

Y4. Properties of Tensor Spherical Harmonics

The scalar spherical harmonics are $Y_l^m(\vartheta, \varphi) = (-1)^m \sqrt{\frac{2L+1}{4\pi}} \sqrt{\frac{(L-m)!}{(L+m)!}} P_l^m(\cos\vartheta) e^{im\varphi}$ (1)

where the associated Legendre functions $P_l^m(\cos\vartheta)$ are real and

$$P_l^{-m} = (-1)^m \frac{(L-m)!}{(L+m)!} P_l^m \quad (2)$$

$$\Rightarrow \underline{Y_l^{-m}} = (-1)^{-m} \sqrt{\frac{2L+1}{4\pi}} \sqrt{\frac{(L+m)!}{(L-m)!}} P_l^{-m} e^{-im\varphi} = \sqrt{\frac{2L+1}{4\pi}} \sqrt{\frac{(L-m)!}{(L+m)!}} P_l^m e^{-im\varphi} = \underline{(-1)^m Y_l^m}^* \quad (3)$$

From this follows that also (using Eqs. (3.4) and (3.13))

$Y_{l,-m}^{G/C}$	$=$	$(-1)^m Y_{lm}^{G/C}$	$*$
$W_{l,-m}$	$=$	$(-1)^m W_{lm}$	$*$
$X_{l,-m}$	$=$	$(-1)^m X_{lm}$	$*$

(4)

From (3.14) we can calculate the W_{lm} and X_{lm} in terms of the $P_l^m(\cos\vartheta)$: (exercise)

$W_{lm}(\vartheta, \varphi) = l(l+1)Y_l^m + 2\partial_\vartheta^2 Y_l^m = \dots = (-1)^m \cdot 2 \sqrt{\frac{2L+1}{4\pi}} \sqrt{\frac{(L-m)!}{(L+m)!}} G_{lm}^+(\cos\vartheta) e^{im\varphi}$	(5)
$-iX_{lm}(\vartheta, \varphi) = \frac{2m}{\sin^2\vartheta} (\partial_\vartheta - \cot\vartheta) Y_l^m = \dots = (-1)^m \cdot 2 \sqrt{\frac{2L+1}{4\pi}} \sqrt{\frac{(L-m)!}{(L+m)!}} G_{lm}^-(\cos\vartheta) e^{im\varphi}$	

where

$G_{lm}^+(\cos\vartheta) \equiv - \left[\frac{L-m^2}{\sin^2\vartheta} + \frac{1}{2}l(l-1) \right] P_l^m(\cos\vartheta) + (L+m) \frac{\cos\vartheta}{\sin^2\vartheta} P_{l-1}^m(\cos\vartheta)$	(6)
$G_{lm}^-(\cos\vartheta) \equiv \frac{m}{\sin^2\vartheta} \left[(l-1)\cos\vartheta \cdot P_l^m(\cos\vartheta) - (L+m)P_{l-1}^m(\cos\vartheta) \right]$	

are real,

where we have used the associated Legendre equation

$$\frac{1}{\sin^2\vartheta} \frac{d}{d\vartheta} \left(\sin^2\vartheta \cdot \frac{d}{d\vartheta} P_l^m(\cos\vartheta) \right) + \left[l(l+1) - \frac{m^2}{\sin^2\vartheta} \right] P_l^m(\cos\vartheta) = 0 \quad (7)$$

$$\Rightarrow \frac{d^2}{d\vartheta^2} P_l^m(\cos\vartheta) = x \frac{dP_l^m}{dx} + \frac{m^2}{1-x^2} P_l^m - l(l+1)P_l^m \quad \text{where } x = \cos\vartheta$$

and the recursion relation (Abramowitz & Stegun 8.5.4)

$$(x^2-1) \frac{d}{dx} P_l^m = lx P_l^m - (L+m)P_{l-1}^m \quad (8)$$

We see that the φ -dependence of $W_{lm}, X_{lm}, Y_{lm}^G, Y_{lm}^C$ is trivial, $e^{im\varphi}$, and that for $m=0$: W_{l0} is real and X_{l0} is imaginary

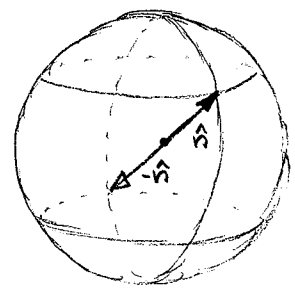
Antipodes

- The two directions opposite to each other, \hat{n} and $-\hat{n}$, have ends

$$\hat{n} = (\vartheta, \varphi) \quad \text{and} \quad -\hat{n} = (\vartheta, \bar{\varphi}) = (\pi - \vartheta, \varphi + \pi)$$

$$\Rightarrow x^- = \cos \vartheta^- = -\cos \vartheta = -x$$

$$\underline{e^{im\varphi^-} = e^{im\pi} e^{im\varphi} = (-1)^m e^{im\varphi}}$$



- The parity of the associated Legendre functions is $\underline{P_l^m(-x) = (-1)^{L+m} P_l^m(x)}$ (9)

$$\therefore \underline{Y_l^m(-\hat{n}) = (-1)^m \cdot (-1)^{L+m} Y_l^m(\hat{n}) = (-1)^L Y_l^m(\hat{n})}$$
 (10)

- From the definitions of the G_{lm}^\pm , Eq. (6), we get easily (exercise)

$$\underline{G_{lm}^\pm(-x) = \pm (-1)^{L+m} G_{lm}^\pm(x)}$$

\Rightarrow

$$\begin{aligned} W_l^m(-\hat{n}) &= (-1)^L W_l^m(\hat{n}) \\ X_l^m(-\hat{n}) &= -(-1)^L X_l^m(\hat{n}) \end{aligned}$$

(11)

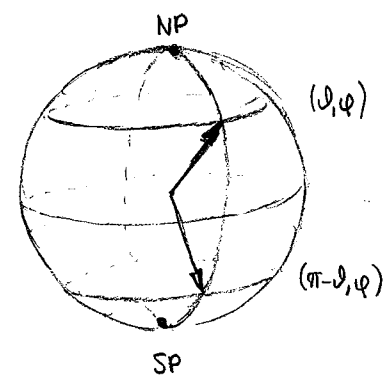
North-South reflection

- If we just make a reflection wrt the equator, i.e.,

$$\vartheta \rightarrow \pi - \vartheta \quad \Rightarrow \quad \cos \vartheta \rightarrow -\cos \vartheta$$

but keep φ the same, we do not get the

$e^{im\pi} = (-1)^m$ factor, and we have the results



$$\begin{aligned} Y_l^m(\pi - \vartheta, \varphi) &= (-1)^{L+m} Y_l^m(\vartheta, \varphi) \\ W_l^m(\pi - \vartheta, \varphi) &= (-1)^{L+m} W_l^m(\vartheta, \varphi) \\ X_l^m(\pi - \vartheta, \varphi) &= -(-1)^{L+m} X_l^m(\vartheta, \varphi) \end{aligned}$$

(12)

North and South Poles

- The north and south poles (NP and SP) of the ϑ, φ coord. system correspond to $\vartheta = 0, \pi$.

$$x = \cos\vartheta = +1, -1$$

The φ coordinate is multivalued at these points (all values of φ correspond to the same point).

The $P_l^m(x)$ have the values $P_l^m(\pm 1) = 0$ for $m \neq 0$ (13)

$$P_l(1) = 1 \quad \text{and} \quad P_l(-1) = (-1)^l$$

From this we get

$$\begin{aligned} Y_l^m(\text{NP}) &= \sqrt{\frac{2l+1}{4\pi}} \delta_{m0} \\ Y_l^m(\text{SP}) &= (-1)^l \sqrt{\frac{2l+1}{4\pi}} \delta_{m0} \end{aligned} \quad (14)$$

- The $G_{lm}^\pm(\pm 1)$ are tricky to derive, since from (6) & (13) we just get $\frac{0}{0}$ or $\infty - \infty$, as $\sin\vartheta = 0$ at the poles. To evaluate them, we need to take the limits $\vartheta \rightarrow 0, \pi$.

From Cabella & Kamionkowski astro-ph/0403392 Eqs. (75) and (76),

$$\begin{aligned} P_l^m(\cos\vartheta) &\sim \frac{(-1)^m}{2^m m!} \frac{(l+m)!}{(l-m)!} \vartheta^m \quad \text{for } m > 0 \\ &\quad \text{as } \vartheta \rightarrow 0 \\ P_l(\cos\vartheta) &\sim 1 - \frac{1}{4} l(l+1) \vartheta^2 \end{aligned} \quad (15)$$

Using this in (6) we get (exercise) for $m \geq 0$, that

$$\begin{aligned} G_{l2}^\pm(1) &= \frac{1}{4} \frac{(l+2)!}{(l-2)!} \Rightarrow W_{l2}(0, \varphi) = \frac{1}{N_l} \sqrt{\frac{2l+1}{8\pi}} e^{i2\varphi} \\ X_{l2}(0, \varphi) &= \frac{i}{N_l} \sqrt{\frac{2l+1}{8\pi}} e^{i2\varphi} \end{aligned}$$

but other $G_{lm}^\pm(1) = W_{lm}(0, \varphi) = X_{lm}(0, \varphi) = 0$.

Using (4) to get the negative m , we have for the North Pole

$$\begin{aligned} W_l^m(0, \varphi) &= \frac{1}{N_l} \sqrt{\frac{2l+1}{8\pi}} (\delta_{m2} e^{i2\varphi} + \delta_{m,-2} e^{-i2\varphi}) \\ X_l^m(0, \varphi) &= \frac{i}{N_l} \sqrt{\frac{2l+1}{8\pi}} (\delta_{m2} e^{i2\varphi} - \delta_{m,-2} e^{-i2\varphi}) \end{aligned} \quad (16)$$

We get the South Pole using (12):

$$\begin{aligned} W_l^m(\pi, \varphi) &= \frac{(-1)^l}{N_l} \sqrt{\frac{2l+1}{8\pi}} (\delta_{m2} e^{i2\varphi} + \delta_{m,-2} e^{-i2\varphi}) \\ X_l^m(\pi, \varphi) &= \frac{-i(-1)^l}{N_l} \sqrt{\frac{2l+1}{8\pi}} (\delta_{m2} e^{i2\varphi} - \delta_{m,-2} e^{-i2\varphi}) \end{aligned} \quad (17)$$

Thus the $W_l^m(\vartheta, \varphi)$ and $X_l^m(\vartheta, \varphi)$ are multivalued at the poles for $m = \pm 2$!