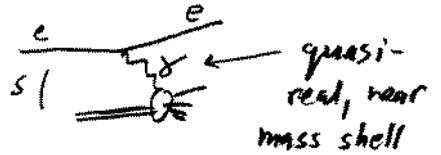
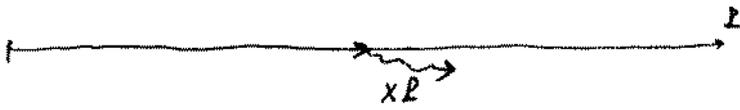


Collinear factorisation of gluonic MHV amps

- Weizsäcker-Williams photons



$$f_{\gamma/P}(x) = \frac{\alpha}{2\pi} \log \frac{s}{4m^2} \frac{1+(1-x)^2}{x}$$

- AP splitting functions:

$$f_{g/g}(x) = \frac{\alpha_s}{g\pi} \log Q^2 P_{gg}(x)$$



$$x = \frac{P^+}{P^+}$$

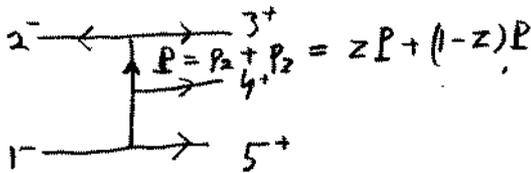
$$6 \left[\frac{1-x}{x} + \frac{x}{1-x} + x(1-x) + \dots \right]$$

evol.

$$\begin{aligned} \frac{dg(x, Q^2)}{d \log Q^2} &= \frac{\alpha_s(Q^2)}{2\pi} \int_x^1 \frac{dz}{z} P_{gg}(z) g\left(\frac{x}{z}, Q^2\right) + \dots \\ &= \int_0^1 \int_0^1 dy dz P_{gg}(z) g(y, Q^2) \delta(x - zy) \end{aligned}$$

How is this seen in MHV amps? Take

$$A_5(1^- 2^- 3^+ 4^+ 5^+) = ig^3 \frac{\langle 12 \rangle^4}{\langle 12 \rangle \langle 23 \rangle \langle 34 \rangle \langle 45 \rangle \langle 51 \rangle} \text{ when } \langle 23 \rangle \rightarrow 0:$$



$$\langle 34 \rangle = \lambda_3^a \lambda_{4a}$$

$$P_{3a\dot{a}} = \lambda_{3a} \tilde{\lambda}_{3\dot{a}} \quad \tilde{\lambda} = \lambda^*$$

$$= (1-z) P_{a\dot{a}} = \lambda_{2a} \tilde{\lambda}_{2\dot{a}}$$

$$\Rightarrow \langle 34 \rangle = \sqrt{1-z} \langle P4 \rangle$$

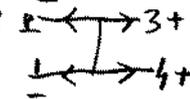
$$\langle 12 \rangle = \sqrt{z} \langle 1P \rangle$$

$$\approx \frac{1}{\langle 23 \rangle} g \sqrt{\frac{z^4}{z(1-z)}} ig^2 \frac{\langle 1P \rangle^4}{\langle 1P \rangle \langle P4 \rangle \langle 45 \rangle \langle 51 \rangle}$$

$$P^2 = g^2 p_2 \cdot p_3 = \langle 23 \rangle [23] \rightarrow 0$$

splitting fn

quasireal



$$= (\text{polarised A-P})^{\frac{1}{2}}$$

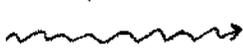
Gluons → Gravitons; weak field gravity

Berends-Giele-Kuijf PLB 911 (88) 91 Kawai-Lewellen-Tye NPB 269 (86) 1

Nair hep-th/0501143

Cachazo 0509160
-Svrcek

$$P_{a\dot{a}} = \lambda_{\alpha} \tilde{\lambda}_{\dot{\alpha}}$$



$$A(\xi) = \frac{i}{M_{Pl}^2} T^{\mu\nu} \frac{i P_{\mu\nu\alpha\beta}}{q^2 + i\epsilon} T^{\alpha\beta}$$

$$P_{\mu\nu\alpha\beta} = \frac{1}{2} [\gamma_{\mu\alpha} \gamma_{\nu\beta} + \gamma_{\mu\beta} \gamma_{\nu\alpha} - \gamma_{\mu\nu} \gamma_{\alpha\beta}]$$

$q^{\mu} T_{\mu\nu} = 0$ harmonic gauge

B. DeWitt PR 162 (1967)

$$\xi_{\mu\nu}^{(+)} \rightarrow \xi_{a\dot{a}b\dot{b}}^{(+)} = \xi_{a\dot{a}}^{(+)} \xi_{b\dot{b}}^{(+)}$$

Donoghue gr-qc/9512024
Bjerrum-Bohr hep-th/0410097

$$\frac{M_a \tilde{\lambda}_a}{[\lambda\mu]}$$

Free gravitons

Linearize,

$$S = \frac{1}{2\kappa_d^2} \int d^d x \sqrt{g} R = \frac{1}{2} \int d^d x h^{\mu\nu} (R_{\mu\nu} - \frac{1}{2} \gamma_{\mu\nu} \gamma^{\alpha\beta} R_{\alpha\beta})_{Linear}$$

$2\kappa_d^2 = 16\pi G_d$ $d=4$ $g_{\mu\nu} = \gamma_{\mu\nu} + \kappa h_{\mu\nu}$ $g^{\mu\nu} = \gamma^{\mu\nu} - \kappa h^{\mu\nu} + \kappa^2 h^{\mu\alpha} h_{\alpha}^{\nu} + \dots$

$$\kappa_4 = \sqrt{8\pi G_N}$$

$\gamma^{\mu\nu} \gamma_{\mu\nu} = \gamma^{\mu}_{\mu}$

$$R_{\mu\nu}^{Lin} = [-\partial^2 \gamma_{\mu\alpha} \gamma_{\nu\beta} + \partial_{\mu} \partial_{\alpha} \gamma_{\nu\beta} + \partial_{\nu} \partial_{\alpha} \gamma_{\mu\beta} - \partial_{\mu} \partial_{\nu} \gamma_{\alpha\beta}] h^{\alpha\beta}$$

$$R^{\mu\nu L} = (-\partial^2 \gamma^{\mu\alpha} \gamma^{\nu\beta} + \partial^{\mu} \partial_{\alpha} \gamma^{\nu\beta} - \partial^{\nu} \partial_{\alpha} \gamma^{\mu\beta}) h^{\alpha\beta} = \partial^{\mu} \partial_{\alpha} \gamma^{\nu\beta} - \partial^{\nu} \partial_{\alpha} \gamma^{\mu\beta} h^{\alpha\beta}$$

$$S_{Lin} = \int d^d x h^{\mu\nu} [\partial^2 \gamma_{\mu\alpha} \gamma_{\nu\beta} + \partial_{\mu} \partial_{\alpha} \gamma_{\nu\beta} + \partial_{\nu} \partial_{\alpha} \gamma_{\mu\beta} - \partial_{\mu} \partial_{\nu} \gamma_{\alpha\beta} + \gamma_{\mu\nu} (\partial^2 \gamma_{\alpha\beta} - \partial_{\alpha} \partial_{\beta})] h^{\alpha\beta}$$

$D_{\mu\nu, \alpha\beta}^{(0)-1}$ inverse graviton propagator

Invariant under

$$h_{\mu\nu} \rightarrow h_{\mu\nu} + \partial_{\mu} \epsilon_{\nu}(x) + \partial_{\nu} \epsilon_{\mu}(x) \quad (A_{\mu} \rightarrow A_{\mu} + \partial_{\mu} \epsilon)$$

Reminder:

$$R_{\mu\nu\lambda\sigma} = \frac{1}{2} (\partial_{\sigma} \partial_{\lambda} g_{\mu\nu} - \partial_{\sigma} \partial_{\mu} g_{\nu\lambda} - \partial_{\lambda} \partial_{\nu} g_{\mu\sigma} + \partial_{\lambda} \partial_{\mu} g_{\nu\sigma}) + g_{\alpha\beta} (\Gamma_{\lambda\mu}^{\alpha} \Gamma_{\nu\sigma}^{\beta} - \Gamma_{\sigma\mu}^{\alpha} \Gamma_{\nu\lambda}^{\beta})$$

$$\nabla_{\nu} V_{\mu} = \partial_{\nu} V_{\mu} - \Gamma_{\mu\nu}^{\alpha} V_{\alpha} \quad (D_{\mu} = \partial_{\mu} + i g A_{\mu})$$

$$(\nabla_{\lambda} \nabla_{\nu} - \nabla_{\nu} \nabla_{\lambda}) V_{\mu} = -V_{\alpha} R^{\alpha}_{\mu\nu\lambda} \quad ([D_{\mu}, D_{\nu}] = \frac{i}{4} F_{\mu\nu})$$

$$\Gamma^{\sigma}_{\mu\nu} = \frac{1}{2} g^{\alpha\sigma} (\partial_{\mu} g_{\nu\alpha} + \partial_{\nu} g_{\mu\alpha} - \partial_{\alpha} g_{\mu\nu})$$

Contractions

Unique

scalar!

Field eq.

$$R_{\mu\nu} \lambda^{\mu} \lambda^{\nu} = 0 \quad \epsilon^{\mu\nu\lambda\sigma} R_{\mu\nu\lambda\sigma} = 0$$

$$R_{\nu\beta} = R_{\mu\nu\lambda\sigma} \lambda^{\mu} \lambda^{\lambda} \lambda^{\sigma} \quad R = R_{\mu\nu\lambda\sigma} \lambda^{\mu} \lambda^{\nu} \lambda^{\lambda} \lambda^{\sigma} = -R_{\mu\nu\lambda\sigma} \lambda^{\mu} \lambda^{\nu} \lambda^{\lambda} \lambda^{\sigma}$$

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu} \quad \text{or} \quad R_{\mu\nu} = 8\pi G (T_{\mu\nu} - \frac{1}{2} g_{\mu\nu} T^{\alpha}_{\alpha})$$

Due to inv. under $h_{\mu\nu} \rightarrow h_{\mu\nu} + \partial_\mu \epsilon_\nu + \partial_\nu \epsilon_\mu$ lots of extra dofs
 6 off diag
 + 4 diag comps

$$\partial_\mu F_{\mu\nu} = \partial_\mu (\partial_\mu A_\nu - \partial_\nu A_\mu) = \partial^2 A_\nu - \partial_\nu \partial^\mu A_\mu = 0$$

if $\partial_\mu A^\mu = 0$

Gauge fixing: $\rightarrow \partial_\mu \partial_\alpha h^{\alpha\mu} = \frac{1}{2} \partial^2 h^\alpha{}_\alpha \Rightarrow \partial^2 A_\nu = 0$

4 relations $\partial_\alpha h^\alpha{}_\mu - \frac{1}{2} \partial_\mu h^\alpha{}_\alpha = 0$

$$-\partial^2 h_{\mu\nu} + \underbrace{\partial_\mu \partial_\alpha h^\alpha{}_\nu + \partial_\nu \partial_\alpha h^\alpha{}_\mu - \partial_\mu \partial_\nu h^\alpha{}_\alpha}_{\text{removes these terms}} + \eta_{\mu\nu} (\partial^2 h^\alpha{}_\alpha - \underbrace{\partial_\alpha \partial^\alpha h^\mu{}_\mu}_{= \frac{1}{2} \partial^2 h^\alpha{}_\alpha})$$

$$\Rightarrow \boxed{\partial^2 h_{\mu\nu} - \frac{1}{2} \eta_{\mu\nu} \partial^2 h^\alpha{}_\alpha = 16\pi G T_{\mu\nu}} + \frac{1}{2} \eta_{\mu\nu} \partial^2 h^\alpha{}_\alpha$$

$$\eta^{\mu\nu} : \partial^2 h_{\mu\nu} - \frac{1}{2} \eta^{\mu\nu} \partial^2 h = -\partial^2 h$$

$$\tilde{h}_{\mu\nu} \equiv h_{\mu\nu} - \frac{1}{2} \eta_{\mu\nu} h \quad \partial^2 \tilde{h}_{\mu\nu} = 16\pi G T_{\mu\nu}$$

$$\tilde{h} \equiv \tilde{h}^\mu{}_\mu = h - \frac{1}{2} \cdot 4h = -h$$

Plane waves

$$h_{\mu\nu}(x) = \text{Re } \epsilon_{\mu\nu} e^{-i p \cdot x}$$

$$T_{\mu\nu} = 0 \quad \partial^2 h = 0$$

$$\partial^2 h_{\mu\nu} = 0 \Rightarrow p^2 = 0$$

gauge fixing $p_\alpha \epsilon^\alpha{}_\mu = \frac{1}{2} p_\mu \epsilon^\alpha{}_\alpha$ 6 dofs

inv. under $\epsilon^\mu{}_\nu \rightarrow \epsilon^\mu{}_\nu + \gamma^\mu{}_\rho \epsilon^\rho{}_\nu + \gamma^\nu{}_\rho \epsilon^\mu{}_\rho$

$$\begin{aligned} A_\mu &= \text{Re } \epsilon_\mu e^{-i p \cdot x} && 10 \text{ dofs} \\ \square A_\mu &= 0 \Rightarrow p^2 = 0 && 4 \text{ dofs} \\ \partial^\mu A_\mu &= 0 && p \cdot \epsilon = 0 \quad 3 \text{ dofs} \\ A_\mu &\rightarrow A_\mu + \partial_\mu \epsilon && \epsilon^\mu \rightarrow \epsilon^\mu + \gamma^\mu{}_\rho \epsilon^\rho && 2 \text{ dofs} \end{aligned}$$

Choose $p^\mu = (E, 0, 0, E)$ and work out explicitly what the above conditions imply for $\epsilon_{\mu\nu}$ (Weinberg sect 10.2). See that by proper choice of η_μ can choose

$$\epsilon_{\mu\nu} = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & \epsilon_{11} & \epsilon_{12} & 0 \\ 0 & \epsilon_{12} & -\epsilon_{11} & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix}$$

Now rotate around z axis ^{by angle θ} , leaving p^μ inv. Find

$$\epsilon_{11} \mp i\epsilon_{12} \rightarrow e^{\pm 2i\theta} (\epsilon_{11} \mp i\epsilon_{12}) \quad !!$$

↑
helicity 2

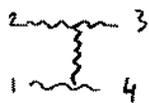
We have redone the little group analysis of reps of Poincaré!

⇒ gravitational waves

[Here pp. 33-36]

Back to MHV ^{of} gravitons Higher orders give couplings

and one can evaluate diags like



nonrenormalisable! coupling const
 $\sqrt{G} \sim \frac{1}{M_{Pl}}$

But tree diags can be evaluated!

$$A_4(1^- 2^- 3^+ 4^+) = 8\pi G \frac{\langle 12 \rangle^8 [12]}{\langle 12 \rangle \langle 13 \rangle \langle 14 \rangle \langle 23 \rangle \langle 24 \rangle \langle 34 \rangle \langle 44 \rangle} \leftarrow \text{note both } \langle \rangle \text{ and } []$$

$\equiv N(4)$

$\dim A_4 = \frac{1}{M_{Pl}^2} \dim \langle 12 \rangle^2 = 1$

$$A_5(1^- 2^- 3^+ 4^+ 5^+) = -4i (8\pi G)^{3/2} \frac{\langle 12 \rangle^8}{N(5)} \frac{\epsilon_{\mu\nu\alpha\beta} p_1^\mu p_2^\nu p_3^\alpha p_4^\beta}{4i [[12] \langle 23 \rangle [34] \langle 41 \rangle - h.c.]}$$

$$\dim A_5 = \frac{1}{M_{Pl}^3} \frac{\langle 12 \rangle^8}{\langle 12 \rangle^9} p^4 = \frac{p}{\langle 12 \rangle} = \frac{1}{p} \quad \dim A_n = GeV^{4-n}$$

Much is happening here, too!

Work out $\epsilon_{\mu\nu}$ for $p^\mu = (E \ 0 \ 0 \ E)$:

Conditions: $p^\alpha \epsilon_{\alpha\mu} = \frac{1}{2} p_\mu \epsilon^\alpha{}_\alpha$

$$\Rightarrow E(\epsilon_{0\mu} + \epsilon_{3\mu}) = \frac{1}{2} p_\mu (\epsilon_{00} - \epsilon_{11} - \epsilon_{22} - \epsilon_{33})$$

$$\mu=0 \quad \epsilon_{00} + \epsilon_{30} = \frac{1}{2}(\epsilon_{00} - \epsilon_{11})$$

$$\mu=1,2 \quad \epsilon_{01} + \epsilon_{31} = \epsilon_{02} + \epsilon_{32} = 0$$

$$\mu=3 \quad \epsilon_{03} + \epsilon_{33} = -\frac{1}{2}(\epsilon_{00} - \epsilon_{11})$$

$$\epsilon_{00} + \epsilon_{03} = -\epsilon_{03} - \epsilon_{33}$$

↓

$$\epsilon_{00} + \epsilon_{33} = -2\epsilon_{03}$$

1st eq: $\epsilon_{00} - \frac{1}{2}(\epsilon_{00} + \epsilon_{33}) = \frac{1}{2}(\epsilon_{00} - \epsilon_{33} - \epsilon_{11} - \epsilon_{22})$

$$\Rightarrow \epsilon_{11} + \epsilon_{22} = 0$$

4 eqs are, e.g.
$$\begin{cases} \epsilon_{01} = -\epsilon_{31} & \epsilon_{02} = -\epsilon_{32} & \epsilon_{03} = -\frac{1}{2}(\epsilon_{00} + \epsilon_{33}) \\ \epsilon_{22} = -\epsilon_{11} \end{cases}$$

But now do $\epsilon_{\mu\nu} \rightarrow \epsilon_{\mu\nu} + \eta_\mu p_\nu + \eta_\nu p_\mu$ to leave only $\epsilon_{11}, \epsilon_{12}$:

$$\begin{pmatrix} \epsilon_{00} & -\epsilon_{13} & -\epsilon_{23} & -\frac{1}{2}(\epsilon_{00} + \epsilon_{33}) \\ & \epsilon_{11} & \epsilon_{12} & \epsilon_{13} \\ \text{symm} & & -\epsilon_{11} & \epsilon_{23} \\ & & & \epsilon_{33} \end{pmatrix} \Rightarrow \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & \epsilon_{11} & \epsilon_{12} & 0 \\ 0 & \epsilon_{12} & -\epsilon_{11} & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix}$$

$$\epsilon_{00} \rightarrow \epsilon_{00} + 2\eta_0 E = 0$$

$$\eta_0 = -\frac{\epsilon_{00}}{2E}$$

$$\epsilon_{13} \rightarrow \epsilon_{13} - \eta_1 E = 0$$

$$\eta_1 = \frac{\epsilon_{13}}{E}$$

$$\eta^\mu = \frac{1}{E} \left(\frac{\epsilon_{00}}{2}, \epsilon_{13}, \epsilon_{23}, \frac{\epsilon_{33}}{2} \right)$$

$$\epsilon_{23} \rightarrow \epsilon_{23} - \eta_2 E = 0$$

$$\eta_2 = \frac{\epsilon_{23}}{E}$$

$$\epsilon_{33} \rightarrow \epsilon_{33} - 2\eta_3 E = 0$$

$$\eta_3 = \frac{\epsilon_{33}}{2E}$$

Two polarisation states

$$\epsilon_{\mu\nu}^{(1)} = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & \frac{1}{\sqrt{2}} & 0 & 0 \\ 0 & 0 & -\frac{1}{\sqrt{2}} & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix} \quad \epsilon_{\mu\nu}^{(2)} = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & \frac{1}{\sqrt{2}} & 0 \\ 0 & \frac{1}{\sqrt{2}} & 0 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix} \quad \sum_{\epsilon_1} \epsilon_{\mu\nu}^{(r)} \sum_{\epsilon_2} \epsilon_{\alpha\beta}^{(s)} = \frac{1}{2} (\eta_{\mu\alpha} \eta_{\nu\beta} + \eta_{\mu\beta} \eta_{\nu\alpha} - \eta_{\mu\nu} \eta_{\alpha\beta})$$

Now rotate around 3-axis $x^\mu \rightarrow x^{\mu'} = \Lambda^{\mu'}_{\nu} x^\nu = R^{\mu'}_{\nu} x^\nu$

$$\epsilon_{\mu\nu} \rightarrow R_{\mu}^{\alpha} R_{\nu}^{\beta} \epsilon_{\alpha\beta}$$

$$R_{\mu}^{\alpha} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & \cos\theta & \sin\theta & 0 \\ 0 & -\sin\theta & \cos\theta & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

$$\begin{aligned} \epsilon'_{11} &= R_1^{\alpha} R_1^{\beta} \epsilon_{\alpha\beta} = R_1^1 R_1^1 \epsilon_{11} + (R_1^1 R_1^2 + R_1^2 R_1^1) \epsilon_{12} + R_1^2 R_1^2 \epsilon_{22} \\ &= (\cos^2\theta - \sin^2\theta) \epsilon_{11} + \sin 2\theta \epsilon_{12} \end{aligned}$$

$$\begin{aligned} \epsilon'_{12} &= (R_1^1 R_2^1 - R_1^2 R_2^2) \epsilon_{11} + (R_1^1 R_2^2 + R_1^2 R_2^1) \epsilon_{12} \\ &= -2\sin\theta \cos\theta \epsilon_{11} + (\cos^2\theta - \sin^2\theta) \epsilon_{12} \end{aligned}$$

$$\begin{cases} \epsilon'_{11} = \cos 2\theta \epsilon_{11} + \sin 2\theta \epsilon_{12} \\ \epsilon'_{12} = -\sin 2\theta \epsilon_{11} + \cos 2\theta \epsilon_{12} \end{cases}$$

$$\epsilon'_{11} - i \epsilon'_{12} = e^{i2\theta} \epsilon_{11} + (-i \cos 2\theta + \sin 2\theta) \epsilon_{12} - i e^{i2\theta}$$

$$\epsilon^{(+)} \boxed{\epsilon'_{11} - i \epsilon'_{12} = e^{i2\theta} (\epsilon_{11} - i \epsilon_{12})}$$

The state $(\epsilon_{11} - i \epsilon_{12}) e^{-ipx}$ has helicity $+\frac{2}{\hbar}$;
 (there are also two $h=0$ comps $\epsilon_{00}, \epsilon_{33}$ and $h=\pm 1$ comps $\epsilon_{31}, \epsilon_{23}$ which were gauge transformed away)

Graviton scattering

in quantised weak-field gravity

$$\text{Full: } S = \int d^4x \left\{ -\frac{1}{16\pi G} \sqrt{|g|} (R - \frac{2\Lambda}{d}) + \sqrt{|g|} \mathcal{L}_m(g_{\mu\nu}) \right\} \quad \frac{1}{2} g^{\mu\nu} \nabla_\mu \varphi \nabla_\nu \varphi$$

related to

$$\delta\sqrt{|g|} = -\frac{1}{2}\sqrt{|g|} g_{\mu\nu} \delta g^{\mu\nu}$$

$$T_{\mu\nu} = \epsilon_{vac} g_{\mu\nu}$$

$$\Rightarrow \frac{\delta S}{\delta g^{\mu\nu}} = - \int d^4x \left\{ \frac{1}{16\pi G} \sqrt{|g|} \left(R_{\mu\nu} - \frac{1}{2} (R - 2\Lambda) g_{\mu\nu} + 8\pi G (g_{\mu\nu} \mathcal{L}_m - 2 \frac{\partial \mathcal{L}_m}{\partial g^{\mu\nu}}) \right) \right\}$$

$$\text{EOM: } R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} + \Lambda g_{\mu\nu} = 8\pi G T_{\mu\nu}$$

$$2 \frac{\partial \mathcal{L}_m}{\partial g^{\mu\nu}} - g_{\mu\nu} \mathcal{L}_m = -\frac{2}{\sqrt{|g|}} \frac{\delta S}{\delta g^{\mu\nu}}$$

Weak field

$$\begin{cases} \mathcal{L}_m = 0 \\ R - 2R + 4\Lambda = 0 \\ \Rightarrow R = 4\Lambda \Rightarrow R_{\mu\nu} = \Lambda g_{\mu\nu} \end{cases}$$

$$R = g^{\mu\nu} R_{\mu\nu} \quad \begin{cases} g_{\mu\nu} = \eta_{\mu\nu} + \kappa h_{\mu\nu} \\ g^{\mu\nu} = \eta^{\mu\nu} - \kappa h^{\mu\nu} + \kappa^2 h^{\mu\alpha} h_{\alpha}{}^{\nu} + \dots \end{cases}$$

like $F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$

$$\sqrt{|g|} = e^{\frac{1}{2} \kappa \log g} = 1 + \kappa h - \frac{1}{4} \kappa^2 h_{\alpha\beta} h^{\alpha\beta} + \frac{\kappa^2}{8} h^2 + \dots$$

$$R_{\mu\alpha\nu\beta} = R_{[\mu\alpha][\nu\beta]} = -\frac{\kappa}{2} \left[\partial_\mu \partial_\nu h_{\alpha\beta} - \partial_\mu \partial_\beta h_{\alpha\nu} - \partial_\alpha \partial_\nu h_{\mu\beta} + \partial_\alpha \partial_\beta h_{\mu\nu} \right] + \dots$$

antis within pairs, symm in pairs

$$R_{\mu\nu} = \eta^{\alpha\beta} R_{\mu\alpha\nu\beta} = \frac{\kappa}{2} \left[-\partial^2 h_{\mu\nu} + \partial_\mu \partial_\alpha h^\alpha{}_\nu + \partial_\nu \partial_\alpha h^\alpha{}_\mu - \partial_\mu \partial_\nu h \right] + \dots$$

$$R = \eta^{\mu\nu} R_{\mu\nu} = \kappa \left[-\partial^2 h + \partial_\mu \partial_\nu h^{\mu\nu} \right] + \dots$$

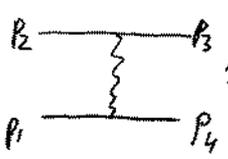
$$S = \int d^4x \left[\frac{1}{2} h^{\mu\nu} D^{-1}_{\mu\nu\alpha\beta} h^{\alpha\beta} + \mathcal{O}(h^3) + \kappa^2 \mathcal{O}(h^4) \right]$$

$$\frac{1}{\text{GeV}^4} \left[\text{GeV} \quad \text{GeV}^2 \quad \text{GeV} + \text{GeV}^2 \frac{1}{\text{GeV}} \text{GeV}^3 + \text{GeV}^2 \frac{1}{\text{GeV}^2} \text{GeV}^4 \right]$$

two momenta

For  and  see DeWitt PR 162 (1967) 1939
 eqs (2.6) & (2.7)
 PP 7777 or Bjerrum-Bohr hep-th/0410097
 171 terms 2850 terms eqs (4.10), (4.11)

Example (from DeWitt): scalar scattering



$$\approx i \delta^4(p) \left\{ \frac{1}{2\sqrt{E_2 E_3}} [P_2^\mu P_3^\nu + P_3^\mu P_2^\nu - \eta_{\mu\nu} (P_2 \cdot P_3 + m^2)] \right.$$

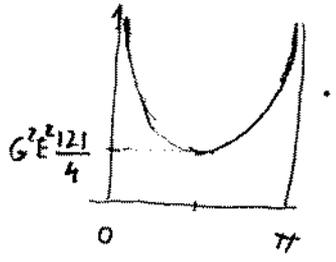
$$\left. \frac{1}{2} (\eta_{\mu\alpha} \eta_{\nu\beta} + \eta_{\mu\beta} \eta_{\nu\alpha} - \eta_{\mu\nu} \eta_{\alpha\beta}) \right\}$$

$$\frac{1}{2\sqrt{E_1 E_4}} [P_1^\alpha P_4^\beta + \dots - \eta^{\alpha\beta} (\dots)]$$

$$\int d^4x \left(\frac{1}{2} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi - \frac{1}{2} m^2 \phi^2 \right) \quad \begin{matrix} p & k \\ \text{---} & \text{---} \\ & \text{---} \end{matrix}$$

$$T_{\mu\nu} = \partial_\mu \phi \partial_\nu \phi - g_{\mu\nu} \mathcal{L} \Rightarrow p_\mu k_\nu + p_\nu k_\mu - \eta_{\mu\nu} (p \cdot k + m^2)$$

$$\frac{d\sigma}{d\Omega} \approx_{\substack{E \gg m \\ \theta \approx 0}} G^2 E^2 \frac{4}{\sin^4 \frac{\theta}{2}} + \dots$$



Overall magnitude $\left(\frac{E}{M_{pl}}\right)^2 \cdot \frac{1}{M_{pl}^2} = 10^{-38} 0.2 \text{ fm}^2 = \text{infinitesimal, of course}$

Gluonic processes $\left| \text{diagram} + \dots \right|^2$ clearly exceedingly complicated, but the MHV amp on p. 32 is again very simple!

$\Rightarrow \frac{d\sigma}{d\Omega}$ as above