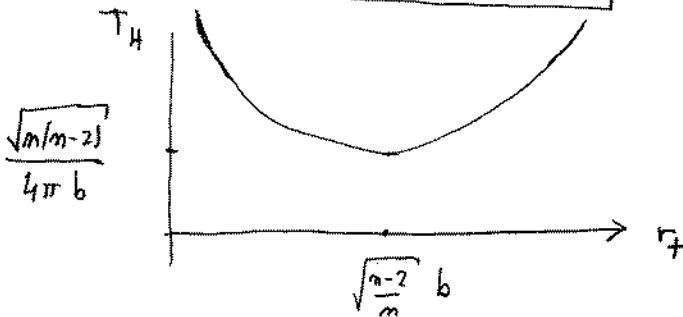
$$g'(r_+) = \frac{2r_+}{b^2} + \frac{(m-2)\omega_m G_{m+1} M}{r_+^{m-1}} = \frac{2r_+}{b^2} + \frac{m-2}{r_+} \left(1 + \frac{r_+^2}{b^2}\right)$$

$$= \frac{m-2}{r_+} + \frac{m r_+}{b^2}$$

$$\frac{\omega_m G_{m+1} M}{r_+^{m-1}} \cdot r_+ = 1 + \frac{r_+^2}{b^2}$$

$$\Rightarrow \boxed{T_H = \frac{m r_+^2 + (m-2) b^2}{4\pi r_+ b^2}}$$

This metric will be related to the 3-brane metric on p. 15



$$\boxed{S = \frac{\Omega_{m-1} r_+^{m-1}}{4G_{m+1}} = \frac{\text{"area of horizon"}}{4G_{m+1}}}$$

$AdS_{m+1} \quad d \Leftrightarrow m+1$   
 $4 = 3+1$