

L2. Line-of-Sight Integration

- We shall consider a single Fourier mode \vec{k} . Thus the brightness function and the linear polarization Stokes parameters in \vec{x}, \hat{n} space are

$$\Theta(\eta, \vec{x}, \hat{n}) = \Theta(\eta, \hat{n}) e^{i\vec{k} \cdot \vec{x}} \quad (1)$$

$$(Q+iU)(\eta, \vec{x}, \hat{n}) = (Q+iU)(\eta, \hat{n}) e^{i\vec{k} \cdot \vec{x}} \quad (2)$$

L2.1 Temperature Anisotropy

- Consider first just free streaming. It's effect in the ordinary (\vec{x}, \hat{n}) space is very simple:
From conformal time η_1 to η_2 each photon travels the comoving distance $r = \eta_2 - \eta_1$ in the direction \hat{n} of its momentum. Since all photons with the same \hat{n} travel together, the brightness function value $\Theta(\eta_1, \vec{x}, \hat{n})$ is transported to $\Theta(\eta_2, \vec{x} + r\hat{n}, \hat{n})$.

In other words, the brightness function at η_2, \vec{x}_2 can be reconstructed from its values at η_1 as (see Figure)

$$\Theta(\eta_2, \vec{x}_2, \hat{n}) = \Theta(\eta_1, \vec{x}_2 - r\hat{n}, \hat{n}) \quad (3)$$

- Consider now multipoles. Eq. (1) can be written as

$$\Theta(\eta, \vec{x}, \hat{n}) = \sum_{lm} (-i)^l \sqrt{\frac{4\pi}{2l+1}} \Theta_l^m(\eta) Y_l^m(\hat{n}) e^{i\vec{k} \cdot \vec{x}} = \sum_{lm} \Theta_l^m(\eta) G_{L\vec{k}}^m(\vec{x}, \hat{n}) \quad (4)$$

where the mode function $G_{L\vec{k}}^m$ is

$$G_{L\vec{k}}^m(\vec{x}, \hat{n}) \equiv (-i)^l \sqrt{\frac{4\pi}{2l+1}} Y_l^m(\hat{n}) e^{i\vec{k} \cdot \vec{x}} \quad (5)$$

and we use the coordinate system (for Y_l^m) where Z-axis is in the \vec{k} direction.

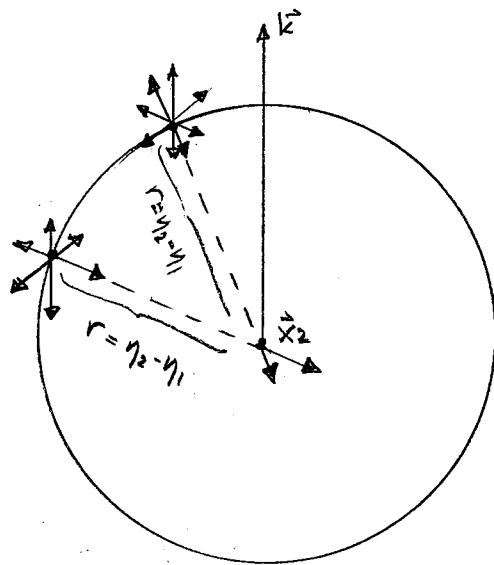


Fig. Free streaming of temperature anisotropy from time η_1 to η_2 .

Eq. (3) then becomes

$$\Theta(\gamma_2, \vec{x}_2, \hat{n}) = \sum_{l,m} \Theta_l^m(\gamma_1) G_{lk}^m(\vec{x}_2 - r\hat{n}, \hat{n}) = e^{i\vec{k} \cdot \vec{x}_2} \overbrace{\sum_{l,m} \Theta_l^m(\gamma_1) G_{lk}^m(r, \hat{n})}^{\Theta(\gamma_2, \hat{n})} \quad (6)$$

where we have defined the line-of-sight mode function $\vec{k} \cdot \vec{x} = -kr\hat{z} \cdot \hat{n} = -kr \cos \theta$

$$G_{lk}^m(r, \hat{n}) \equiv G_{lk}^m(-r\hat{n}, \hat{n}) = (-i)^l \sqrt{\frac{4\pi}{2l+1}} Y_l^m(\hat{n}) e^{-ikr\hat{z} \cdot \hat{n}} \quad (7)$$

(that is, only those pairs \vec{x}, \hat{n} are considered, where \vec{x} is in the $-\hat{n}$ direction).

To relate $\Theta_l^m(\gamma_2)$ to $\Theta_l^m(\gamma_1)$ we just need to pick out the multipole of Eq. (6).

For that we need the spherical harmonic expansion of the plane wave in (7):

$$\begin{aligned} e^{i\vec{k} \cdot \vec{x}} &= 4\pi \sum_{l,m} i^l j_l(kr) Y_l^m(\hat{x}) Y_l^{m*}(\hat{k}) & \begin{matrix} \vec{x} = -r\hat{n} \\ \hat{k} = \hat{z} \end{matrix} &= 4\pi \sum_{l,m} i^l j_l(kr) \underbrace{Y_l^m(-\hat{n})}_{(-1)^l Y_l^m(\hat{n})} \underbrace{Y_l^{m*}(\hat{z})}_{\sqrt{\frac{2l+1}{4\pi}} \delta_{m0}} \\ \Rightarrow e^{-ikr\hat{n} \cdot \hat{z}} &= \sqrt{4\pi} \sum_l (-i)^l \sqrt{2l+1} j_l(kr) Y_l^0(\hat{n}) & (8) & \end{aligned}$$

The angular dependence of $\Theta(\gamma_2, \hat{n})$ is a combination of the angular dependence of the plane wave, i.e. the Y_l^0 in (8), and the initial angular dependence at $\vec{x}_2 - r\hat{n}$, i.e. the Y_l^m in (7). The effect is mathematically like the addition of orbital (the plane wave) and spin (the initial) angular momentum into the total angular momentum.

To get the multipole $\Theta_l^m(\gamma_2)$ of $\Theta(\gamma_2, \hat{n})$, we need the total angular dependence of the line-of-sight mode functions Use now CG series, Eq. (FG.11)

$$\begin{aligned} G_{lk}^m(r, \hat{n}) &= 4\pi \sum_l (-i)^{l+l} \sqrt{\frac{2l+1}{2l'+1}} j_l(kr) Y_l^m(\hat{n}) Y_l^0(\hat{n}) \\ &= 4\pi \sum_l (-i)^{l+l} \sqrt{\frac{2l+1}{2l'+1}} j_l(kr) \sum_j \sqrt{\frac{(2l+1)(2l'+1)}{4\pi(2j+1)}} \langle l0l'm|jm \rangle \langle l0l'0|j0 \rangle Y_j^m(\hat{n}) \\ &= \sum_j (-i)^{l+l} (2l+1) \sqrt{\frac{4\pi}{2j+1}} \langle l0l'm|jm \rangle \langle l0l'0|j0 \rangle j_l(kr) Y_j^m(\hat{n}) \quad (9) \end{aligned}$$

where l represents the plane-wave, l' the initial and j the total angular dependence.

Thus the new brightness function, $\Theta(y_2, \hat{n})$ at Eq.(6) can be written

$$\begin{aligned} \Theta(y_2, \hat{n}) &= \sum_{l'm} \Theta_{l'}^m(y_1) G_{l'k}^m(r, \hat{n}) \\ &= \sum_{j'm} \sum_{l'L} \underbrace{\Theta_{l'}^m(y_1) (-i)^{l'+L} (2L+1) \sqrt{\frac{4\pi}{2j+1}} \langle l'0l'm|j'm \rangle \langle l'0l'0|j'0 \rangle j_L(kr)}_{(-i)^j \sqrt{\frac{4\pi}{2j+1}} \Theta_j^m(y_2)} Y_j^m(\hat{n}) \end{aligned} \quad (10)$$

where all the angular dependence is in the $Y_j^m(\hat{n})$, and thus we can directly pick up the new multipoles

$$\Theta_j^m(y_2) = \sum_{l'L} (-i)^{l'+L-j} (2L+1) \langle l'0l'm|j'm \rangle \langle l'0l'0|j'0 \rangle j_L(ky_2 - ky_1) \Theta_{l'}^m(y_1) \quad (11)$$

This equation integrates the free streaming from y_1 to y_2 .

The sum over l produces a radial function which is a linear combination of spherical Bessel functions j_L . We define such new radial functions $j_j^{(l'm)}$ so that the line-of-sight mode functions $G_{l'k}^m$ can be written as

$$G_{l'k}^m(r, \hat{n}) \equiv \sqrt{4\pi} \sum_j (-i)^j \sqrt{2j+1} j_j^{(l'm)}(kr) Y_j^m(\hat{n}) \quad (12)$$

in analogy at $(\vec{x} = -r\hat{n})$

$$e^{i\vec{k}\cdot\vec{x}} = \sqrt{4\pi} \sum_j (-i)^j \sqrt{2j+1} j_j(kr) Y_j^0(\hat{n}). \quad (13)$$

From (9) we have that

$$\boxed{j_j^{(l'm)}(kr) \equiv \sum_{l'} (-i)^{l'+L-j} \left(\frac{2L+1}{2j+1} \right) \langle l'0l'm|j'm \rangle \langle l'0l'0|j'0 \rangle j_L(kr)} \quad (14)$$

and (11) becomes

$$\boxed{\Theta_j^m(y_2) = (2j+1) \sum_{l'} j_j^{(l'm)}(ky_2 - ky_1) \Theta_{l'}^m(y_1)} \quad (15)$$

where we have so far included only the free streaming effect.

The radial functions $j_j^{(l'm)}$ turn out to be real (since $\langle l'0l'0|j'0 \rangle = 0$ for $l'+L-j$ odd).

- We now have to add also the effect of the loss and source terms:
- The loss term is easy to integrate,

$$\Theta_L^m = -a n_e \sigma_T \Theta_L^m \Rightarrow d \ln \Theta_L^m = -a n_e \sigma_T dy$$

$$\Rightarrow \ln \frac{\Theta_L^m(\eta_2)}{\Theta_L^m(\eta_1)} = - \int_{\eta_1}^{\eta_2} a n_e \sigma_T dy \Rightarrow \Theta_L^m(\eta_2) = e^{-\tau(\eta_1, \eta_2)} \Theta_L^m(\eta_1) \quad (16)$$

where the background quantity is the optical depth from η_1 to η_2 .

$$\tau(\eta_1, \eta_2) \equiv \int_{\eta_1}^{\eta_2} a n_e \sigma_T dy \quad (17)$$

Eq. (16) represents the damping of the anisotropy due to the Thomson scattering loss.

- Combining it with the free streaming effect at (15), we have

$$\Theta_j^m(\eta_2) = (2j+1) e^{-\tau(\eta_1, \eta_2)} \sum_{l'} \Theta_{l'}^m(\eta_1) j_l^{(l'm)}(k\eta_2 - k\eta_1) \quad (18)$$

which represents how the anisotropy present at an earlier time η_1 is damped and transported by free streaming to a later time η_2 .

- However, between η_1 and η_2 , new anisotropy is also generated by the source term,

$$\frac{d\Theta_L^m(\eta)}{d\eta} = S_L^m(\eta) \quad (19)$$

- After being generated at time $\eta \in (\eta_1, \eta_2)$, this new anisotropy is also damped and transported by free streaming, but only from time η to η_2 . Thus this adds a new term to Eq. (18),

$$\Theta_j^m(\eta_2) = (2j+1) e^{-\tau(\eta_1, \eta_2)} \sum_{l'} \Theta_{l'}^m(\eta_1) j_l^{(l'm)}(k\eta_2 - k\eta_1) + (2j+1) \int_{\eta_1}^{\eta_2} d\eta e^{-\tau(\eta, \eta_2)} \sum_{l'} S_{l'}^m(\eta) j_l^{(l'm)}(k\eta_2 - k\eta) \quad (20)$$

This is the full result, that evolves the multipoles Θ_L^m from η_1 to η_2 .

Note that the sum for the source terms $S_{l'}^m$ goes only to $l'=2$, since there are no higher source terms.

- Towards early times the rate of scattering $\sigma_T n_e$ increases, and as $\eta_1 \rightarrow 0$ (or in practice, to some time sufficiently long before photon decoupling),

$$\tau(\eta_1, \eta_2) \rightarrow \infty \quad \Rightarrow \quad e^{-\tau(\eta_1, \eta_2)} \rightarrow 0 \quad (21)$$

so that the original anisotropy at time η_1 is completely damped away, and all the anisotropy at η_2 comes from the source terms.

- Thus we can write for the present-day brightness function

$$\Theta_j^m(\eta_0) = (2j+1) \int_0^{\eta_0} d\eta e^{-\tau(\eta)} \sum_{L'} S_{L'}^m(\eta) j_j^{(L'm)}(k\eta_0 - k\eta) \quad (22)$$

where $\tau(\eta) \equiv \tau(\eta, \eta_0)$ is the optical depth from time η to present time. (23)

- Eq. (22) is the line-of-sight integral for $\Theta_L^m(\eta_0)$. It is a formal solution of Eq. (1.1a) which can be verified by direct substitution. (Doing this is some work; it involves the properties of spherical Bessel functions. $j_L(x)$, like recursion relations and values at $x=0$.) Note that the sources $S_{L'}^m(\eta)$ include $\Theta_0^m(\eta)$, $\Theta_2^m(\eta)$, and $E_2^m(\eta)$. Thus Eq. (22) will not by itself solve Eq. (1.1a). We still need to integrate Eq. (1) numerically for the low L to get the sources $S_{L'}^m(\eta)$. The point of Eq. (22) is that it gives all the higher L from these low multipoles.

- Now in (22) & (14) the L' represents the source, l the plane-wave, and j the total angular dependence; and the radial functions $j_L^{(l'm)}$ are needed only for $L=0,1,2$ and $m = -L, \dots, L$, but they are needed for all l up to $l = j_{\max} + 2$, where j_{\max} is the highest multipole we are interested in.

- For (22), we only need those radial functions $j_j^{(l'm)}$ for which $S_{L'}^m$ are nonzero.

L2.2 Polarization

Polarization is handled the same way. First the free streaming, which gives

$$(Q \pm iU)(\gamma_2, \vec{x}_2, \hat{n}) = (Q \pm iU)(\gamma_1, \vec{x}_2 - r\hat{n}, \hat{n}) \quad (24)$$

The multiple expansion of (2) is

$$\begin{aligned} (Q \pm iU)(\gamma_2, \vec{x}_2, \hat{n}) &= \sum_{lm} (-i)^l \sqrt{\frac{4\pi}{2l+1}} [E_l^m(\gamma) \pm iB_l^m(\gamma)] \pm_2 Y_l^m(\hat{n}) e^{ik \cdot \vec{x}} \\ &\equiv \sum_{lm} [E_l^m(\gamma) \pm iB_l^m(\gamma)] \pm_2 G_{lk}^m(\vec{x}, \hat{n}) \end{aligned} \quad (25)$$

where we have defined the polarization mode functions $\pm_2 G_{lk}^m$ as

$$\pm_2 G_{lk}^m(\vec{x}, \hat{n}) \equiv (-i)^l \sqrt{\frac{4\pi}{2l+1}} \pm_2 Y_l^m(\hat{n}) e^{ik \cdot \vec{x}} \quad (26)$$

Written in terms of the mode functions, Eq. (24) becomes

$$(Q \pm iU)(\gamma_2, \vec{x}_2, \hat{n}) = \sum_{lm} [E_l^m(\gamma_1) \pm iB_l^m(\gamma_1)] \pm_2 G_{lk}^m(\vec{x}_2 - r\hat{n}, \hat{n}) \quad (27)$$

$$\text{where } \pm_2 G_{lk}^m(\vec{x}_2 - r\hat{n}, \hat{n}) = (-i)^l \sqrt{\frac{4\pi}{2l+1}} \pm_2 Y_l^m(\hat{n}) e^{ik \cdot \vec{x}_2} e^{-ikr\hat{z} \cdot \hat{n}} \equiv e^{ik \cdot \vec{x}_2} \pm_2 G_{lk}^m(r, \hat{n}) \quad (28)$$

$$\text{and } \pm_2 G_{lk}^m(r, \hat{n}) \equiv \pm_2 G_{lk}^m(-r\hat{z}, \hat{n}) = (-i)^l \sqrt{\frac{4\pi}{2l+1}} \pm_2 Y_l^m(\hat{n}) e^{-ikr\hat{z} \cdot \hat{n}} \quad (29)$$

is the two-of-right polarization mode function.

Now we just need to pick the multipoles $E_l^m(\gamma_2)$, $B_l^m(\gamma_2)$ of $(Q \pm iU)(\gamma_2, \hat{n})$ from

$$(Q \pm iU)(\gamma_2, \vec{x}_2, \hat{n}) = e^{ik \cdot \vec{x}_2} \underbrace{\sum_{lm} [E_l^m(\gamma_2) \pm iB_l^m(\gamma_2)] (-i)^l \sqrt{\frac{4\pi}{2l+1}} \pm_2 Y_l^m(\hat{n}) e^{-ikr\hat{z} \cdot \hat{n}}}_{\pm_2 G_{lk}^m(r, \hat{n})} \quad (30)$$

Using the expansion (8) of the plane wave, this becomes

$$(Q \pm iU)(\gamma_2, \hat{n}) = \sum_{lm} [E_l^m(\gamma_1) \pm iB_l^m(\gamma_1)] \underbrace{(-i)^l \sqrt{\frac{4\pi}{2l+1}} \cdot \sqrt{4\pi} \sum_l (-i)^l \sqrt{2l+1} j_l(kr) Y_l^0(\hat{n})}_{\pm_2 G_{lk}^m(r, \hat{n})} \cdot \pm_2 Y_l^m(\hat{n}) \quad (31)$$

To pick out the multipoles of Eq. (31), we need to express the total angular dependence of $\pm 2 G_{l'k}^m(r, \hat{n})$ as a linear combination of $s Y_j^m(\hat{n})$. We use the CG series for spin-s harmonics (Eq. C5.5), which gives

$$Y_l^0(\hat{n}) s Y_{l'}^m(\hat{n}) = \sqrt{\frac{(2L+1)(2L'+1)}{4\pi}} \sum_j \frac{1}{\sqrt{2j+1}} \langle l 0 l' m | j m \rangle \langle l 0 j l' -s | j, -s \rangle s Y_j^m(\hat{n}) \quad (32)$$

to write

$$s G_{l'k}^m(r, \hat{n}) = \sum_{l'j} (-i)^{l'+l} (2L+1) \sqrt{\frac{4\pi}{2j+1}} \langle l 0 l' m | j m \rangle \langle l 0 j l' -s | j, -s \rangle j_l(kr) s Y_j^m(\hat{n}). \quad (33)$$

With (33), Eq. (31) becomes

$$(Q \pm iU)(y_2, \hat{n}) = \sum_{jm} \sum_{l'l'} \underbrace{[E_{l'}^m(y_1) \pm i B_{l'}^m(y_1)] (-i)^{l+l'} (2L+1) \sqrt{\frac{4\pi}{2j+1}} \langle l 0 l' m | j m \rangle \langle l 0 j l' \mp 2 | j, \mp 2 \rangle j_l(kr)}_{(-i)^j \sqrt{\frac{4\pi}{2j+1}} [E_j^m(y_2) \pm i B_j^m(y_2)]} \pm 2 Y_j^m(\hat{n}) \quad (34)$$

from which we can directly pick up the polarization multipoles

$$E_j^m(y_2) \pm i B_j^m(y_2) = \sum_{l'l'} (-i)^{l+l'-j} (2L+1) \langle l 0 l' m | j m \rangle \langle l 0 j l' \mp 2 | j, \mp 2 \rangle j_l(kr) [E_{l'}^m(y_1) \pm i B_{l'}^m(y_1)] \quad (35)$$

This equation integrates the free streaming of polarization from y_1 to y_2 .

(To get the E_j^m and B_j^m just take the sum and the difference of the \pm cases).

Again, we define new radial functions $\pm 2 j_j^{(l'm)}$ to correspond to the linear combinations of j_l (sum over l) that appear in (35) for each l^2 term:

$$\pm 2 j_j^{(l'm)}(kr) \equiv \sum_l (-i)^{l+l'-j} \left(\frac{2L+1}{2j+1} \right) \langle l 0 l' m | j m \rangle \langle l 0 j l' \mp 2 | j, \mp 2 \rangle j_l(kr) \quad (36)$$

so that

$$\pm 2 G_{l'k}^m(r, \hat{n}) = \sqrt{4\pi} \sum_j (-i)^j \sqrt{2j+1} \pm 2 j_j^{(l'm)}(kr) \pm 2 Y_j^m(\hat{n}) \quad (37)$$

and

$$E_j^m(y_2) \pm i B_j^m(y_2) = (2j+1) \sum_{l'} \pm 2 j_j^{(l'm)}(ky_2 - ky_1) [E_{l'}^m(y_1) \pm i B_{l'}^m(y_1)] \quad (38)$$

Including then the effect of the loss term, $\frac{d}{dy} (E_L^m \pm iB_L^m) = -\alpha n_e \sigma_T \cdot (E_L^m \pm iB_L^m)$ (39)

and the source term, $\frac{d}{dy} (E_L^m \pm iB_L^m) = Z_L^m$, we get the line-of-sight integral (40)

from y_1 to y_2 (in analogy to Eq.(20)) as

$$\boxed{E_j^m(y_2) \pm iB_j^m(y_2) = (2j+1) e^{-\tau(y_1, y_2)} \sum_{L'} [E_{L'}^m(y_1) \pm iB_{L'}^m(y_1)] \pm 2j_j^{(L'm)}(ky_2 - ky_1) + (2j+1) \int_{y_1}^{y_2} dy e^{-\tau(y_1, y_2)} Z_2^m(y) \pm 2j_j^{(2m)}(ky_2 - ky_1)} \quad (41)$$

Note that there is no sum over L' for the source terms Z_2^m , since the only nonzeros have $L'=2$.

For the present-day polarization multipoles we thus get

$$E_j^m(y_0) \pm iB_j^m(y_0) = (2j+1) \int_0^{y_0} dy e^{-\tau(y)} Z_2^m(y) \pm 2j_j^{(2m)}(ky_0 - ky) \quad (42)$$

It turns out (§L3) that the $\pm 2j_j^{(2m)}$ are complex conjugates of each other

$$-2j_j^{(2m)} = +2j_j^{(2m)*} \quad (43)$$

so that they can be written as $\pm 2j_j^{(2m)} \equiv E_j^m \pm iB_j^m$ (44)

where the radial functions E_j^m and B_j^m are real. Thus we obtain from (42)

$$\boxed{E_L^m(y_0) = (2j+1) \int_0^{y_0} dy e^{-\tau(y)} [-\sqrt{6} \alpha n_e \sigma_T P^{(m)}(y)] E_L^m(ky_0 - ky) \quad (45a)}$$

$$B_L^m(y_0) = (2j+1) \int_0^{y_0} dy e^{-\tau(y)} [-\sqrt{6} \alpha n_e \sigma_T P^{(m)}(y)] B_L^m(ky_0 - ky) \quad (45b)$$

Note that the only difference between (45a) and (45b) is the radial functions.

E and B modes have the same source, and the relation between E and B mode polarization is determined completely by the structure of the E_L^m and B_L^m radial functions (and its relation to the time dependence of $P^{(m)}(y)$). They are needed for $m=0, \pm 1, \pm 2$, corresponding to scalar, vector, and tensor perturbations.