

L5. Delta Function Approximation

(good for illustration only)

- To get more insight into the dominant effects, we note that the visibility function is peaked around $\eta_* \equiv \eta_{dec}$, and approximate it by a delta function:

$$\begin{aligned} g(\eta) &\approx \delta(\eta - \eta_*) & (1) \\ \Rightarrow \quad \bar{e}^{\tau(\eta)} &\approx \Theta(\eta - \eta_*) & (2) \end{aligned}$$

(step function)

That is, we approximate that the last scattering happened for each photon at exactly the same time.

- Here we completely ignore the effect of reionization. A straightforward generalization would be to use two delta functions (exercise)

$$g(\eta) \approx \bar{e}^{\tau_{reion}} \delta(\eta - \eta_*) + (1 - \bar{e}^{\tau_{reion}}) \delta(\eta - \eta_{reion}). \quad (3)$$

- With the approximation (1) & (2), Eq. (4.8) becomes

$$\begin{aligned} \frac{1}{2L+1} \Theta_L^0(\eta_0) &\approx \left[\Theta_0^0(\eta_*) + \phi(\eta_*) \right] j_L(k\eta_0 - k\eta_*) + V_b^{(0)}(\eta_*) j_L'(k\eta_0 - k\eta_*) \\ &+ \underbrace{\int_{\eta_*}^{\eta_0} d\eta (\phi' + \psi') j_L(k\eta_0 - k\eta)}_{\text{integrated Sachs-Wolfe effect}} + \underbrace{\frac{1}{10} \left[\Theta_2^0(\eta_*) - \sqrt{6} E_2^0(\eta_*) \right] j_L^{(20)}(k\eta_0 - k\eta_*)}_{\text{local quadrupole effect}} \end{aligned} \quad (4)$$

The effect of the local quadrupole is relatively small, and we drop it for the further discussion within the delta-function approximation. The result (4) is in the $k_{\perp} m$ space, i.e. for multipoles at a single Fourier mode. The role of the radial functions $j_L^{(l m)}$ is to project the plane-wave spatial dependence onto multipoles.

- In \vec{x}, \hat{v} space, Eq. (4), w/o the local quadrupole term, reads

$$\Theta(\eta_0, \vec{x}_{\text{obs}}, \hat{v}) \simeq \underbrace{\Theta_0^{\circ}(\eta_*) + \phi(\eta_*)}_{\text{effective temperature}} + \hat{v} \cdot \underbrace{\vec{v}_b(\eta_*)}_{\text{Doppler}} + \int_{\eta_*}^{\eta_0} d\eta \underbrace{(\phi' + \psi')}_{\text{ISW}} \quad (5)$$

which is Eq. (49) of Chapter 12 of my Cosmology II lecture notes; except that here we have approximated $\psi \simeq \phi$, and \hat{v} has the opposite sign, since it refers to the observation direction, whereas here it refers to the photon direction.

- For vector ($m = \pm 1$) and tensor ($m = \pm 2$) modes, and for polarization, the delta function approximation gives, from (4.8-10) and (4.12-13),

$$\frac{1}{2L+1} \Theta_L^{\pm 1}(\eta_0) \simeq [V_b^{(\pm 1)}(\eta_*) - B^{(\pm 1)}(\eta_*)] j_L^{(11)}(k\eta_0 - k\eta_*) + \int_{\eta_*}^{\eta_0} d\eta \frac{k}{\sqrt{3}} B^{(\pm 1)}(\eta) j_L^{(21)}(k\eta_0 - k\eta) + \frac{1}{10} [\Theta_2^{\pm 1}(\eta_*) - \sqrt{6} E_2^{\pm 1}(\eta_*)] j_L^{(21)}(k\eta_0 - k\eta_*) \quad (6)$$

$$\frac{1}{2L+1} \Theta_L^{\pm 2}(\eta_0) \simeq \frac{1}{10} [\Theta_2^{\pm 2}(\eta_*) - \sqrt{6} E_2^{\pm 2}(\eta_*)] j_L^{(22)}(k\eta_0 - k\eta_*) - \int_{\eta_*}^{\eta_0} d\eta h^{(\pm 2)}(\eta) j_L^{(20)}(k\eta_0 - k\eta) \quad (7)$$

$$\frac{1}{2L+1} E_L^m(\eta_0) \simeq \frac{-\sqrt{6}}{10} [\Theta_2^m(\eta_*) - \sqrt{6} E_2^m(\eta_*)] E_L^m(k\eta_0 - k\eta_*) \quad (8)$$

$$\frac{1}{2L+1} B_L^m(\eta_0) \simeq \frac{-\sqrt{6}}{10} [\Theta_2^m(\eta_*) - \sqrt{6} E_2^m(\eta_*)] B_L^m(k\eta_0 - k\eta_*) \quad (9)$$

- We note that the temperature anisotropy $\Theta_L^m(\eta_0)$ gets a contribution both from the scattering surface and from the metric perturbations along the line of sight.

The polarization, $E_L^m(\eta_0)$, $B_L^m(\eta_0)$, on the other hand, probes only the scattering surface.

Hence CMB polarization observations have the power to separate scattering-surface effects from line-of-sight effects; that are mixed in temperature anisotropy observations. (But note that reionization introduces another "scattering surface", that we have ignored here; so in polarization observations the effects from these two "scattering surfaces" are mixed. Since the reionization surface is closer by, its effects appear at larger angular scales, i.e., at lower multipoles.)