

## Power Spectra

- Statistical homogeneity  $\Rightarrow$  different Fourier modes are uncorrelated

$$\langle \psi(\vec{l})^* \psi(\vec{l}') \rangle = (2\pi)^2 \delta_D^2(\vec{l} - \vec{l}') P_\psi(\vec{l}) \quad (22)$$

Using (19),

$$\langle x(\vec{l})^* x(\vec{l}') \rangle = \frac{1}{4} L^4 \langle \psi(\vec{l})^* \psi(\vec{l}') \rangle \Rightarrow P_x(\vec{l}) = \frac{1}{4} L^4 P_\psi(\vec{l})$$

$$\langle y_1(\vec{l})^* y_1(\vec{l}') \rangle = \frac{1}{4} (L_1^2 - L_2^2)^2 \langle \psi(\vec{l})^* \psi(\vec{l}') \rangle \Rightarrow P_{y_1}(\vec{l}) = \frac{1}{4} (L_1^2 - L_2^2)^2 P_\psi(\vec{l})$$

$$\langle y_2(\vec{l})^* y_2(\vec{l}') \rangle = L_1^2 L_2^2 \langle \psi(\vec{l})^* \psi(\vec{l}') \rangle \Rightarrow P_{y_2}(\vec{l}) = L_1^2 L_2^2 P_\psi(\vec{l})$$

$$\therefore P_{y_1}(\vec{l}) + P_{y_2}(\vec{l}) = \frac{1}{4} [(L_1^2 - L_2^2)^2 + 4L_1^2 L_2^2] P_\psi(\vec{l}) = \frac{1}{4} L^4 P_\psi(\vec{l}) = P_x(\vec{l})$$

- $\psi$  and  $x$  are scalar quantities. Statistical isotropy  $\Rightarrow P_\psi(\vec{l}) = P_\psi(l)$  and  $P_x(\vec{l}) = P_x(l)$

$y$  is a polar quantity. Its components  $y_1$  and  $y_2$  depend on the orientation of the ch. system. This dependence on the ch. system breaks the isotropy of  $y_1$  and  $y_2$ , and thus  $P_{y_1}(\vec{l})$  and  $P_{y_2}(\vec{l})$  depend on the direction of the 2D wave vector  $\vec{l}$ .

So far we have worked only with Fourier transforms and power spectra of real quantities, to be on more familiar ground. But we can also define a power spectrum for the complex  $y = y_1 + i y_2$ . Do first the correlator

$$\langle y_1(\vec{l})^* y_2(\vec{l}') \rangle = \frac{1}{2} L_1 L_2 (L_1^2 - L_2^2) \langle \psi(\vec{l})^* \psi(\vec{l}') \rangle \Rightarrow y_1(\vec{l}) \text{ and } y_2(\vec{l}) \text{ are}$$

correlated; this correlator is real (since  $\langle \psi(\vec{l})^* \psi(\vec{l}') \rangle$  is)

$$\therefore \langle y(\vec{l})^* y(\vec{l}') \rangle = \langle [y_1(\vec{l})^* - i y_2(\vec{l})^*] [y_1(\vec{l}') + i y_2(\vec{l}')] \rangle$$

$$= \underbrace{\langle y_1(\vec{l})^* y_1(\vec{l}') \rangle}_{\text{these are equal (complex conjugates of a real quantity),}} + i \underbrace{\langle y_1(\vec{l})^* y_2(\vec{l}') \rangle}_{\text{so they cancel}} - i \underbrace{\langle y_2(\vec{l})^* y_1(\vec{l}') \rangle}_{\text{}} + \langle y_2(\vec{l})^* y_2(\vec{l}') \rangle$$

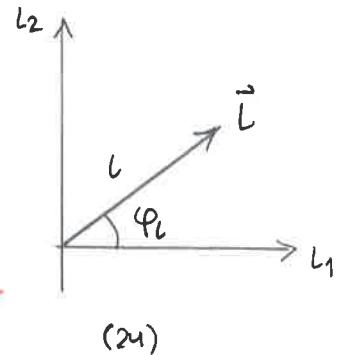
so they cancel

Defining  $P_g(\vec{l})$  by  $\langle g(\vec{l})^* g(\vec{l}') \rangle \equiv (2\pi)^2 S_D^2(\vec{l}-\vec{l}') P_g(\vec{l})$  we have thus

$$\underline{P_g(\vec{l}) = P_{g_1}(\vec{l}) + P_{g_2}(\vec{l}) = P_x(l)} \quad (23) \quad \therefore P_g(l) \text{ is isotropic}$$

In terms of length and direction of the 2D wave vector  $\vec{l}$

$$l_1 = l \cos \varphi_l \quad l_2 = l \sin \varphi_l$$



$$\underline{P_{g_1}(\vec{l}) = \frac{(l_1^2 - l_2^2)^2}{l^4} P_x(l) = (\cos^2 \varphi_l - \sin^2 \varphi_l) P_x(l) = \cos^2 2\varphi_l \cdot P_x(l)}$$

$$\underline{P_{g_2}(\vec{l}) = 4 \frac{l_1^2 l_2^2}{l^4} P_x(l) = 4 \cos^2 \varphi_l \sin^2 \varphi_l P_x(l) = \sin^2 2\varphi_l \cdot P_x(l)}$$

Or directly for  $g(\vec{l}) = g_1(\vec{l}) + i g_2(\vec{l})$ :

$$\underline{g(\vec{l}) = \frac{l_1^2 - l_2^2 + 2il_1l_2}{l^2} h(\vec{l}) = \frac{(l_1 + il_2)^2}{l^2} h(\vec{l}) = e^{i2\varphi_l} h(\vec{l})} \quad (25)$$

$$\Rightarrow P_g(\vec{l}) = |e^{i2\varphi_l}|^2 P_x(l) = P_x(l)$$

Here  $l = \sqrt{l_1^2 + l_2^2} = |\vec{l}| = |l_1 + il_2|$  is both the length of the vector  $\vec{l}$

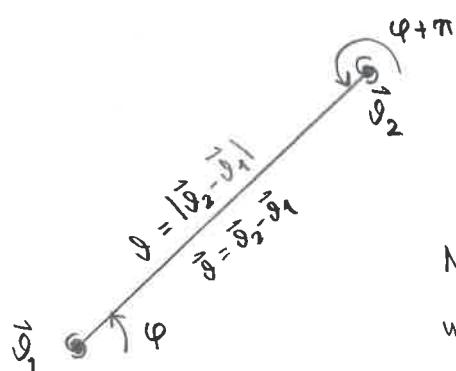
and the modulus (absolute value) of the complex number  $l_1 + il_2$ .

Thus we can relate the convergence power spectrum  $P_x(l)$  to the shear power spectrum  $P_g(l)$  - they are the same!

However, what we can get directly from observations, are the shear correlation functions rather than their power spectra.

### §5.4 Shear Correlation Functions

- We define the tangential and cross component of shear for a pair of points (galaxy images) using their separation line as the reference direction:



$$\begin{aligned}\gamma_t &\equiv -\text{Re}(\gamma e^{-i2\phi}) \\ \gamma_x &\equiv -\text{Im}(\gamma e^{-i2\phi})\end{aligned}\quad (26)$$

Note that since  $e^{-i2\phi} = e^{-i2(\phi+\pi)}$ , it doesn't matter which way the separation direction is defined.

- We define the correlation functions

$$\begin{aligned}\langle \gamma_t(\vec{\delta}_1) \gamma_t(\vec{\delta}_2) \rangle &= \langle \gamma_t \gamma_t \rangle(\vec{\delta}) = \langle \gamma_t \gamma_t \rangle(\vartheta) \\ \langle \gamma_x(\vec{\delta}_1) \gamma_x(\vec{\delta}_2) \rangle &= \langle \gamma_x \gamma_x \rangle(\vec{\delta}) = \langle \gamma_x \gamma_x \rangle(\vartheta)\end{aligned}\quad (27)$$

$$\langle \gamma_t(\vec{\delta}_1) \gamma_x(\vec{\delta}_2) \rangle = \langle \gamma_t \gamma_x \rangle(\vec{\delta}) = \underbrace{\langle \gamma_t \gamma_x \rangle(\vartheta)}_{\text{statistical isotropy}} = \xi_x(\vartheta)$$

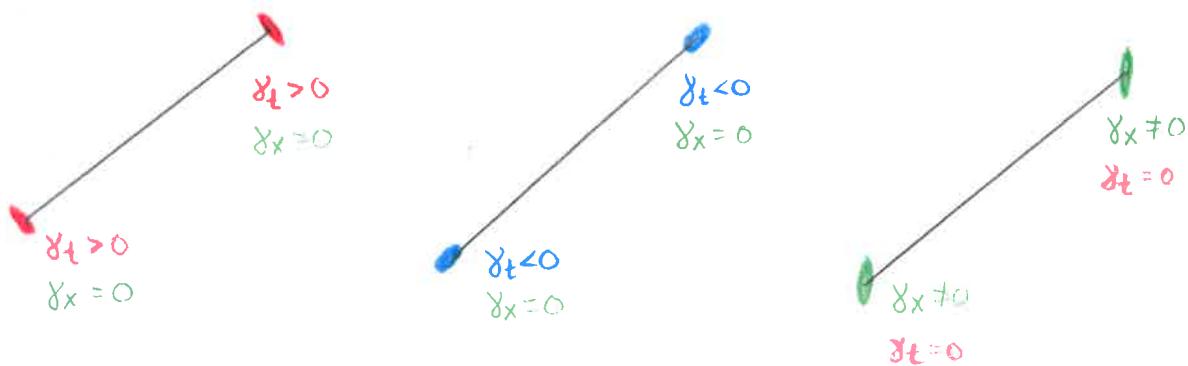
$$\xi_{\pm}(\vartheta) = \underbrace{\langle \gamma_t \gamma_t \rangle(\vartheta) \pm \langle \gamma_x \gamma_x \rangle(\vartheta)}_{\text{statistical homogeneity}} \quad (28)$$

- These are evaluated from the data simply by using galaxy image ellipticities as shear estimators: For each bin  $\vartheta \pm \frac{1}{2}\Delta\vartheta$ , find all  $N$  image pairs whose separation falls in this range, and calculate the average:

$$\hat{\xi}_{\pm}(\vartheta) = \frac{1}{N} \sum_{i,j} (\epsilon_{i,t} \epsilon_{j,t} \pm \epsilon_{i,x} \epsilon_{j,x}) \quad (29)$$

Note that the tangential and cross components  $\epsilon_t, \epsilon_x$  for each galaxy image are defined differently for each pair that it is part of.

- Meaning of  $\gamma_t$  and  $\gamma_x$  in pictures.



The reason for the - sign in the definition of  $\gamma_t$ : This way it is typically expected to be positive: the more nearby of the pair of galaxies lenses the more distant one in the tangential direction. (Without the - sign we should call  $\gamma_t$  the radial component.)

- I didn't bother to check the sign of  $\gamma_x$  in the third picture; it will depend on how we define the coordinate axes: if we define a right-handed 3D coordinate system, it depends on whether we choose the third axis as pointing towards or away from the observer.
- Parity symmetry  $\Rightarrow \xi_x(0)$  is expected to vanish, since in a parity transformation (mirror universe)  $\gamma_t \rightarrow \gamma_t$  but  $\gamma_x \rightarrow -\gamma_x$ .
- We should now relate  $\xi_{\pm}(0)$  to  $P_{\pm}(L)$ .

I don't know if there's a clever way to do this directly with the complex shear; I did this by falling back to  $\gamma_1$  and  $\gamma_2$ , defining first the correlation functions

$$\xi_1(\vec{\delta}) = \langle \gamma_1(\vec{\delta}_1) \gamma_1(\vec{\delta}_2) \rangle \quad \text{where } \vec{\delta} = \vec{\delta}_2 - \vec{\delta}_1$$

$$\xi_2(\vec{\delta}) = \langle \gamma_2(\vec{\delta}_1) \gamma_2(\vec{\delta}_2) \rangle$$

$$\xi_{12}(\vec{\delta}) = \langle \gamma_1(\vec{\delta}_1) \gamma_2(\vec{\delta}_2) \rangle$$

These correlation functions are not expected to be isotropic, since the components depend on the orientation of the lab. system.

But statistical homogeneity  $\Rightarrow$  they depend only on the separation  $\vec{\delta}$ .  
 (On next page I call  $\vec{\delta}_1 = \vec{\delta}_0$  and  $\vec{\delta}_2 = \vec{\delta}_0 + \vec{\delta}$ ).

The correlation functions  $\tilde{\gamma}_1(\vec{\delta})$  and  $\tilde{\gamma}_2(\vec{\delta})$  are Fourier transforms of their power spectra  $P_{\gamma_1}(\vec{l})$  and  $P_{\gamma_2}(\vec{l})$ .

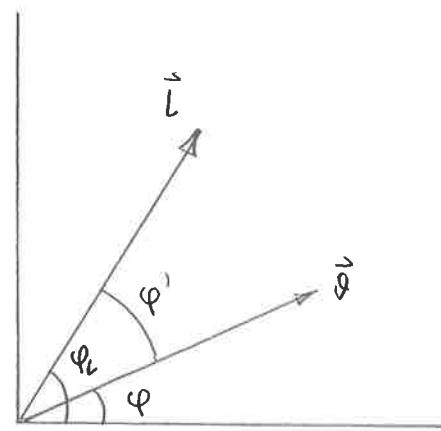
We have to deal with two directions of 2D vectors:

that at  $\vec{l}$  ( $\varphi_l$ )

and that at  $\vec{\delta}$  ( $\varphi$ )

$\varphi' \equiv \varphi_l - \varphi$  is the angle between  $\vec{l}$  and  $\vec{\delta}$

$$\Rightarrow \vec{l} \cdot \vec{\delta} = l \delta \cos \varphi'$$



$$\begin{aligned} \tilde{\gamma}_1(\vec{\delta}) &\equiv \langle \gamma_1(\vec{\delta}_0) \gamma_1(\vec{\delta}_0 + \vec{\delta}) \rangle = \frac{1}{(2\pi)^2} \int d^2 l P_{\gamma_1}(l) e^{i \vec{l} \cdot \vec{\delta}} = \frac{1}{(2\pi)^2} \int d^2 l P_{\gamma_1}(l) e^{i l \delta \cos \varphi'} \\ &= \frac{1}{(2\pi)^2} \int d^2 l P_x(l) \cos^2 2\varphi_l e^{i l \delta \cos \varphi'} \end{aligned}$$

Using  $\int d^2 l = \int_0^\infty l dl \int_0^{2\pi} d\varphi_l = \int_0^\infty l dl \int_0^{2\pi} d\varphi$  and the integral representation of

Bessel functions  $J_n(x) = \frac{(-i)^n}{\pi} \int_0^\pi e^{ix \sin \varphi} \cos n\varphi d\varphi = \frac{(-i)^n}{2\pi} \int_0^{2\pi} e^{ix \sin \varphi} \cos n\varphi d\varphi$  (30)

one finds (exercise)

$$\tilde{\gamma}_1(\vec{\delta}) = \frac{1}{4\pi} \int_0^\infty l dl P_x(l) [J_0(l\delta) + (\omega^2 2\varphi - \sin^2 2\varphi) J_4(l\delta)] \quad (31)$$

$$\tilde{\gamma}_2(\vec{\delta}) = \frac{1}{4\pi} \int_0^\infty l dl P_x(l) [J_0(l\delta) + (\sin^2 2\varphi - \omega^2 2\varphi) J_4(l\delta)]$$

Also

$$\begin{aligned} \tilde{\gamma}_{12}(\vec{\delta}) &\equiv \langle \gamma_1(\vec{\delta}_0) \gamma_2(\vec{\delta}_0 + \vec{\delta}) \rangle = \frac{1}{(2\pi)^4} \int d^2 l d^2 l' e^{-i \vec{l} \cdot \vec{\delta}_0} e^{+i \vec{l}' \cdot (\vec{\delta}_0 + \vec{\delta})} \langle \gamma_1(\vec{l})^* \gamma_2(\vec{l}') \rangle \\ &= \dots = \frac{1}{4\pi} \int_0^\infty l dl P_x(l) \cdot 2 \sin 2\varphi \omega 2\varphi \cdot J_4(l\delta) \end{aligned} \quad \text{statistical homogeneity} \quad (32)$$

$$\text{and } \tilde{\gamma}_{21}(\vec{\delta}) \equiv \langle \gamma_2(\vec{\delta}_0) \gamma_1(\vec{\delta}_0 + \vec{\delta}) \rangle = \langle \gamma_1(\vec{\delta}_0 + \vec{\delta}) \gamma_2(\vec{\delta}_0) \rangle = \langle \gamma_1(\vec{\delta}_0) \gamma_2(\vec{\delta}_0 - \vec{\delta}) \rangle$$

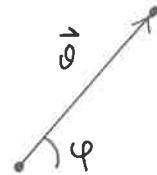
$$\underline{\underline{\tilde{\gamma}_{12}(-\vec{\delta}) = \tilde{\gamma}_{12}(\vec{\delta})}} \quad \text{since } \sin 2\varphi \omega 2\varphi \text{ is invariant in } \varphi \rightarrow \varphi + \pi \quad (\vec{\delta} \rightarrow -\vec{\delta})$$

Return now to  $\gamma_t$  and  $\gamma_x$ :

$$\gamma e^{-i2\varphi} = (\gamma_1 + i\gamma_2)(\cos 2\varphi - i \sin 2\varphi) = \gamma_1 \cos 2\varphi + \gamma_2 \sin 2\varphi + i[-\gamma_1 \sin 2\varphi + \gamma_2 \cos 2\varphi]$$

$$\Rightarrow \gamma_t \equiv -\operatorname{Re}(\gamma e^{-i2\varphi}) = -\gamma_1 \cos 2\varphi - \gamma_2 \sin 2\varphi \quad (33)$$

$$\gamma_x \equiv -\operatorname{Im}(\gamma e^{-i2\varphi}) = \gamma_1 \sin 2\varphi - \gamma_2 \cos 2\varphi$$



$$\Rightarrow \langle \gamma_t \gamma_t \rangle(\vec{\vartheta}) = \dots = \frac{1}{4\pi} \int_0^\infty l dl P_x(l) \cdot [J_0(l\vartheta) + J_4(l\vartheta)]$$

$$\langle \gamma_x \gamma_x \rangle(\vec{\vartheta}) = \dots = \frac{1}{4\pi} \int_0^\infty l dl P_x(l) \cdot [J_0(l\vartheta) - J_4(l\vartheta)]$$

$$\therefore S_+(\vartheta) \equiv \langle \gamma_t \gamma_t \rangle + \langle \gamma_x \gamma_x \rangle = \frac{1}{2\pi} \int_0^\infty l dl J_0(l\vartheta) P_x(l) \quad (34)$$

$$S_-(\vartheta) \equiv \langle \gamma_t \gamma_t \rangle - \langle \gamma_x \gamma_x \rangle = \frac{1}{2\pi} \int_0^\infty l dl J_4(l\vartheta) P_x(l)$$

Also  $S_x(\vec{\vartheta}) \equiv \langle \gamma_t \gamma_x \rangle(\vec{\vartheta}) = \dots = 0$

We can invert these equations using the Bessel function closure (orthogonality) equation

$$\int_0^\infty J_n(\alpha x) J_n(\alpha' x) x dx = \frac{1}{\alpha} \delta_D(\alpha - \alpha') \quad (35)$$

$$2\pi \int_0^\infty d\vartheta \delta S_+(\vartheta) J_0(l\vartheta) = \int_0^\infty d\vartheta \int_0^\infty d\vartheta' l' \delta J_0(l\vartheta) J_0(l'\vartheta) P_x(l')$$

$$= \int_0^\infty d\vartheta' l' P_x(l') \underbrace{\int_0^\infty J_0(l\vartheta) J_0(l'\vartheta) \delta d\vartheta}_{\frac{l}{l'} \delta_D(l-l')} = P_x(l) \quad (36a)$$

likewise

$$2\pi \int_0^\infty d\vartheta \delta S_-(\vartheta) J_4(l\vartheta) = \dots = P_x(l) \quad (36b)$$

Thus the two shear correlation functions  $\xi_+(\vartheta)$  and  $\xi_-(\vartheta)$  are not independent (as the two components of shear are both derived from the same scalar quantity, the deflection potential  $\psi$ ). We can express them in terms of each other by inserting (36b) into (34a) and (36a) into (34b). Schreiber (p.364) gives the results

$$\xi_+(\vartheta) = \frac{1}{2\pi} \int_0^\infty L dL J_0(L\vartheta) P_{xx}(L) \stackrel{(36b)}{=} \int_0^\infty L dL \int_0^\infty \vartheta' d\vartheta' J_0(L\vartheta) J_4(L\vartheta') \xi_-(\vartheta') \quad (37a)$$

$$= \int_0^\infty \vartheta' d\vartheta' \xi_-(\vartheta') \int_0^\infty L dL J_0(L\vartheta) J_4(L\vartheta') = \dots = \xi_-(\vartheta) + \int_0^\infty \frac{d\vartheta'}{\vartheta'} \xi_-(\vartheta') \left( 4 - 12 \frac{\vartheta'^2}{\vartheta^2} \right)$$

$$\xi_-(\vartheta) = \frac{1}{2\pi} \int_0^\infty L dL J_4(L\vartheta) P_{xx}(L) = \dots = \xi_+(\vartheta) + \int_0^\vartheta \frac{d\vartheta'}{\vartheta'^2} \xi_+(\vartheta') \left( 4 - 12 \frac{\vartheta'^2}{\vartheta^2} \right) \quad (37b)$$

I haven't so far managed to do those. Tools for doing the integral over the Bessel functions product include:

- Bessel function recursion formulae, e.g.,  $J_{n-1}(x) + J_{n+1}(x) = \frac{2n}{x} J_n(x)$
- The closure relation (35), which can be used to give the first term in (37a,b)
- Gradshteyn & Ryzhik integral 6.512.3:

$$\int_0^\infty J_n(\alpha x) J_{n-1}(\beta x) = \begin{cases} 0 & \text{if } \alpha < \beta \\ \frac{1}{2\beta} & \text{if } \alpha = \beta \\ \frac{\beta^{n-1}}{\alpha^n} & \text{if } \alpha > \beta \end{cases}$$

which can be used to cut the integral  $\int_0^\infty d\vartheta'$  to  $\int_0^\infty d\vartheta'$  or  $\int_\vartheta^\infty d\vartheta'$

(The exact forms of (37a,b) are maybe not that important; just the principle is.)